

# Controlling light and matter using cooperative radiation

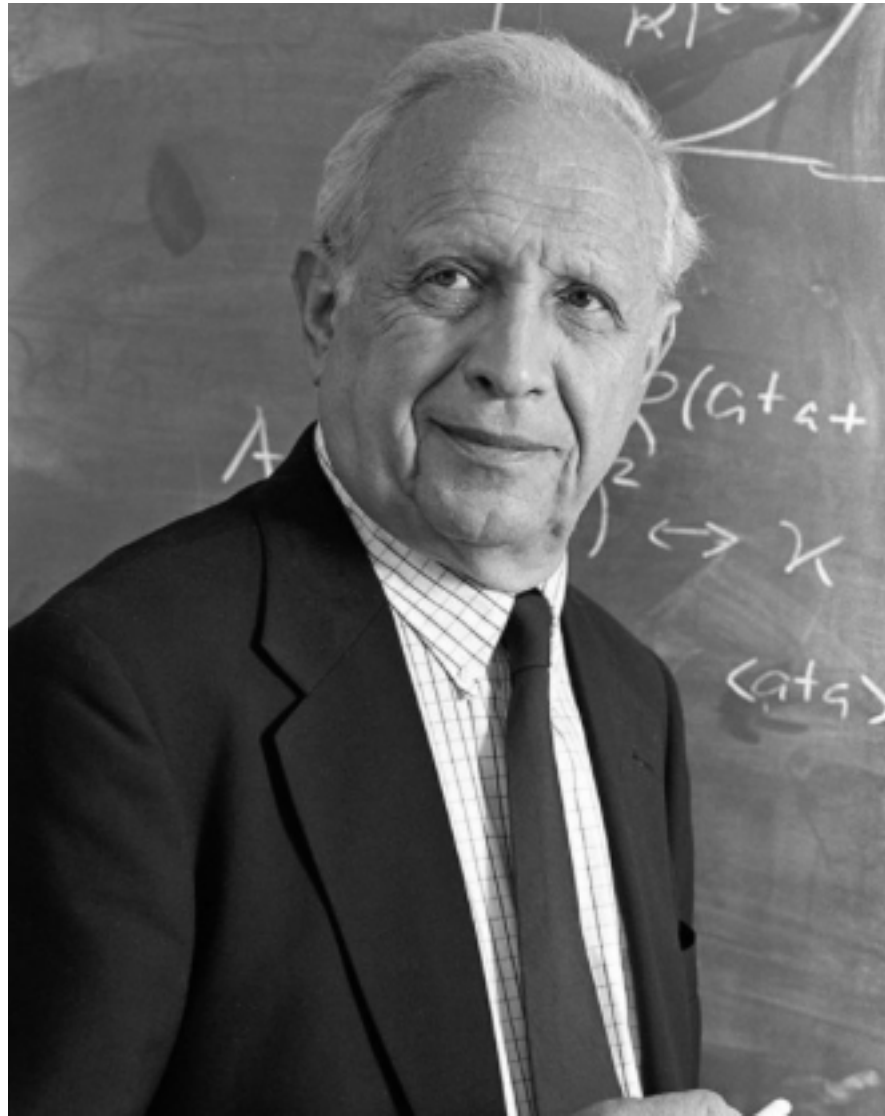
## Part I: Dicke states, cooperative effects, entanglement

Susanne Yelin

University of Connecticut

Harvard University

Herrsching, March 6, 2019

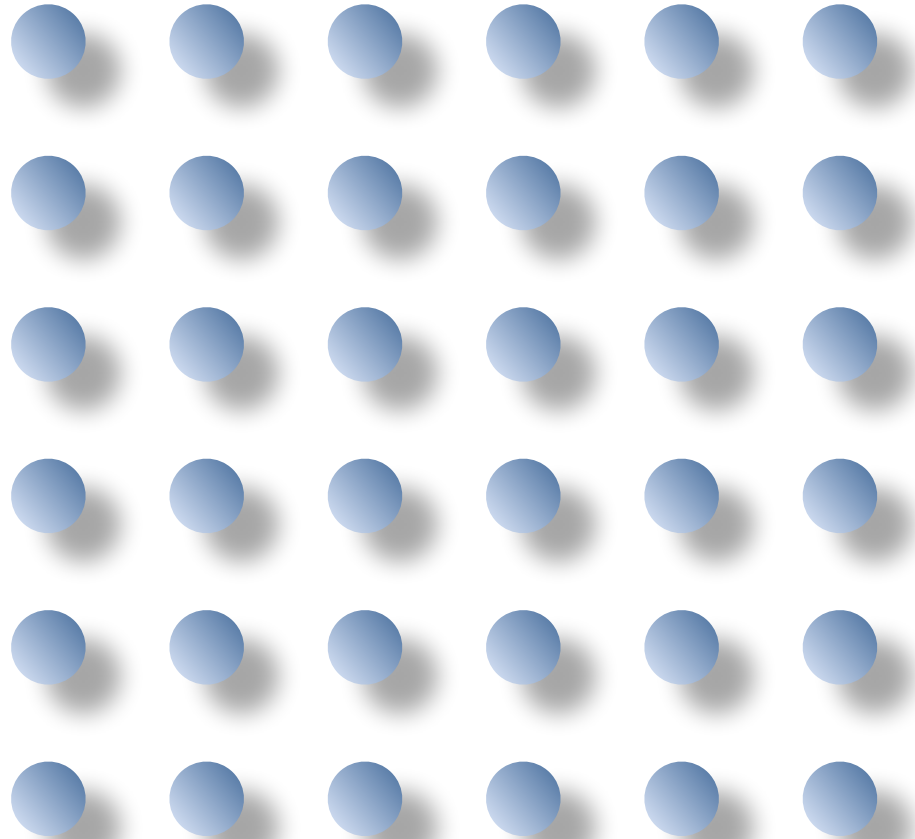


Roy Glauber, 1925 - 2018

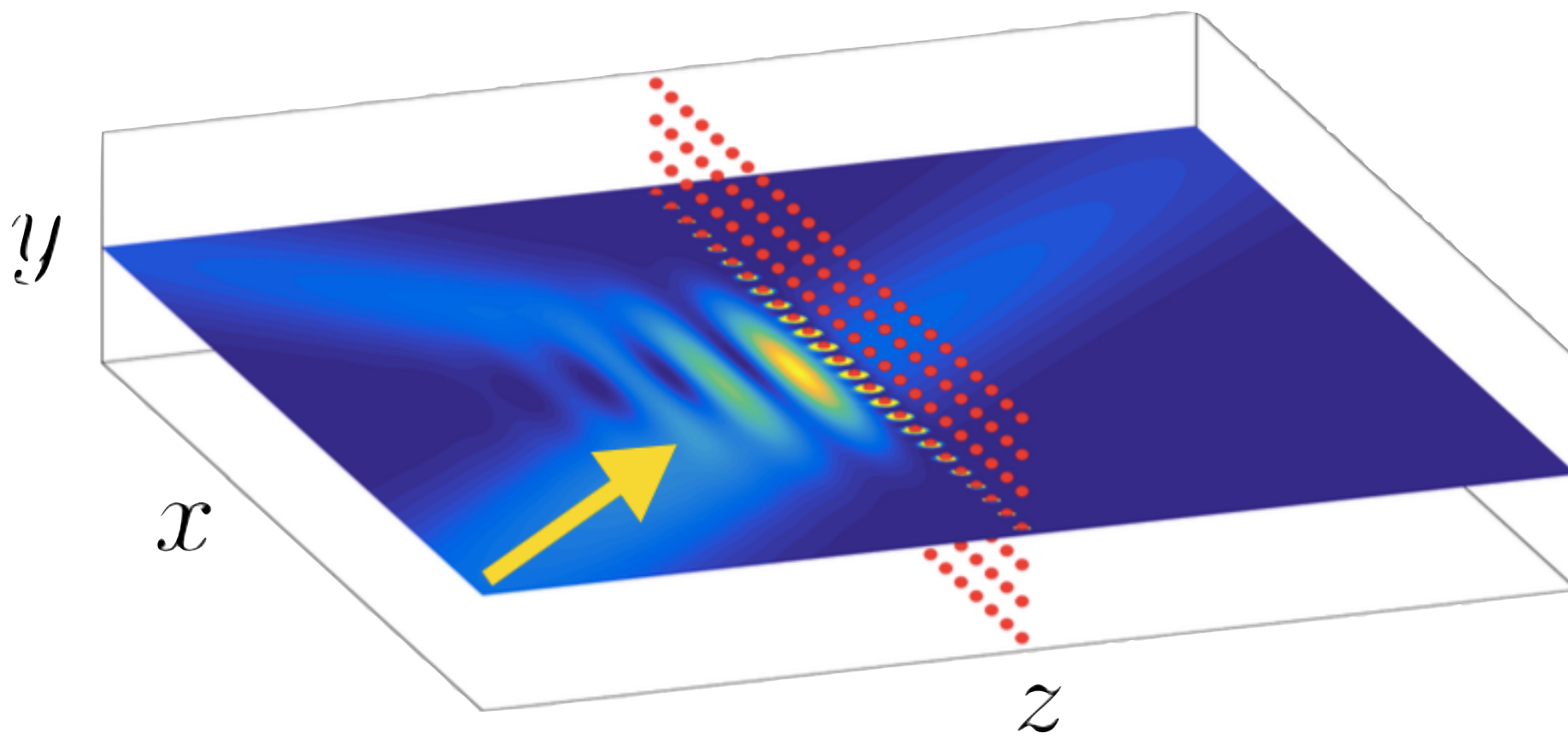
# Goals of the lectures

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- Use 2D array...

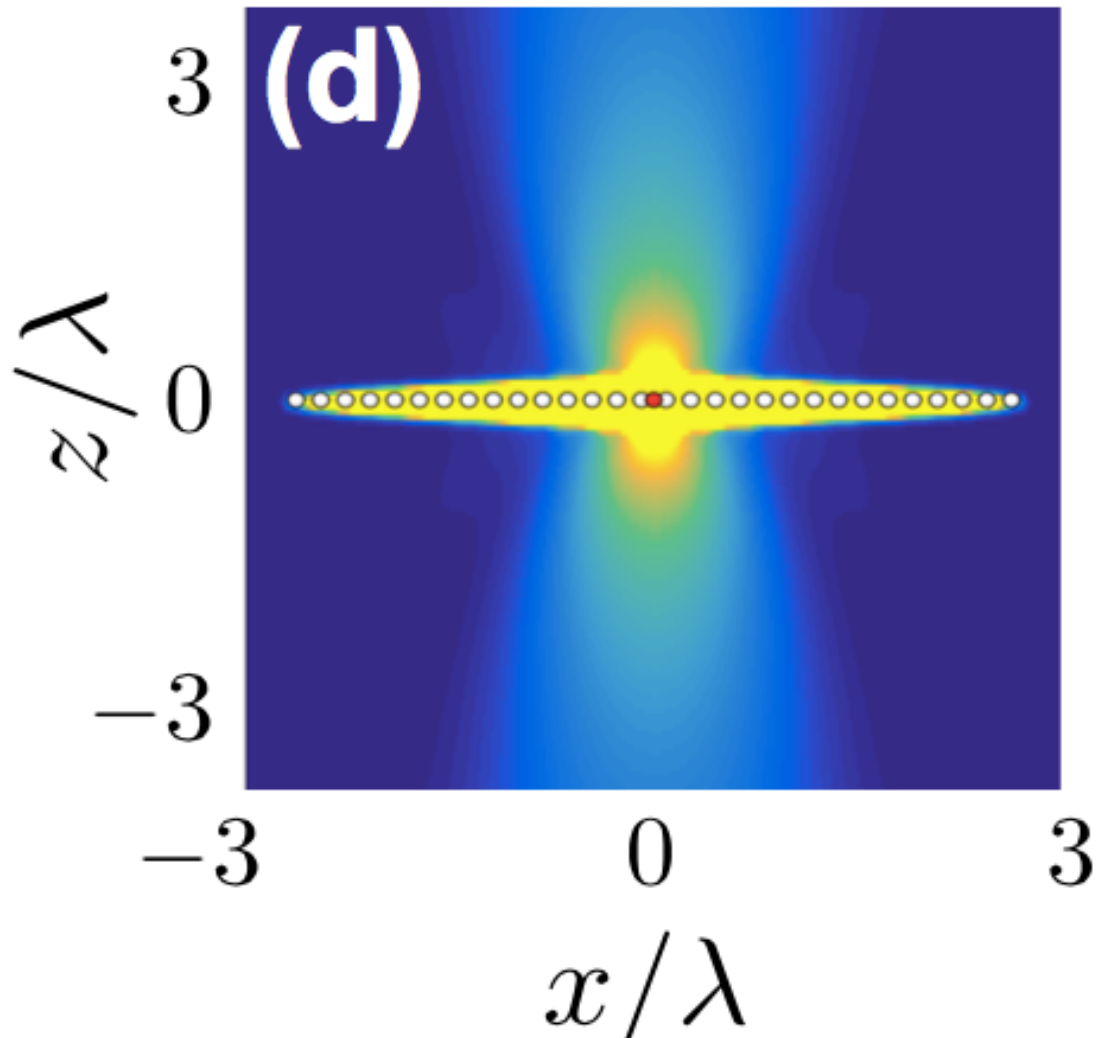


- ... to reflect light ...





- ... emit in controlled way ...



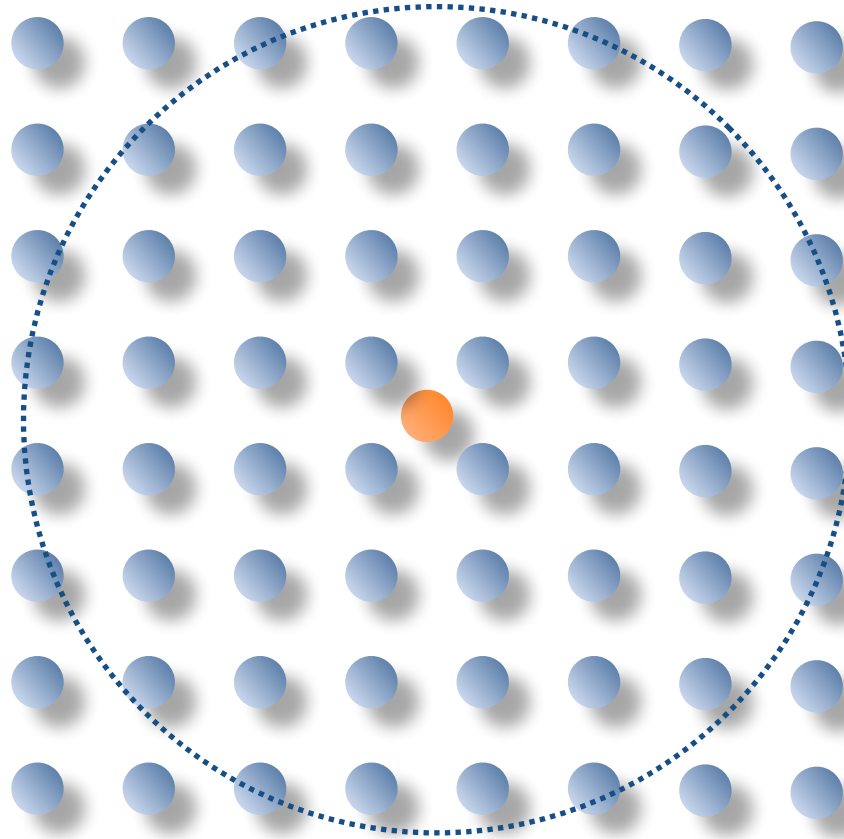
# Increase (impurity) cross section?

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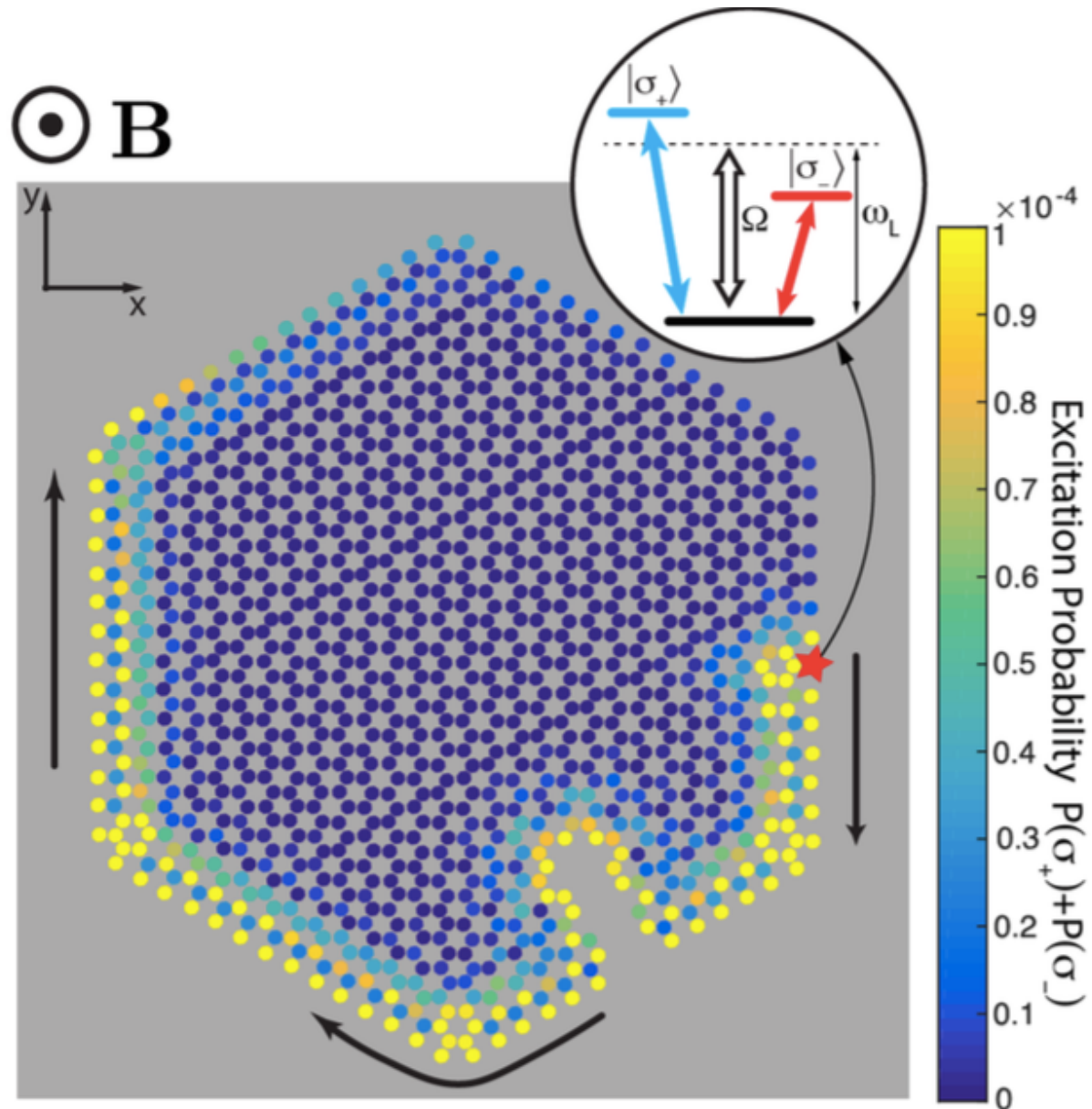


# Increase (impurity) cross section?

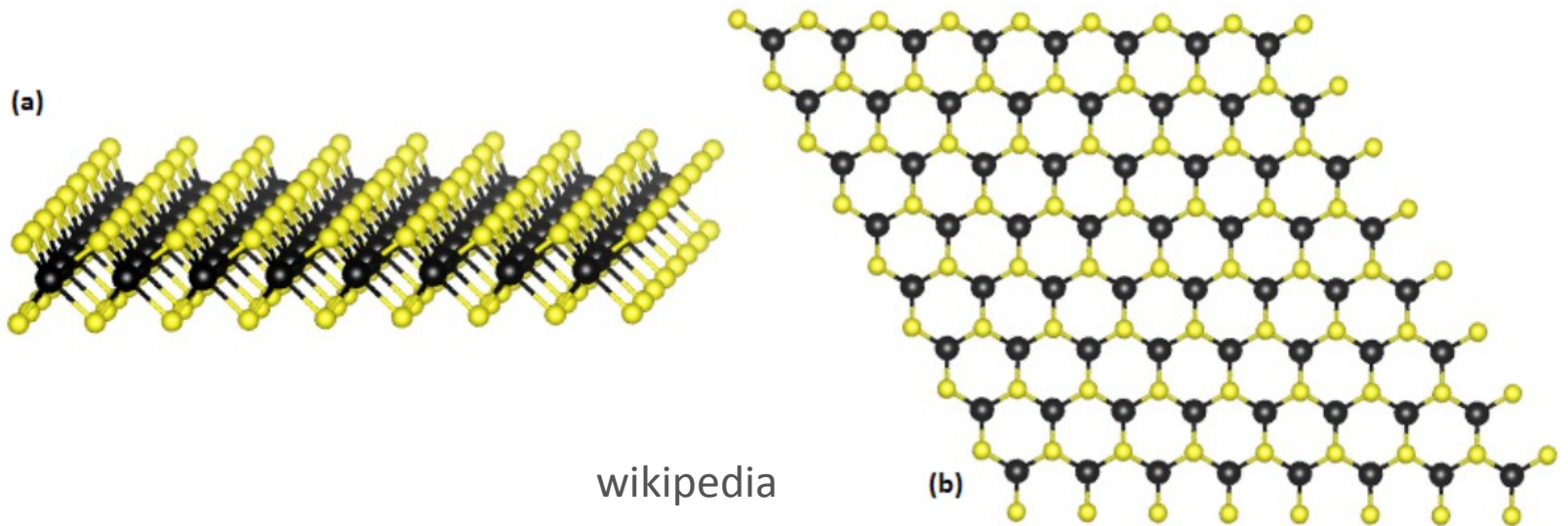
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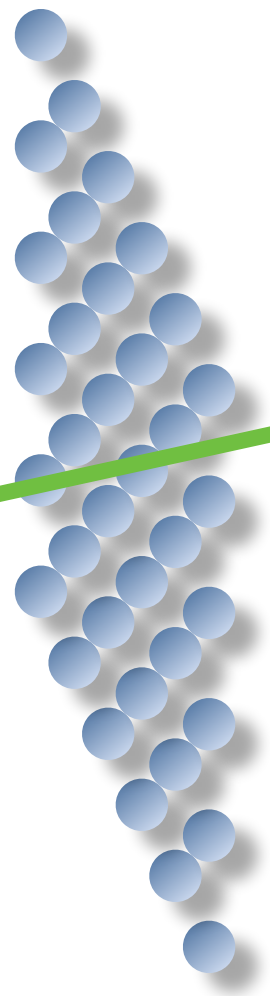
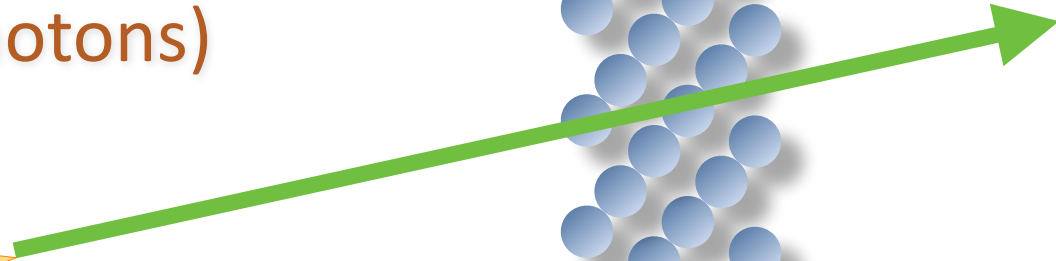
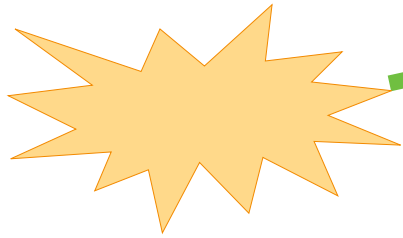
- ... edge states with photons?



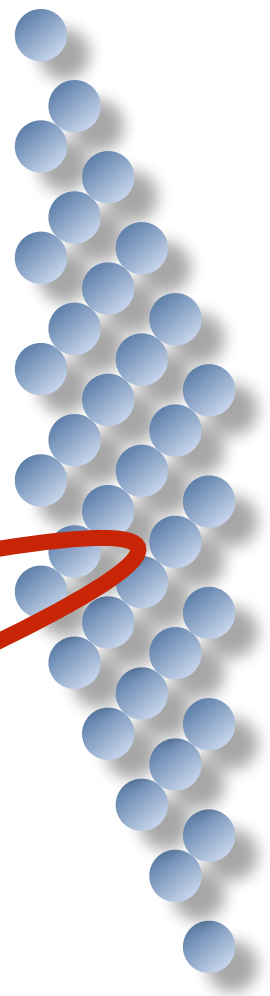
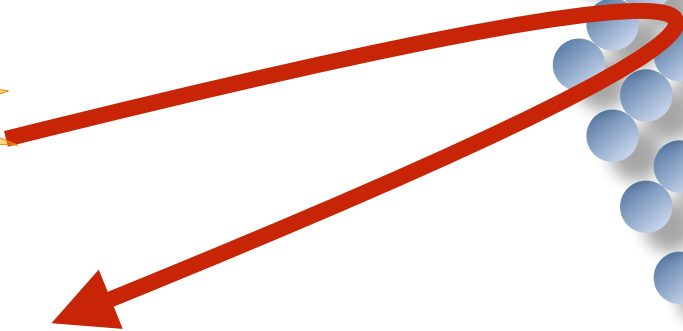
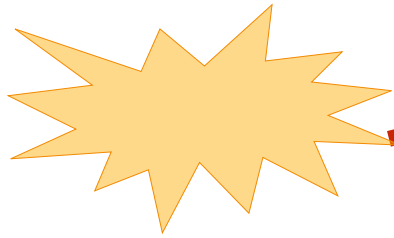
- 2D materials like graphene or TMDs



- ... to switch  
(single photons)



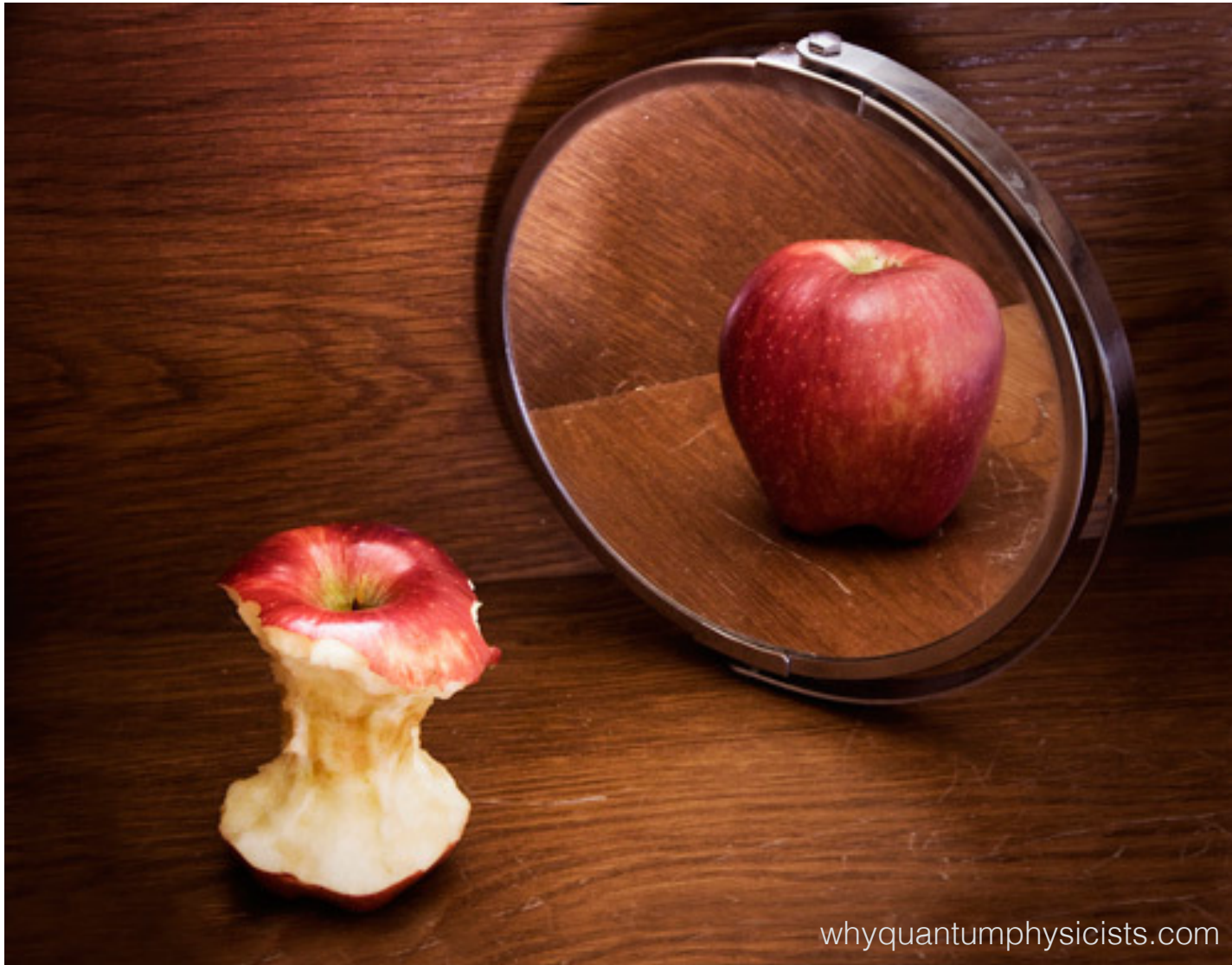
- ... to switch  
(single photons)





# Quantum mirror: Refraction superposition

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**Why “cooperative effects”??**

# Cooperative radiation: superradiance

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For want of a better term, a gas which is radiating strongly because of coherence will be called “superradiant.”

— Robert H Dicke, 1954.

PHYSICAL REVIEW

VOLUME 93, NUMBER 1

JANUARY 1, 1954

## Coherence in Spontaneous Radiation Processes

R. H. DICKE

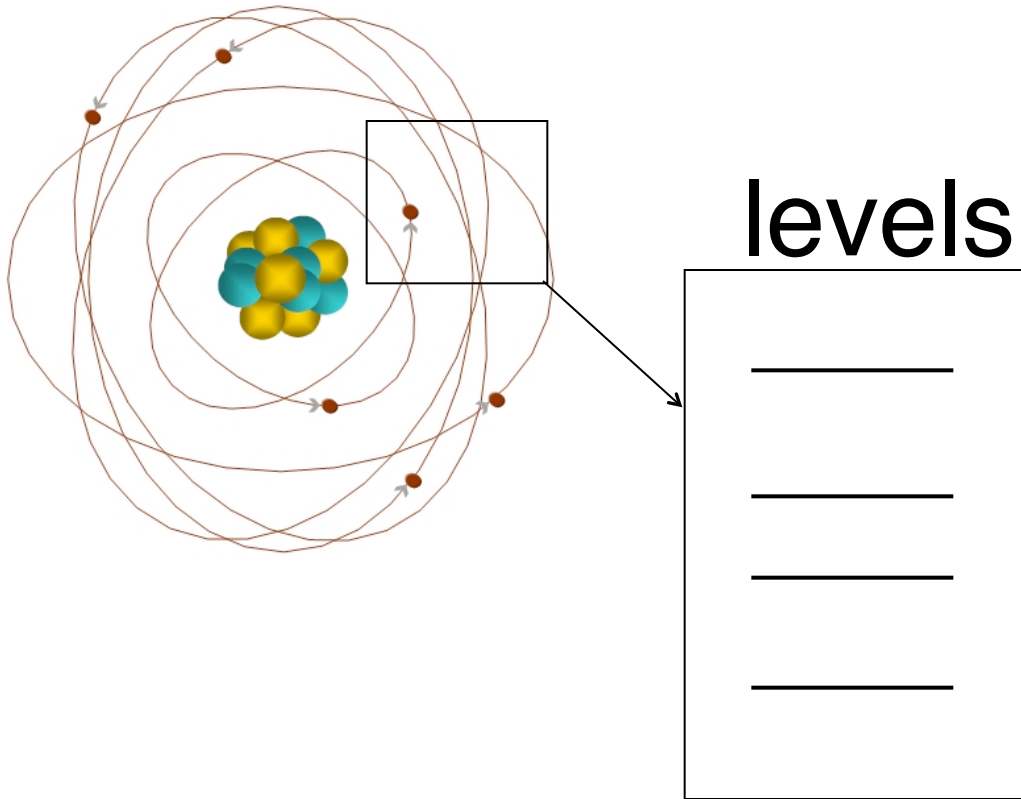
*Palmer Physical Laboratory, Princeton University, Princeton, New Jersey*

(Received August 25, 1953)

By considering a radiating gas as a single quantum-mechanical system, energy levels corresponding to certain correlations between individual molecules are described. Spontaneous emission of radiation in a transition between two such levels leads to the emission of coherent radiation. The discussion is limited first to a gas of dimension small compared with a wavelength. Spontaneous radiation rates and natural line

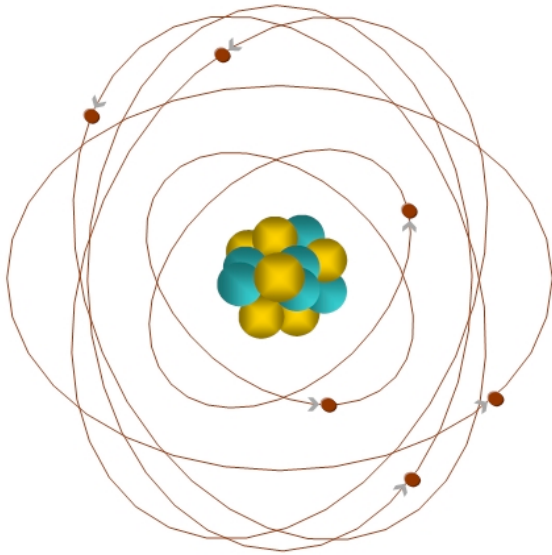
# What is an atom?

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# What is an atom?

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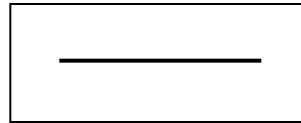
levels

\_\_\_\_\_

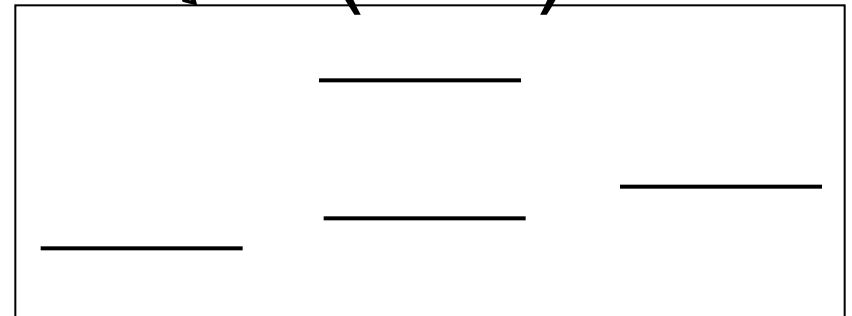
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(sub)levels

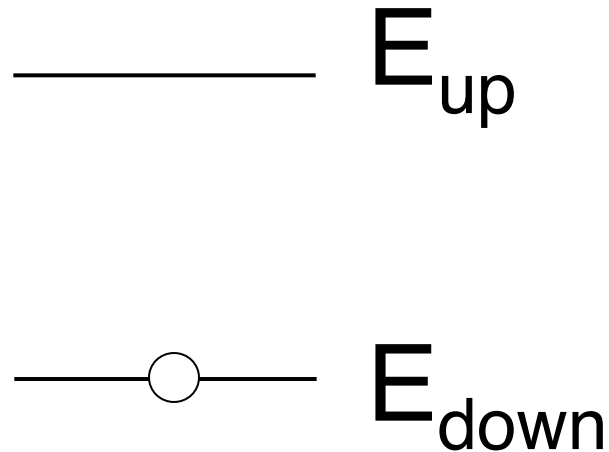
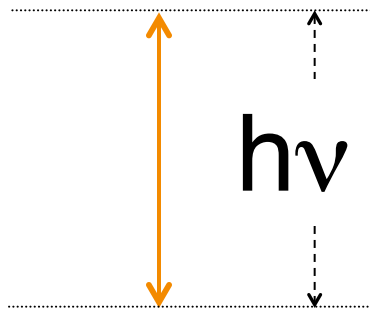


# How does it couple to light?

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photon energy:  $h\nu$

atomic transition energy:  $E_{\text{up}} - E_{\text{down}}$

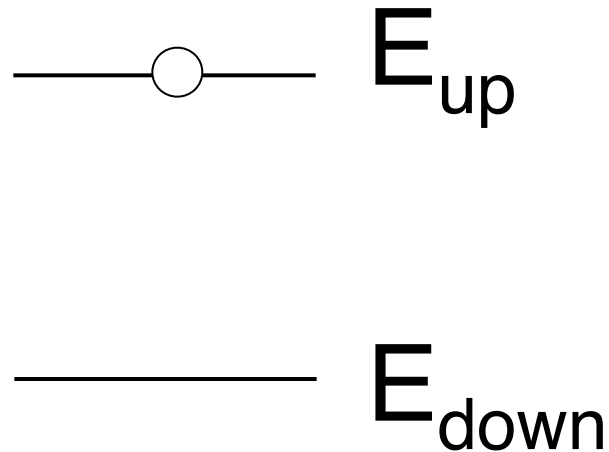
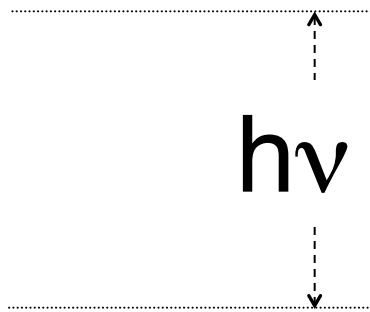


# How does it couple to light?

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photon energy:  $h\nu$

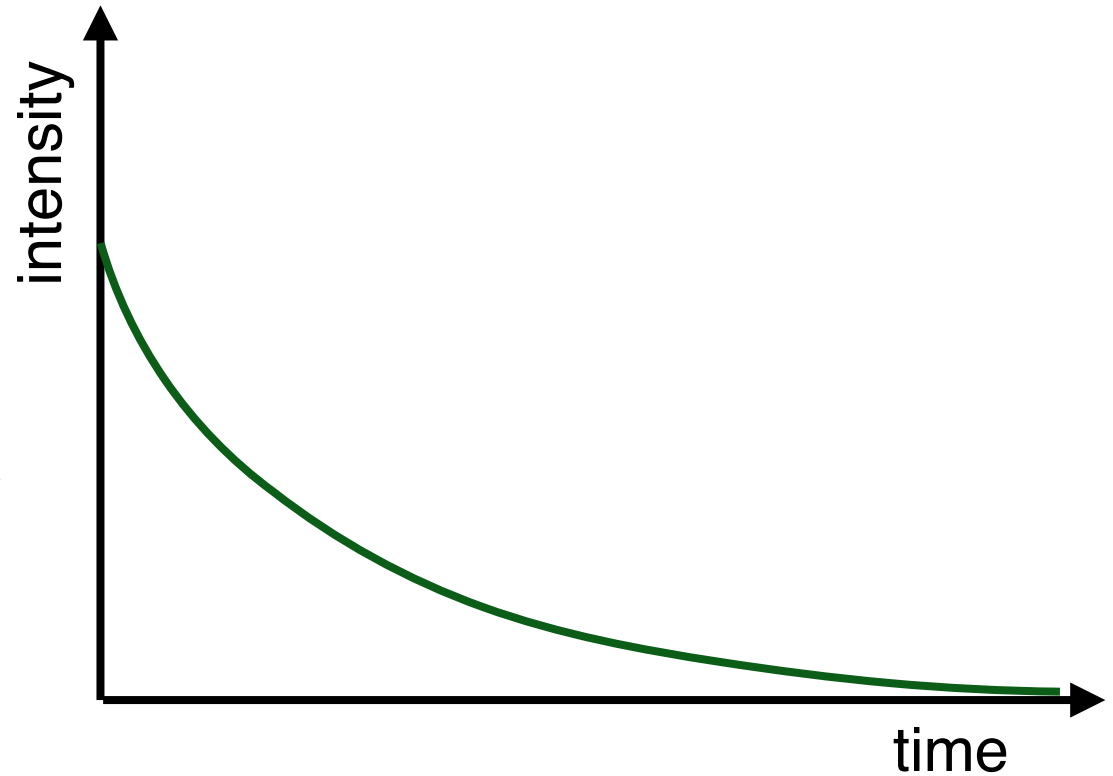
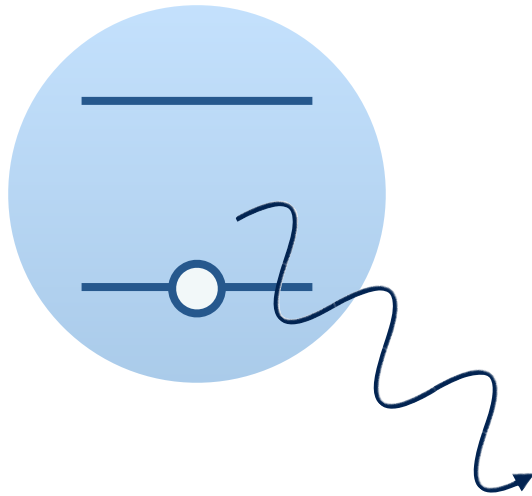
atomic transition energy:  $E_{\text{up}} - E_{\text{down}}$



# Cooperative effects in radiation

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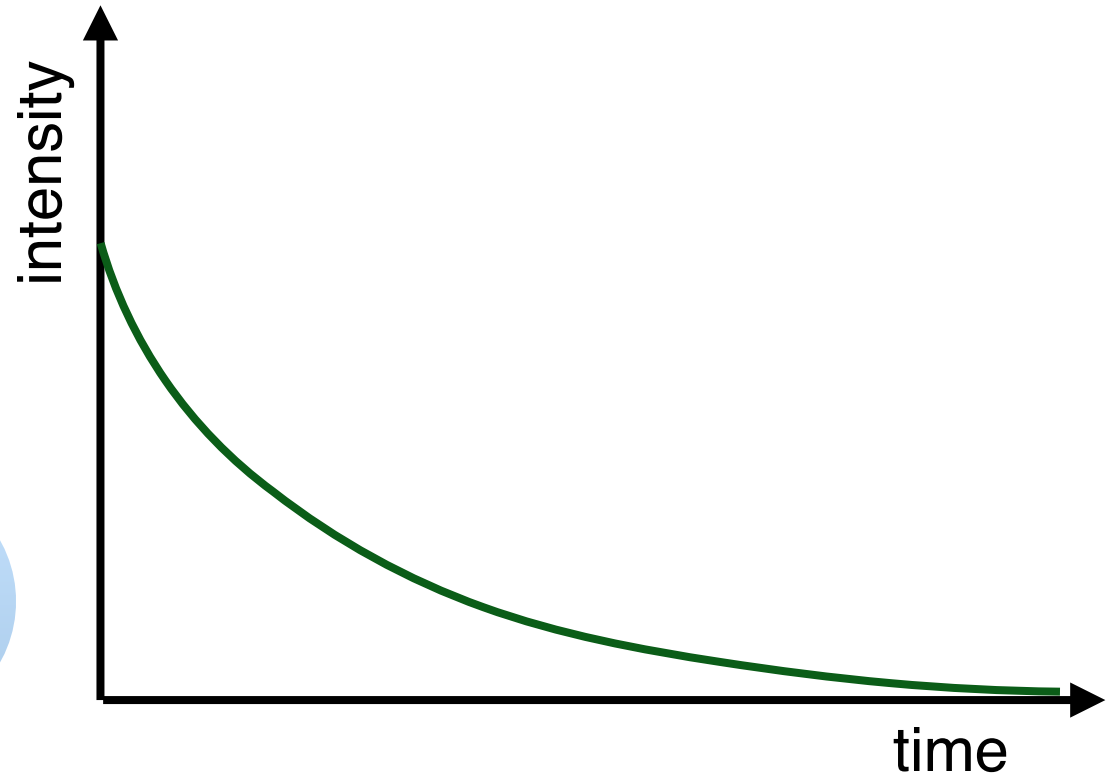
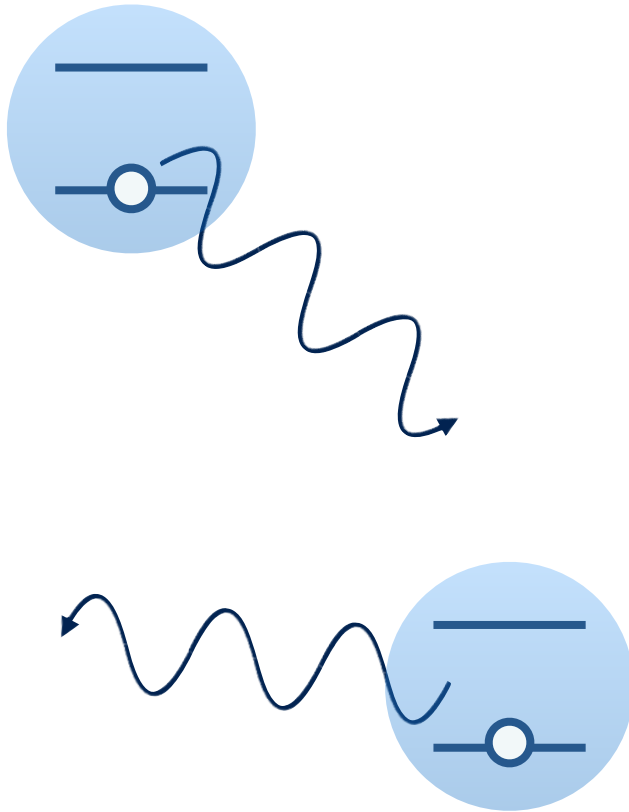
single atom



# Cooperative effects in radiation

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two far-away atoms

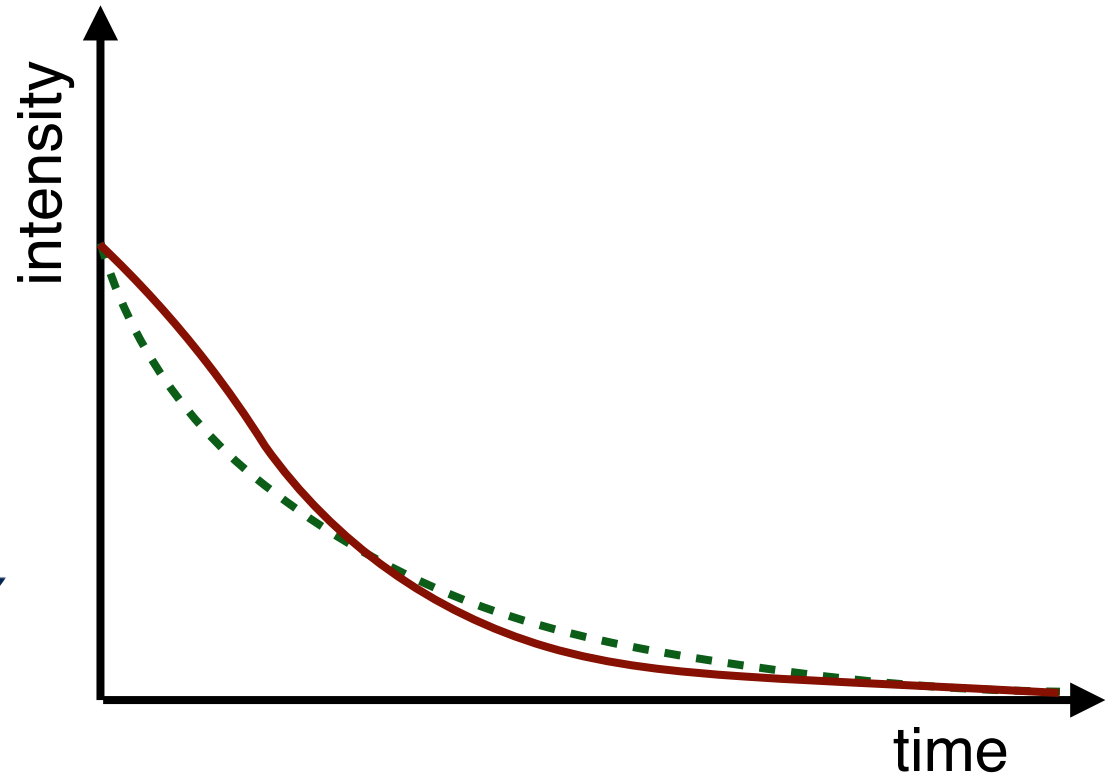
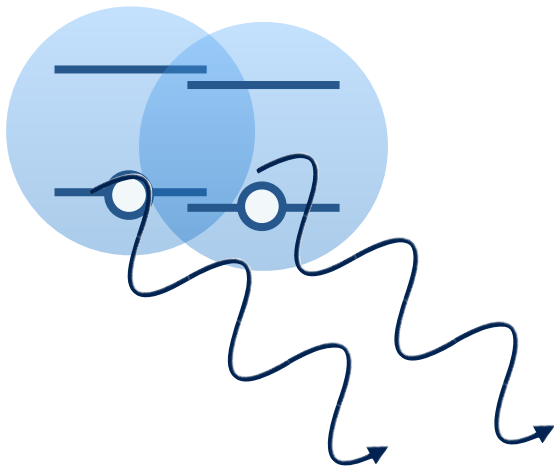




# Cooperative effects in radiation

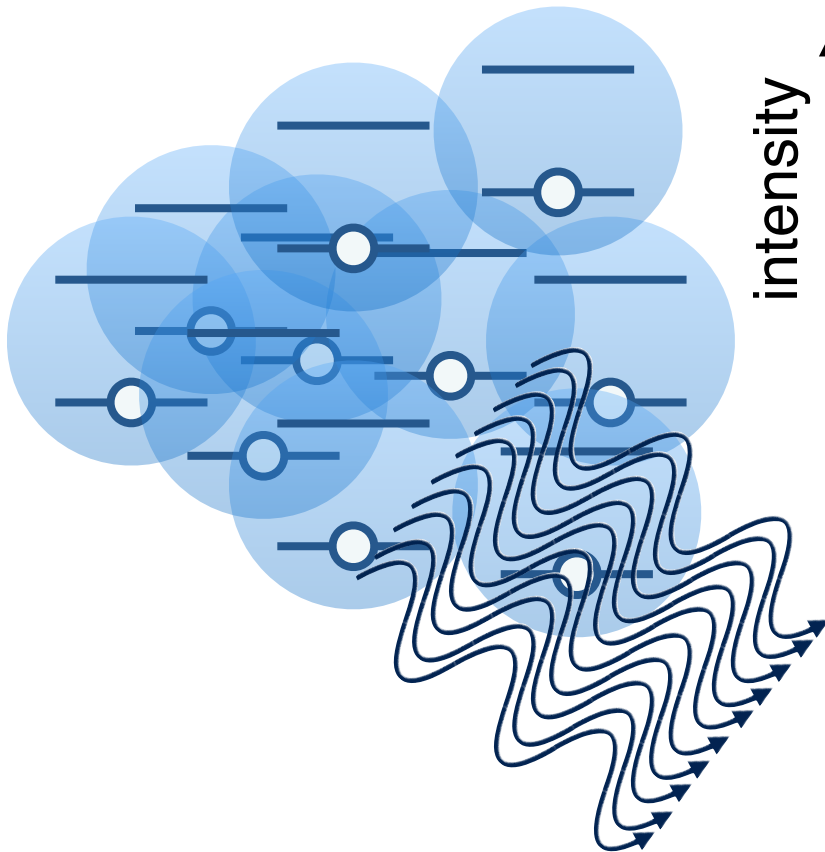
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two close atoms

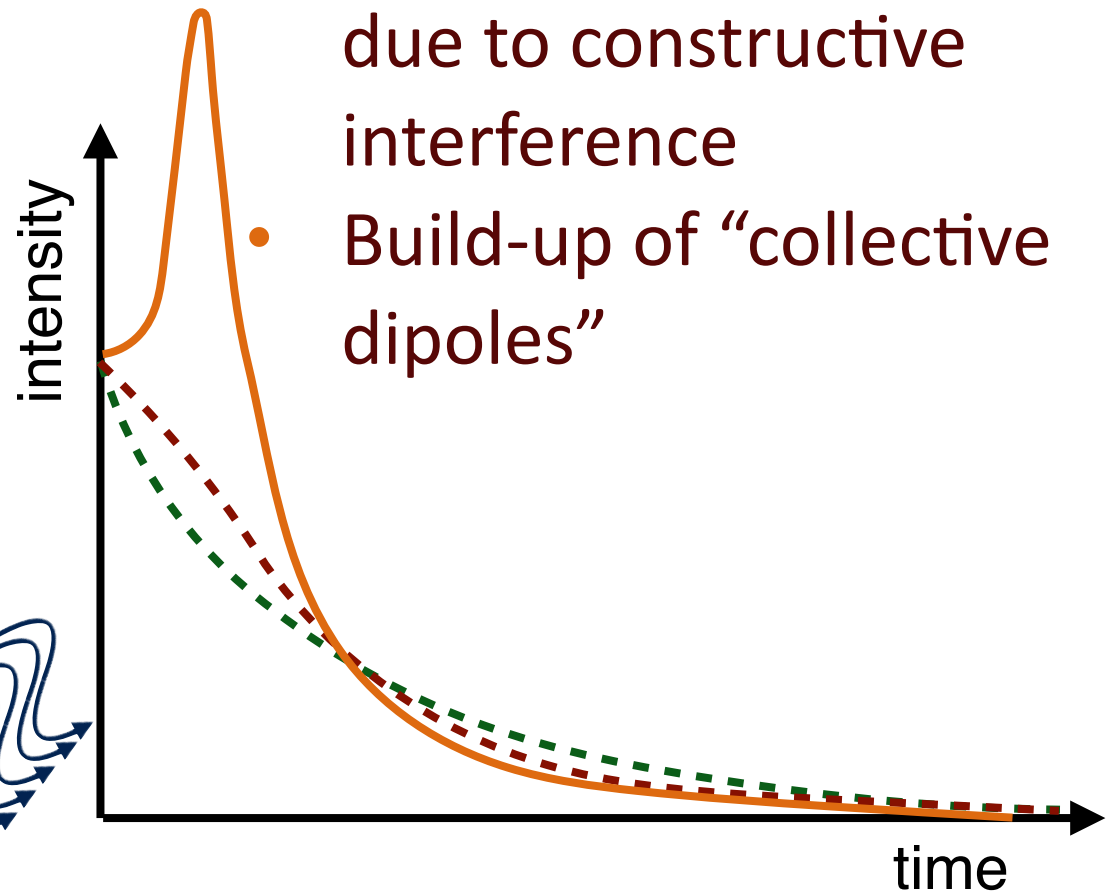


# Cooperative effects in radiation

many close atoms



- Superradiance  $\propto N^2$  due to constructive interference
- Build-up of “collective dipoles”



# Cooperative Effects

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- Two important aspects:
  - collective effects (many particles)

# Cooperative Effects

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- Two important aspects:
  - collective effects (many particles)
    - e.g., much higher chance for photons to interact with many atoms than one
  - “exchange” due to dipole-dipole interaction
    - excitations are exchanged
- “Cooperative” is more than just “collective”!

# Collective effects: Example of quantized field

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- Interaction Hamiltonian with classical light (Rabi frequency  $\Omega$ )

$$H/\hbar = \Omega (|e\rangle\langle g| + |g\rangle\langle e|)$$

- Interaction Hamiltonian with quantized light (coupling element  $g$ , annihilation operator  $a$ , number of atoms  $N$ )

$$H/\hbar = g\sqrt{N} (|e\rangle\langle g| a + a^\dagger |g\rangle\langle e|)$$

# Cooperative effects

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- simplest form of “exchange interaction”

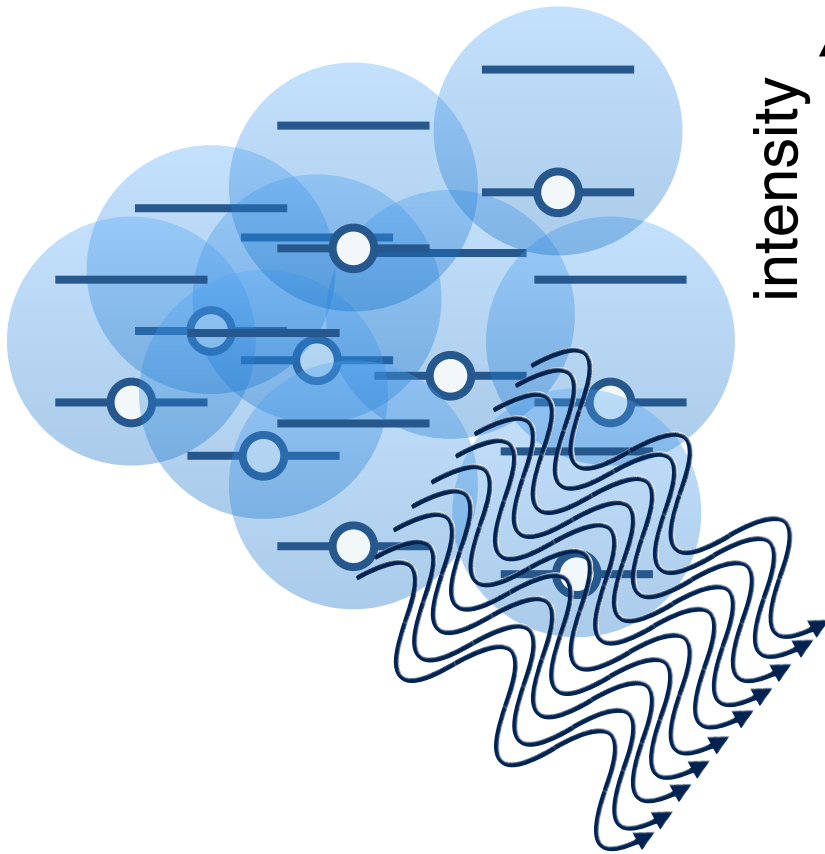
# Cooperative effects

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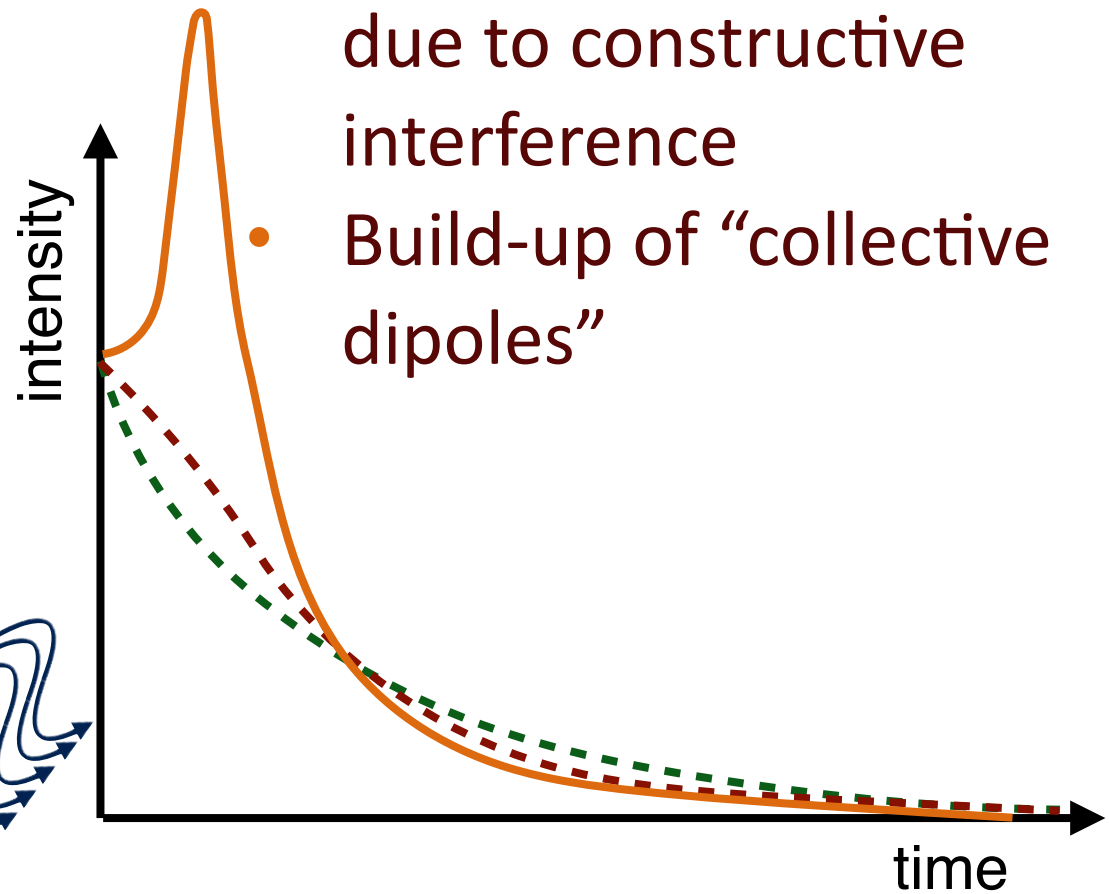
- Traditional example: superradiance

# Cooperative effects in radiation

many close atoms



- Superradiance  $\propto N^2$  due to constructive interference
- Build-up of “collective dipoles”





# What is **super** in superradiance?

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man (Clark)



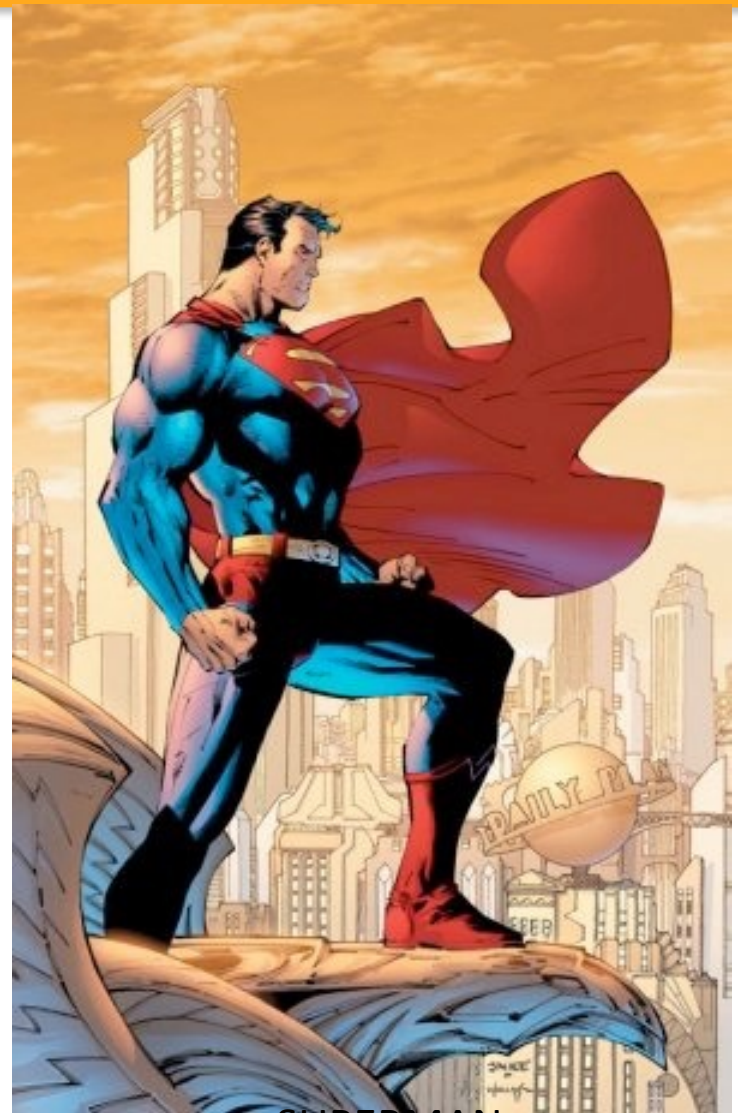
many men (Clarks)

# What is **super** in superradiance?

man (Clark)



many men (Clarks)



SUPERMAN

# What is **super** in superradiance?

man (Clark)



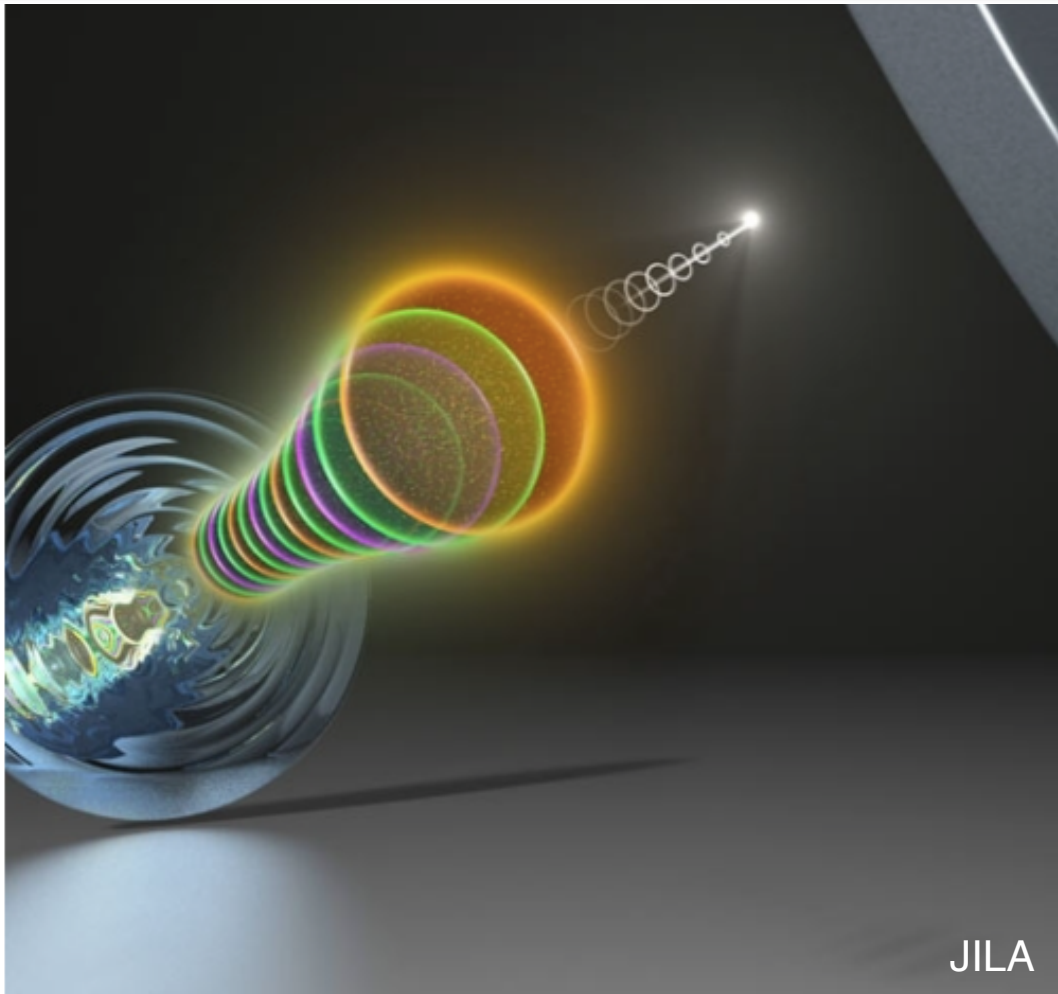
many men (Clarks)



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# Examples

# Superradiant laser



$\# \text{Atoms} \gg \# \text{Photons}$



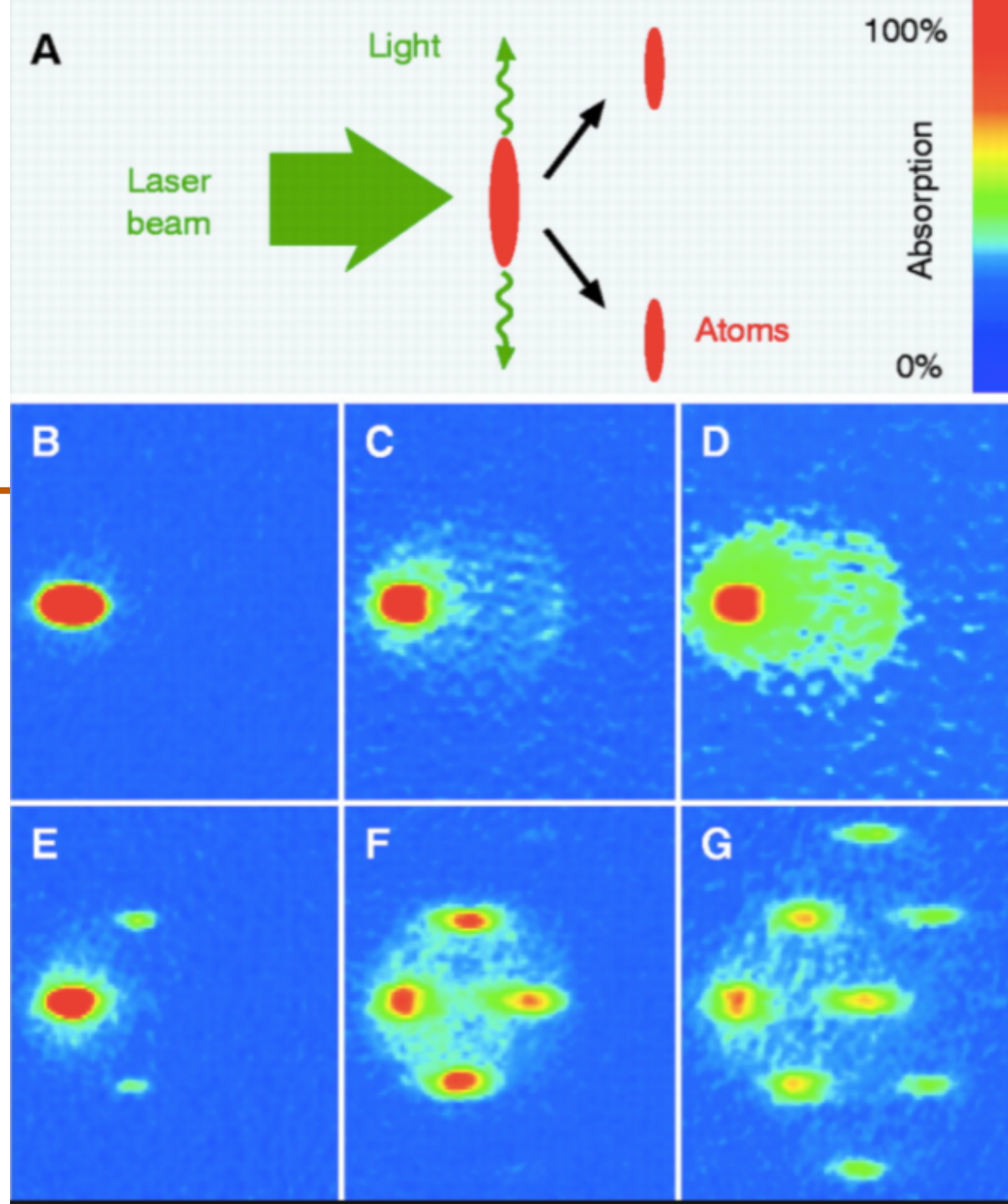
much more coherent  
lasing



# Super-radiance in BECs

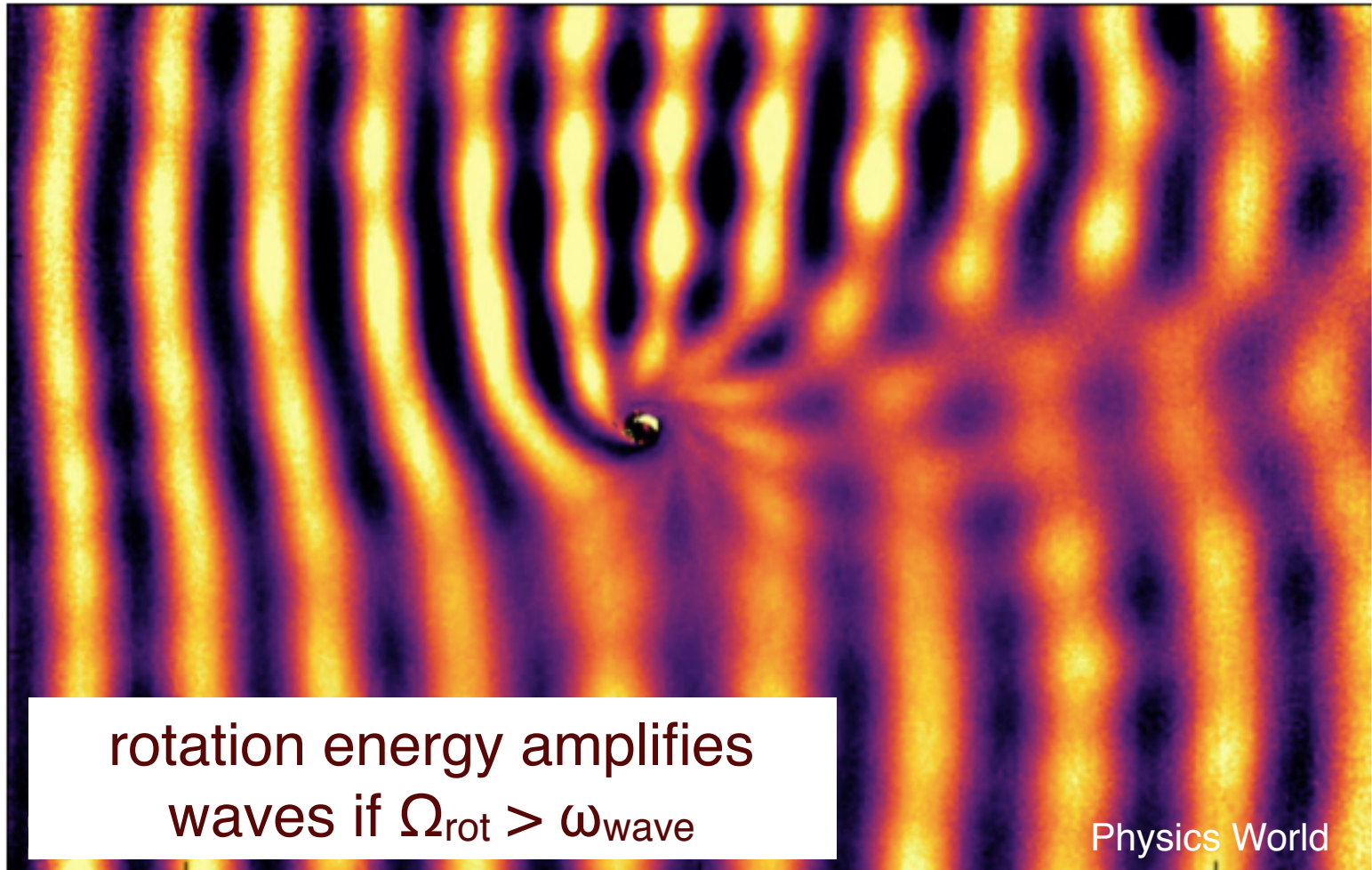
momentum conservation  
+  
interference of matter waves

Inouye, Chikkatur, Stamper-Kurn,  
Stenger, Pritchard, Ketterle  
Science 23 (1999)



# Rotational superradiance spotted as water swirls down a drain

13 Jun 2017 Hamish Johnston



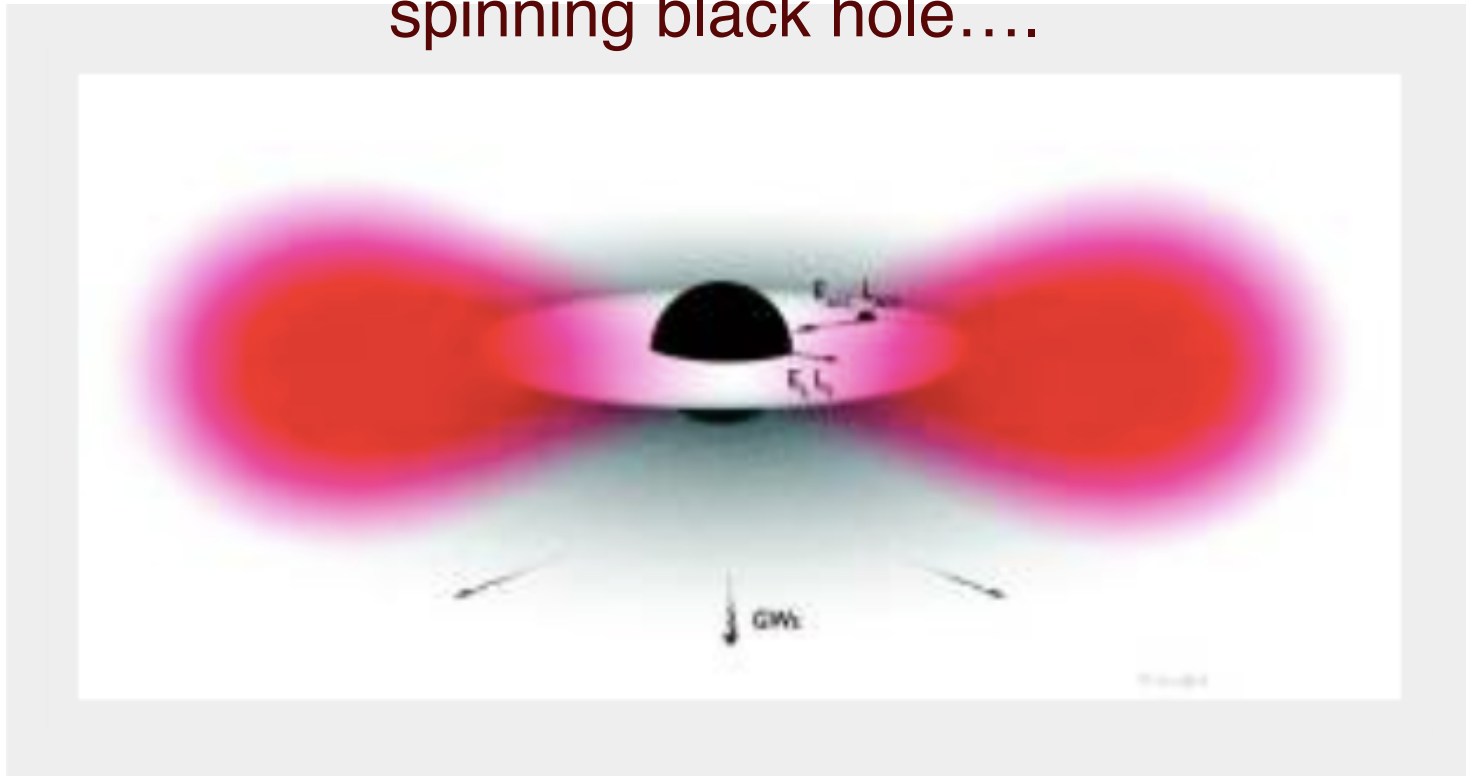
Whirling around: waves scatter from a vortex

Torres, Patrick, Coutant, Richartz, Tedford, Weinfurter  
Nature Physics 13, 833 (2017).

# Black hole superradiance

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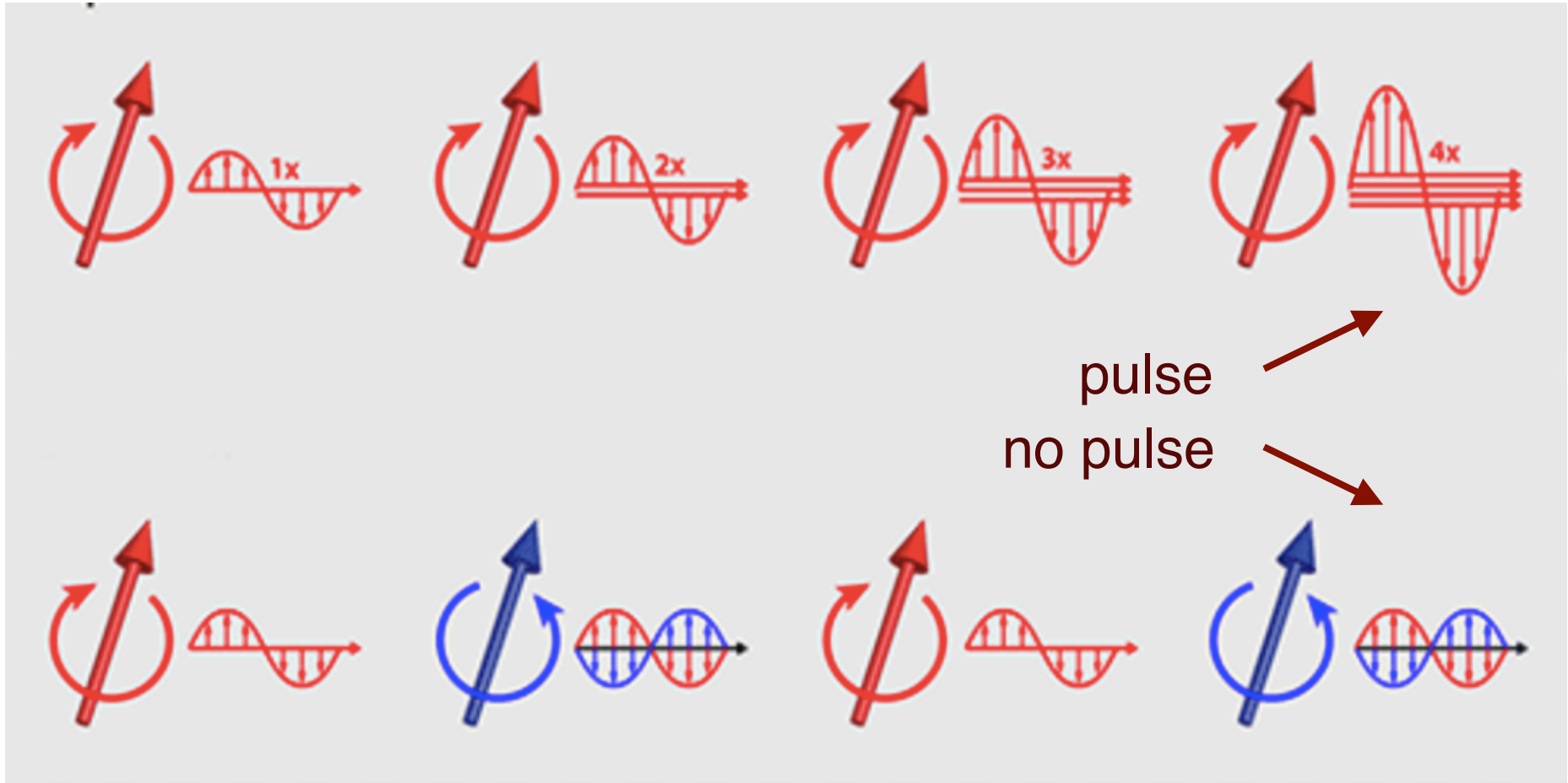
spinning black hole....



... amplifies gravitational waves  
(basically same as for water waves)



# Subradiance



# These lectures

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- Cooperative effects in complex systems
- New application: atomically thin mirrors

# These lectures

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- Cooperative effects in complex systems
  - ▶ Collective (Lamb) level shifts
  - ▶ Subradiance
  - ▶ Entanglement
- New application: atomically thin mirrors

# These lectures

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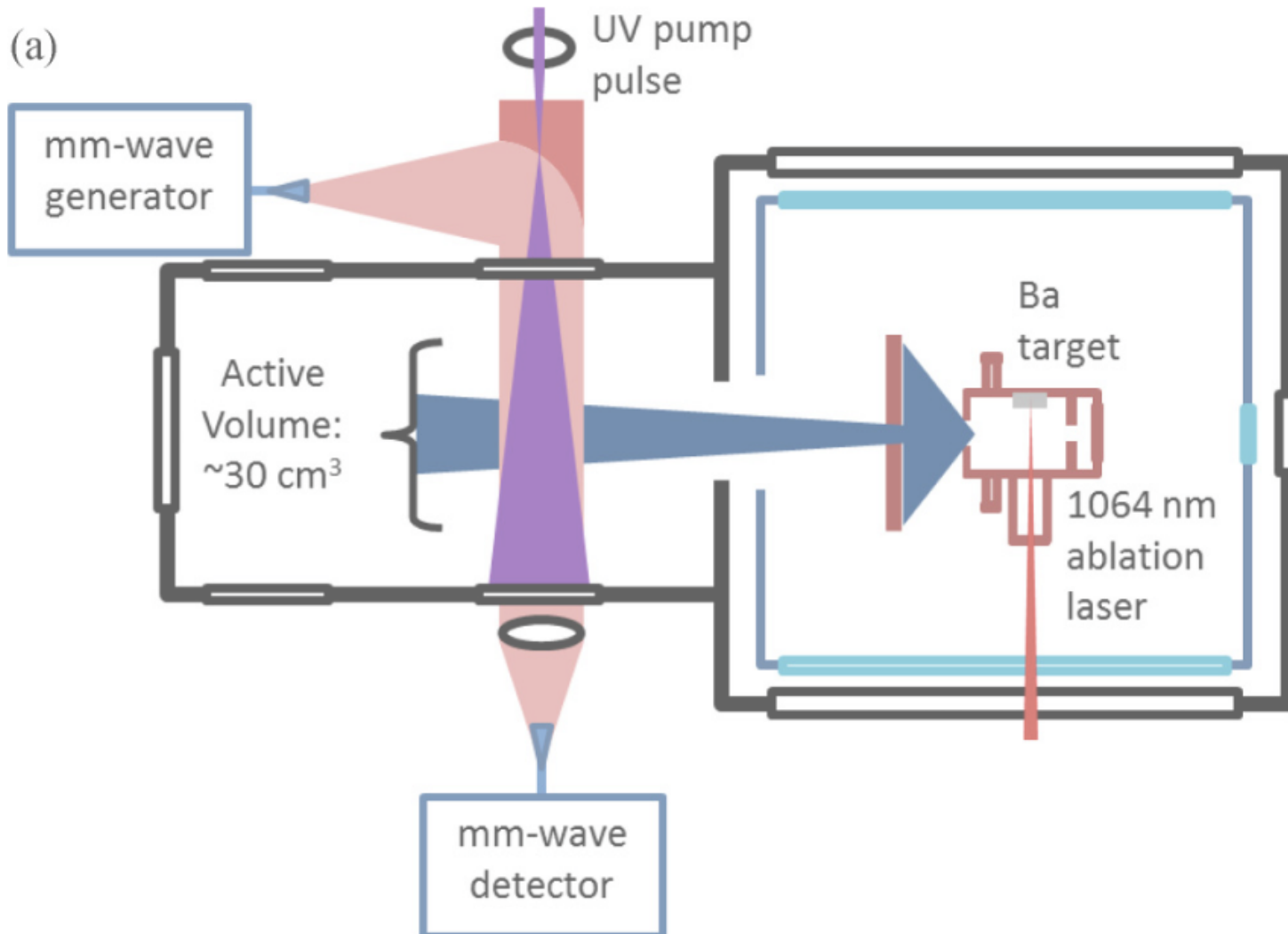
- Cooperative effects in complex systems
- New application: atomically thin mirrors
  - ▶ Cooperative resonances
  - ▶ Applications:

# These lectures

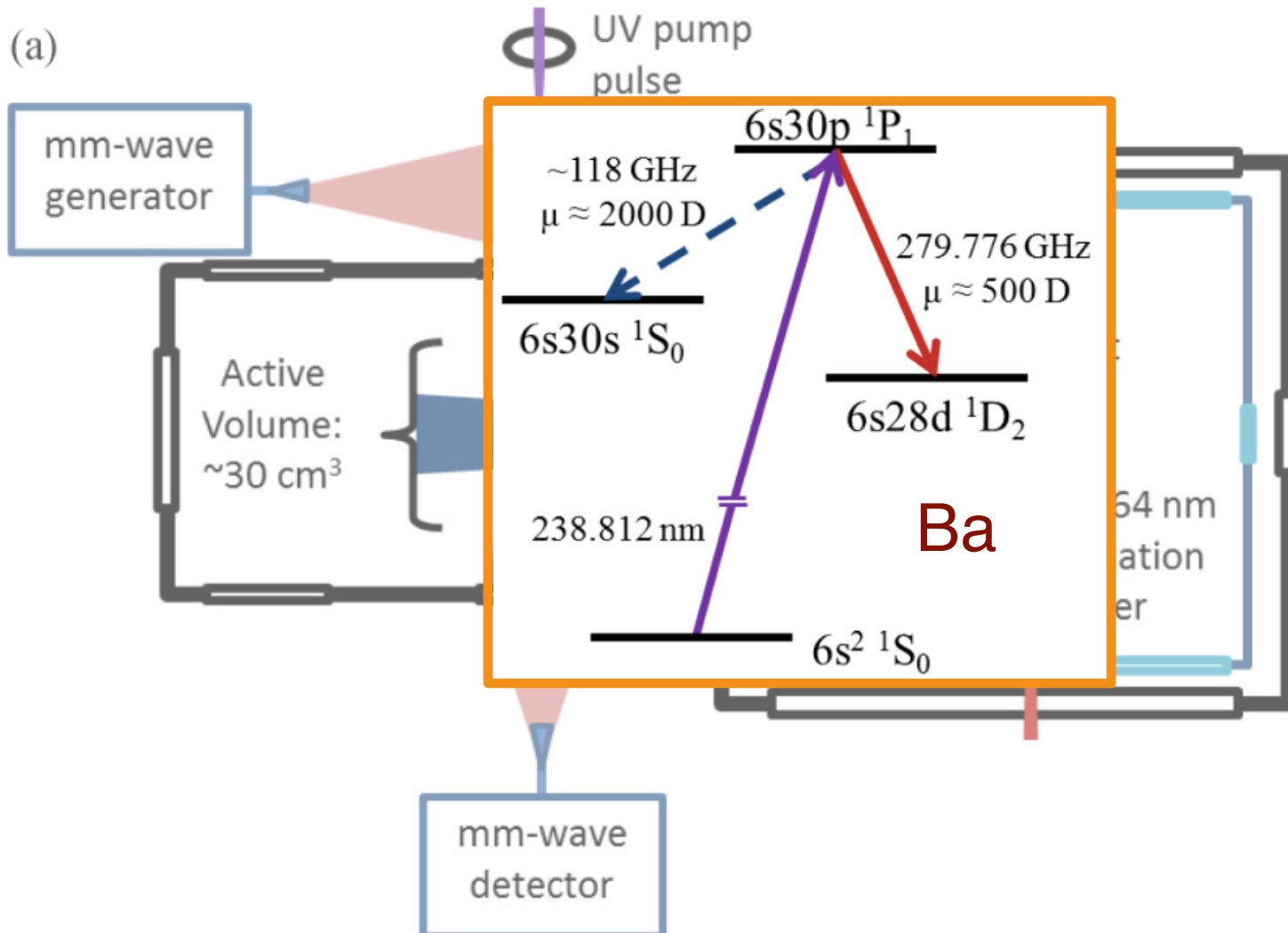
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- Cooperative effects in complex systems
- New application: atomically thin mirrors
  - ▶ Cooperative resonances
  - ▶ Applications:
    - topology with photons
    - nonlinear quantum optics
    - Quantum metasurfaces

# Superradiance: recent experiment

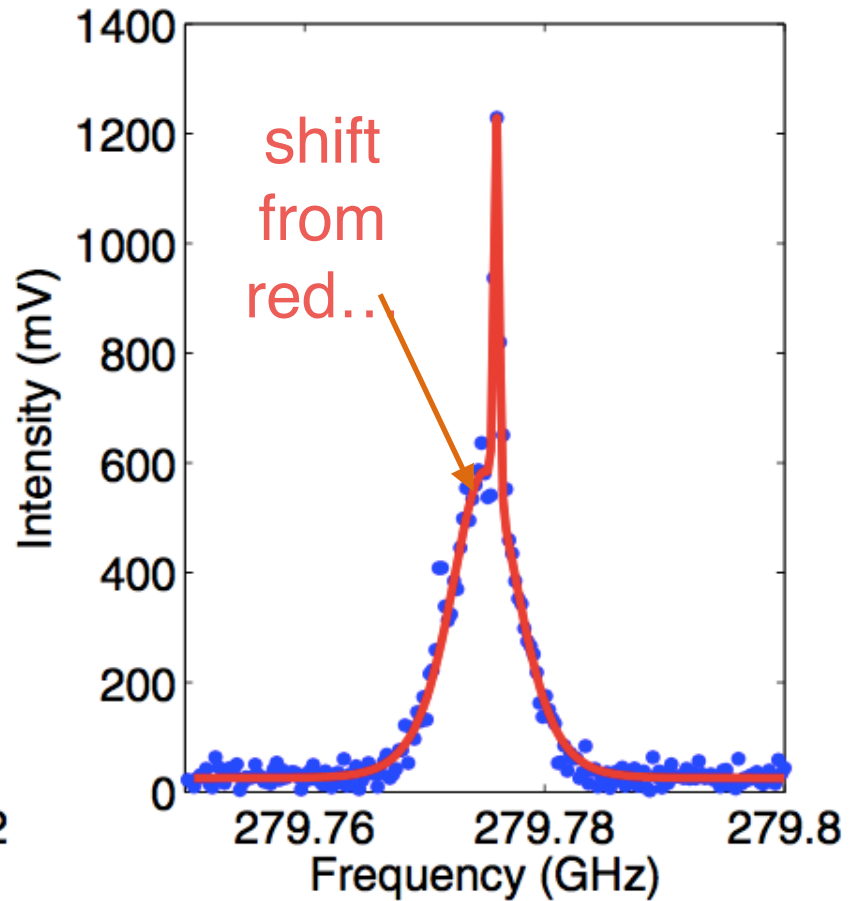
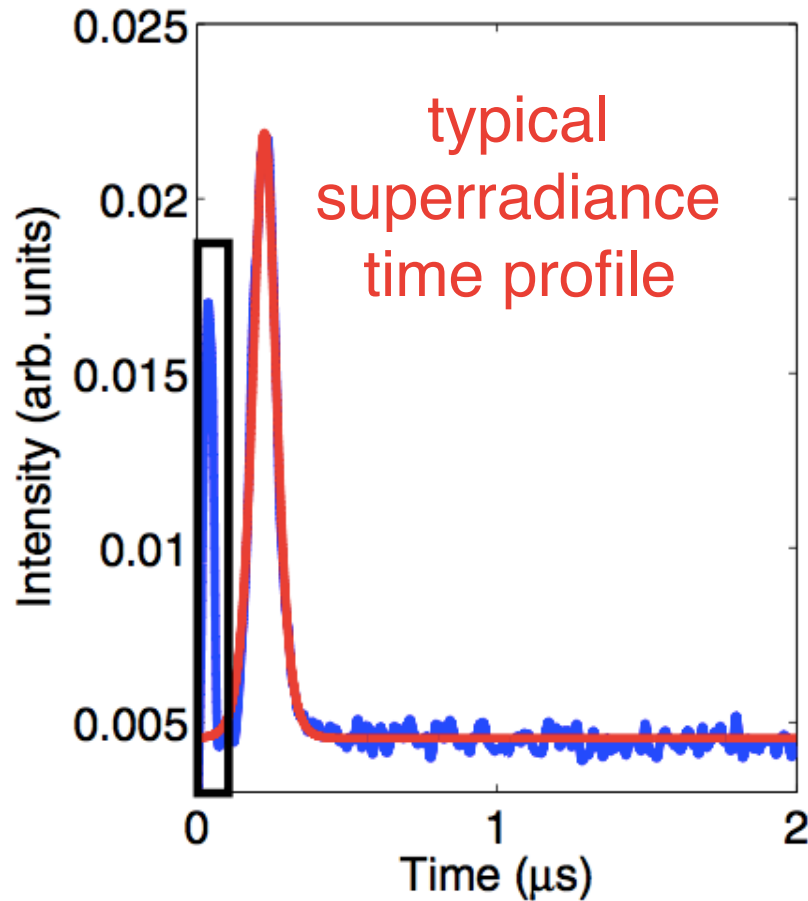


# Superradiance: recent experiment



# Superradiance in Ba Rydbergs

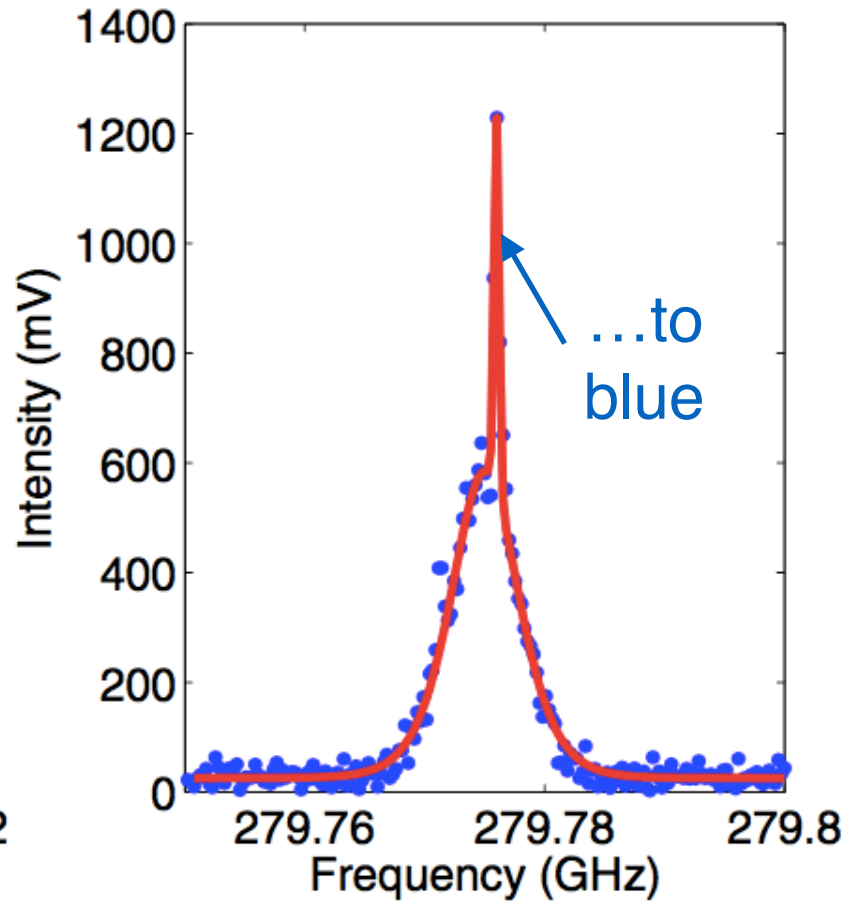
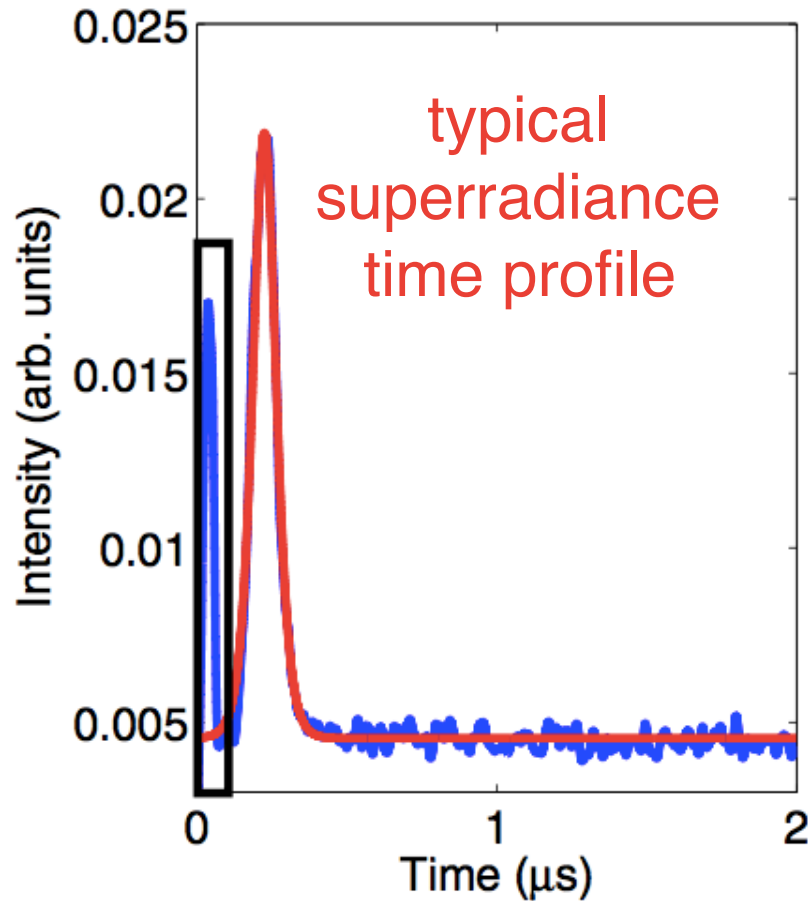
## Collective shift





# Superradiance in Ba Rydbergs

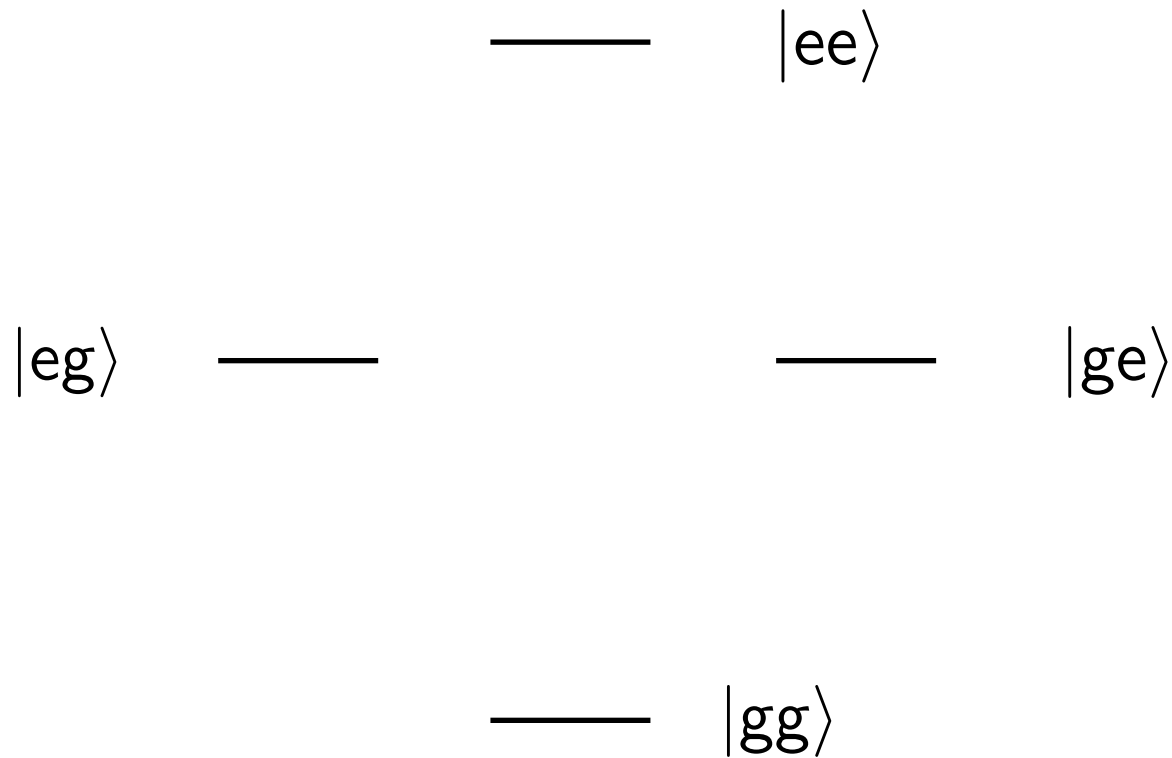
## Collective shift



# Cooperative radiation: two atoms

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atoms distinguishable



# Cooperative radiation: two atoms

---

atoms indistinguishable

$$\text{————} \quad |ee\rangle \equiv |1, 1\rangle$$

$$\frac{1}{\sqrt{2}} (|eg\rangle + |ge\rangle) \equiv |1, 0\rangle \text{————}$$

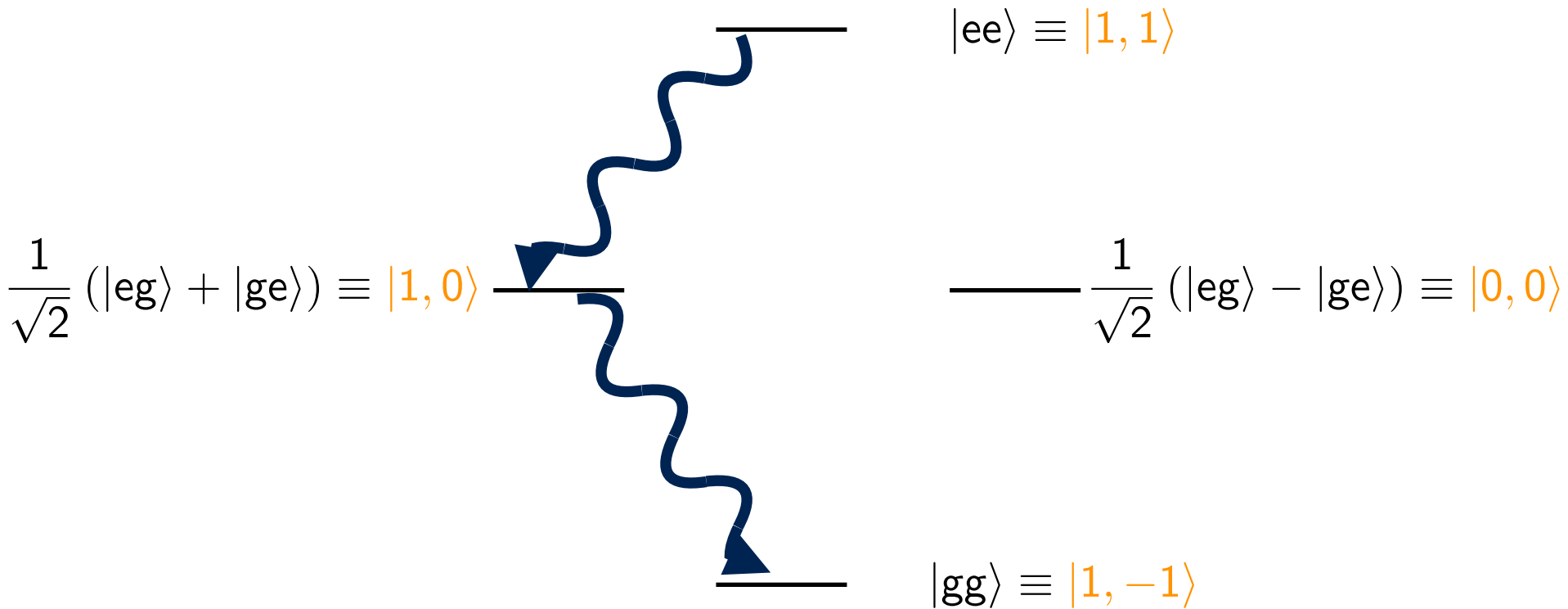
$$\text{————} \frac{1}{\sqrt{2}} (|eg\rangle - |ge\rangle) \equiv |0, 0\rangle$$

$$\text{————} \quad |gg\rangle \equiv |1, -1\rangle$$

# Cooperative radiation: two atoms

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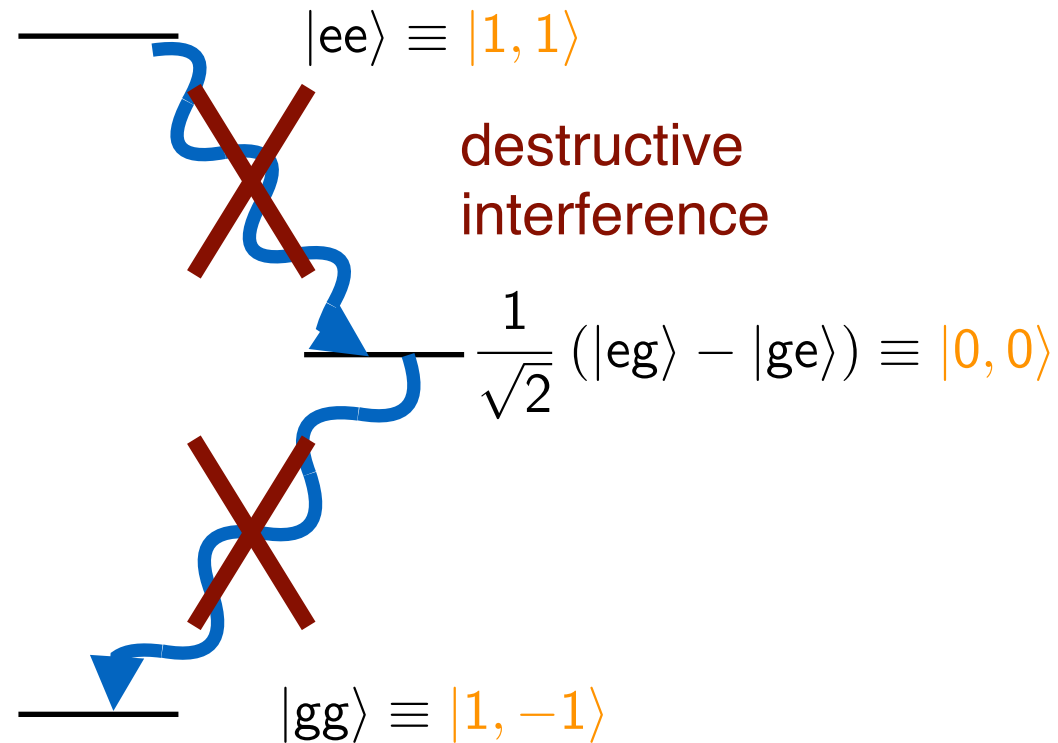
atoms indistinguishable



# Cooperative radiation: two atoms

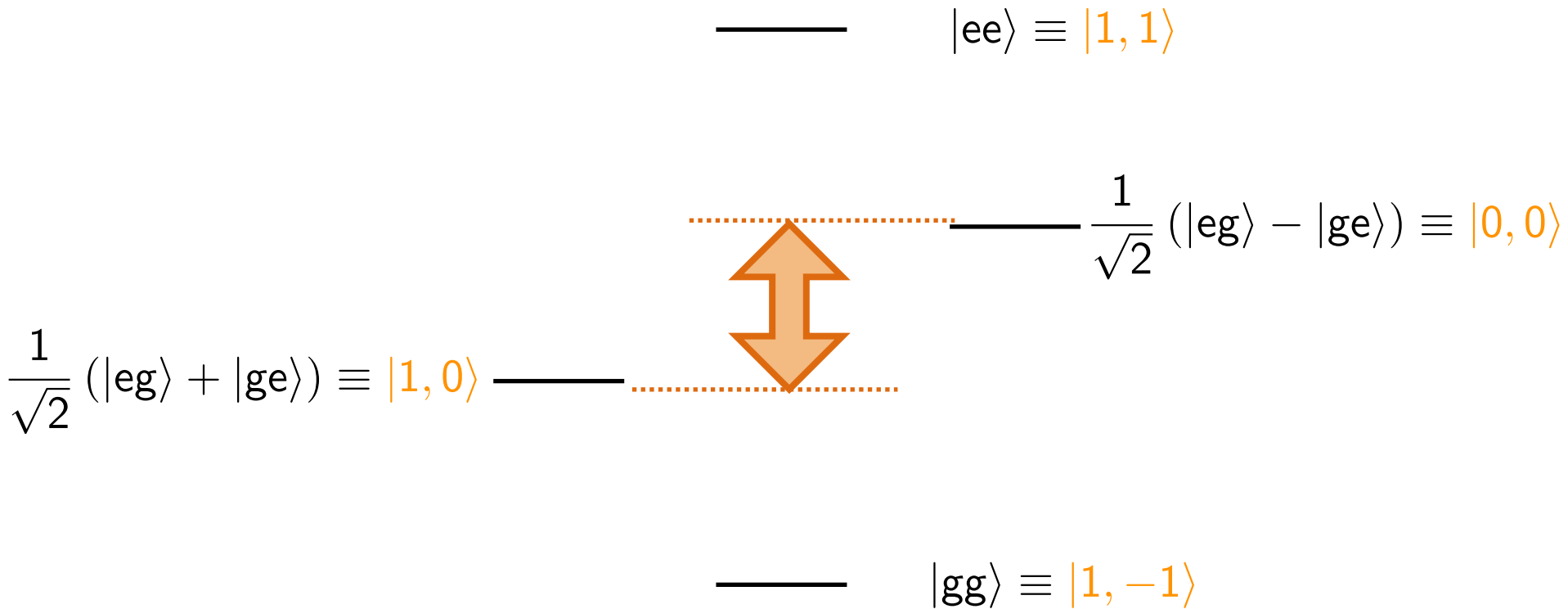
atoms indistinguishable

$$\frac{1}{\sqrt{2}} (|eg\rangle + |ge\rangle) \equiv |1, 0\rangle$$



# Cooperative radiation: two atoms

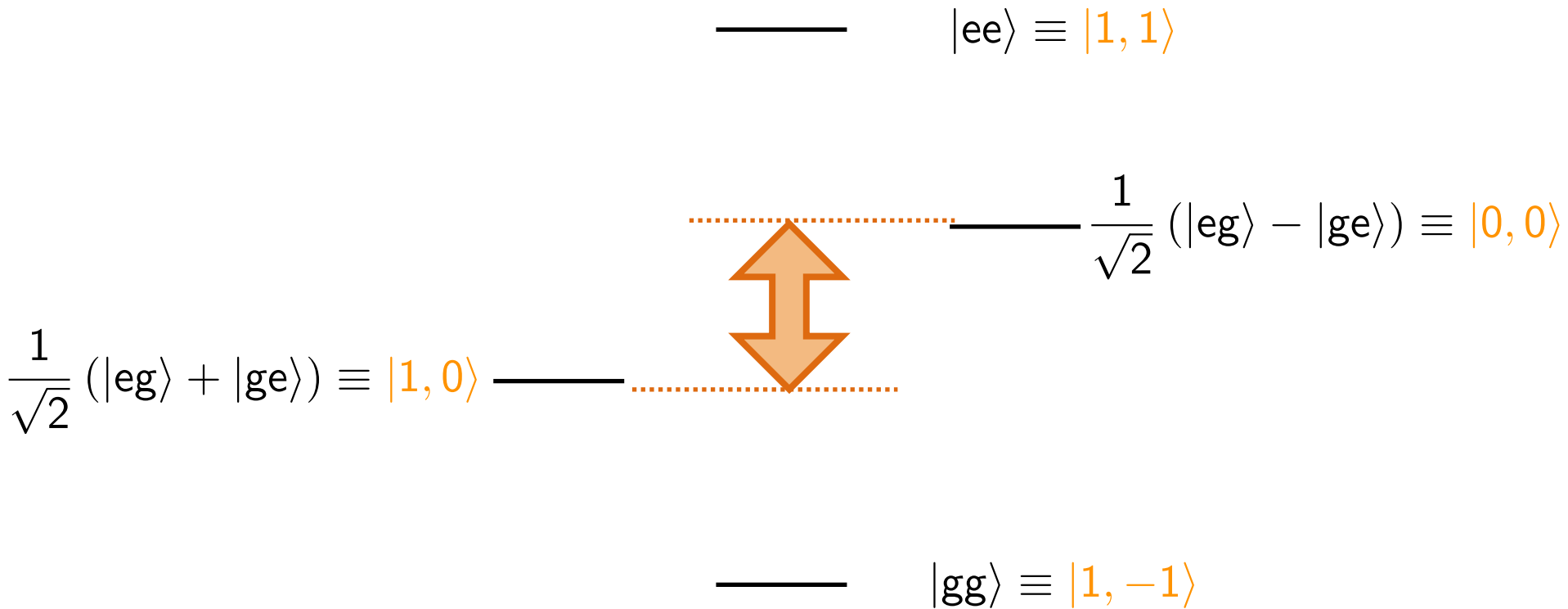
atoms indistinguishable



dipole-dipole exchange interaction

# Cooperative radiation: two atoms

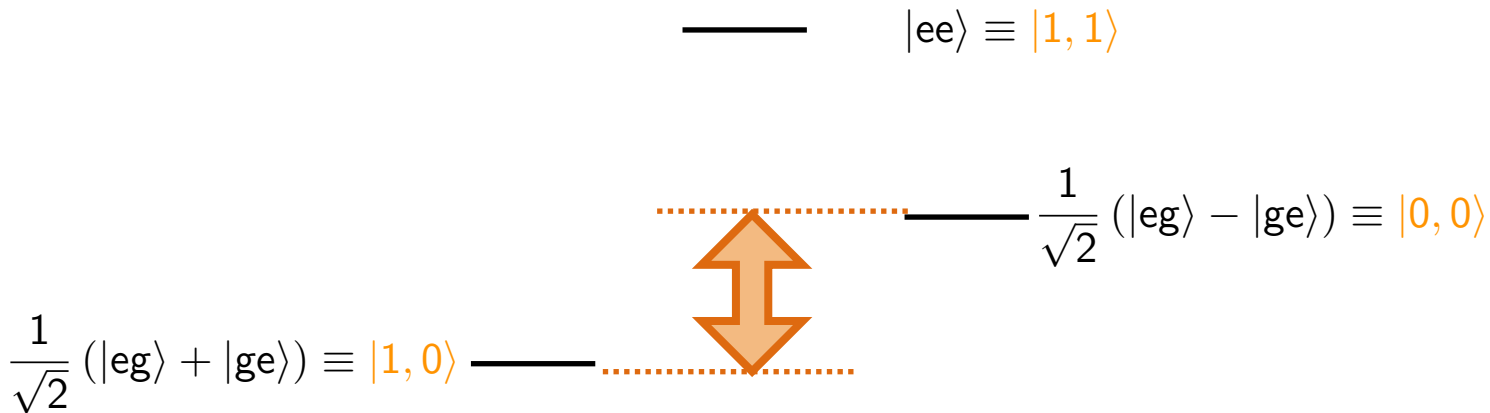
atoms indistinguishable



dipole-dipole exchange interaction

# Cooperative radiation: two atoms

atoms indistinguishable



exchange interaction:

- usually dipole-dipole mediated
- creates shift and broadening (Kramers-Kronig)

dipole



# What is “superradiance”?

---

1. Everything that involves Dicke states
  - (e.g., collective  $\sqrt{N}$  effects,
  - bad-cavity limit,
  - ...)

# Dicke states

---

Fully symmetric state of  $n$  excitations in  $N$  particles, for example

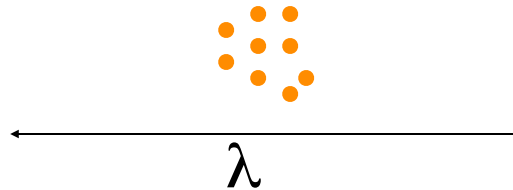
$$|2\rangle_4 = \frac{1}{\sqrt{6}} (|1100\rangle + |1010\rangle + |1001\rangle + |0110\rangle + |0101\rangle + |0011\rangle)$$

$N$ -particle Dicke states decay with up to  $N^2$  speedup

# Dicke states

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- Question: When do we have a system that consists only of Dicke states?
- Answer 1: When there exists no mechanism to distinguish atoms. Example:



- Problem: interactions (dip-dip) drive system out of purely symmetric state! How to deal?
- (Answer 2: When the exchange interaction is infinitely high. In this case, other states cannot be reached. Example: “many-body protected manifolds”)

# Describing superradiance

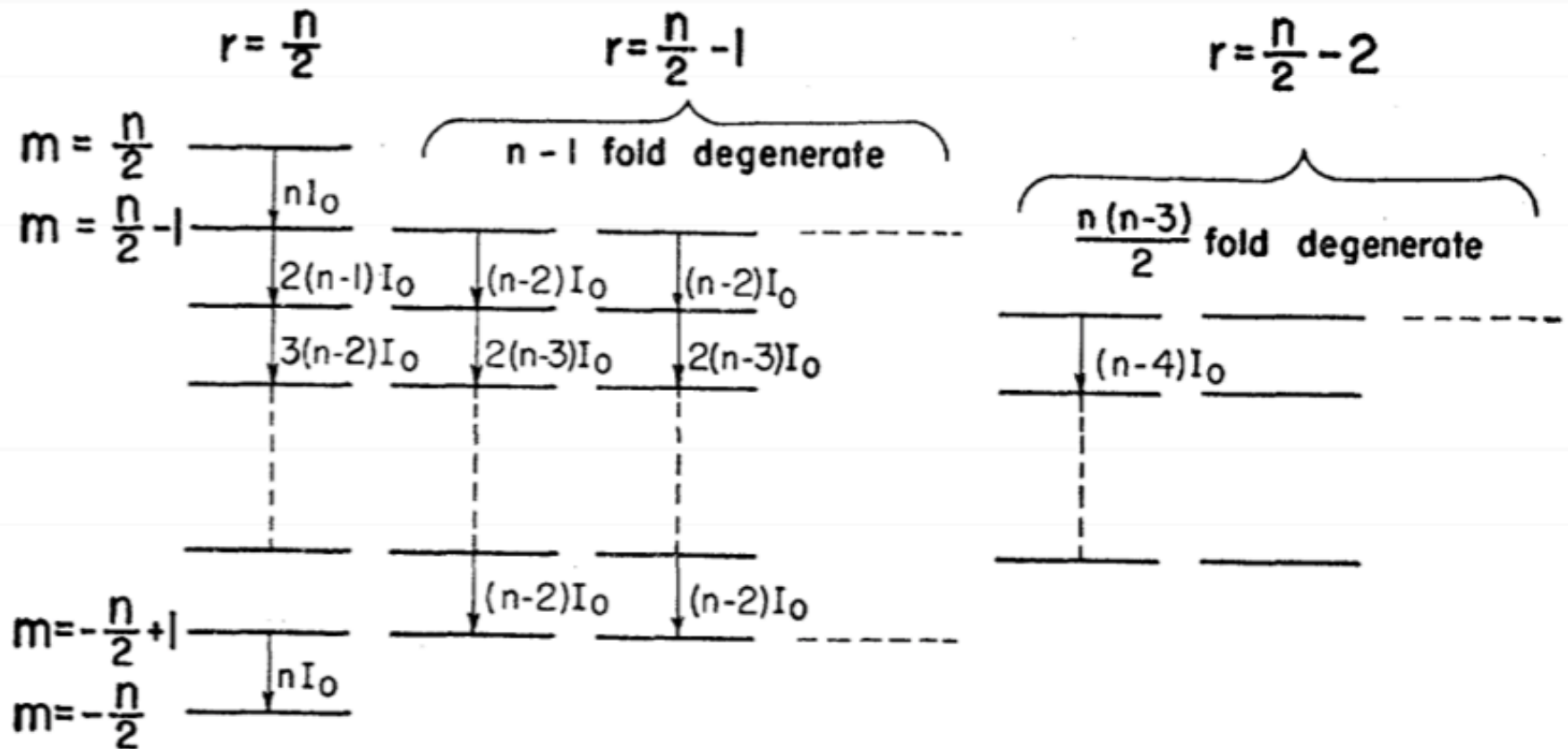
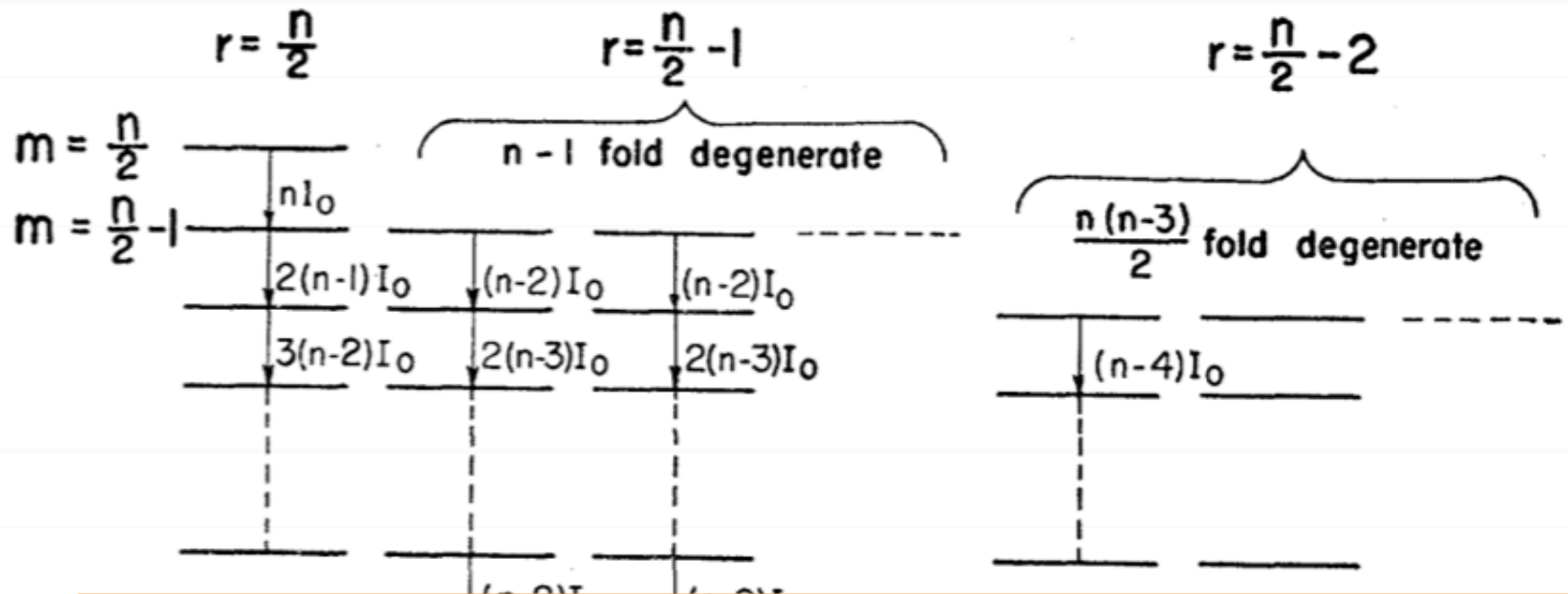


FIG. 1. Energy level diagram of an  $n$ -molecule gas, each molecule having 2 nondegenerate energy levels. Spontaneous radiation rates are indicated.  $E_m = mE$ .

# Describing superradiance



Use angular momentum form:  
 $|J, m\rangle$  denotes system with  $m$  excitations,  $m = -J \dots J$ .

FIG. 1. Energy level diagram of an  $n$ -molecule gas, each molecule having 2 nondegenerate energy levels. Spontaneous radiation rates are indicated.  $E_m = mE$ .

# Angular momentum states

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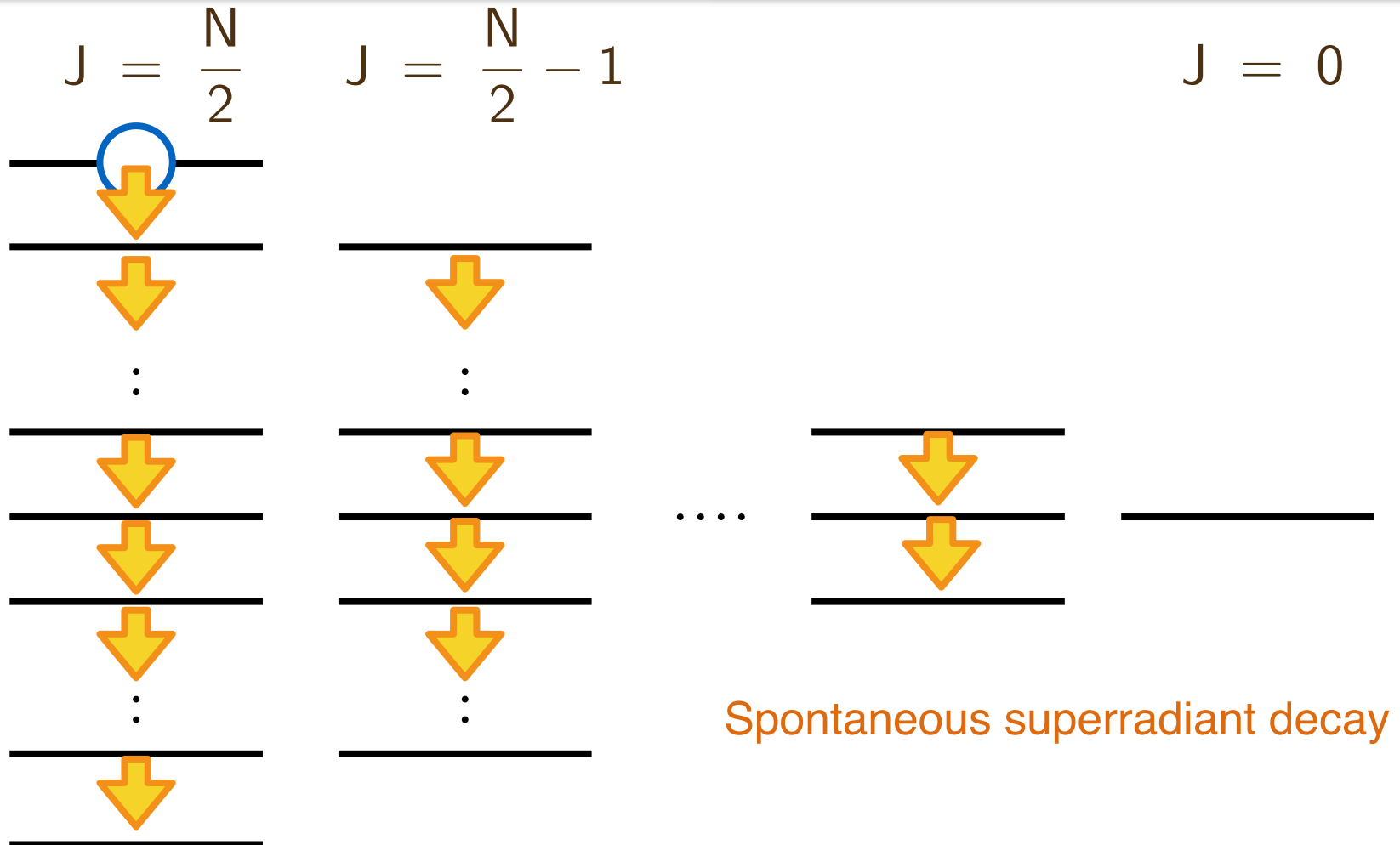
- Form of 3-atom Dicke state?
- What are the other states?
- 4 atoms?
- 40?

# Population of symmetric states

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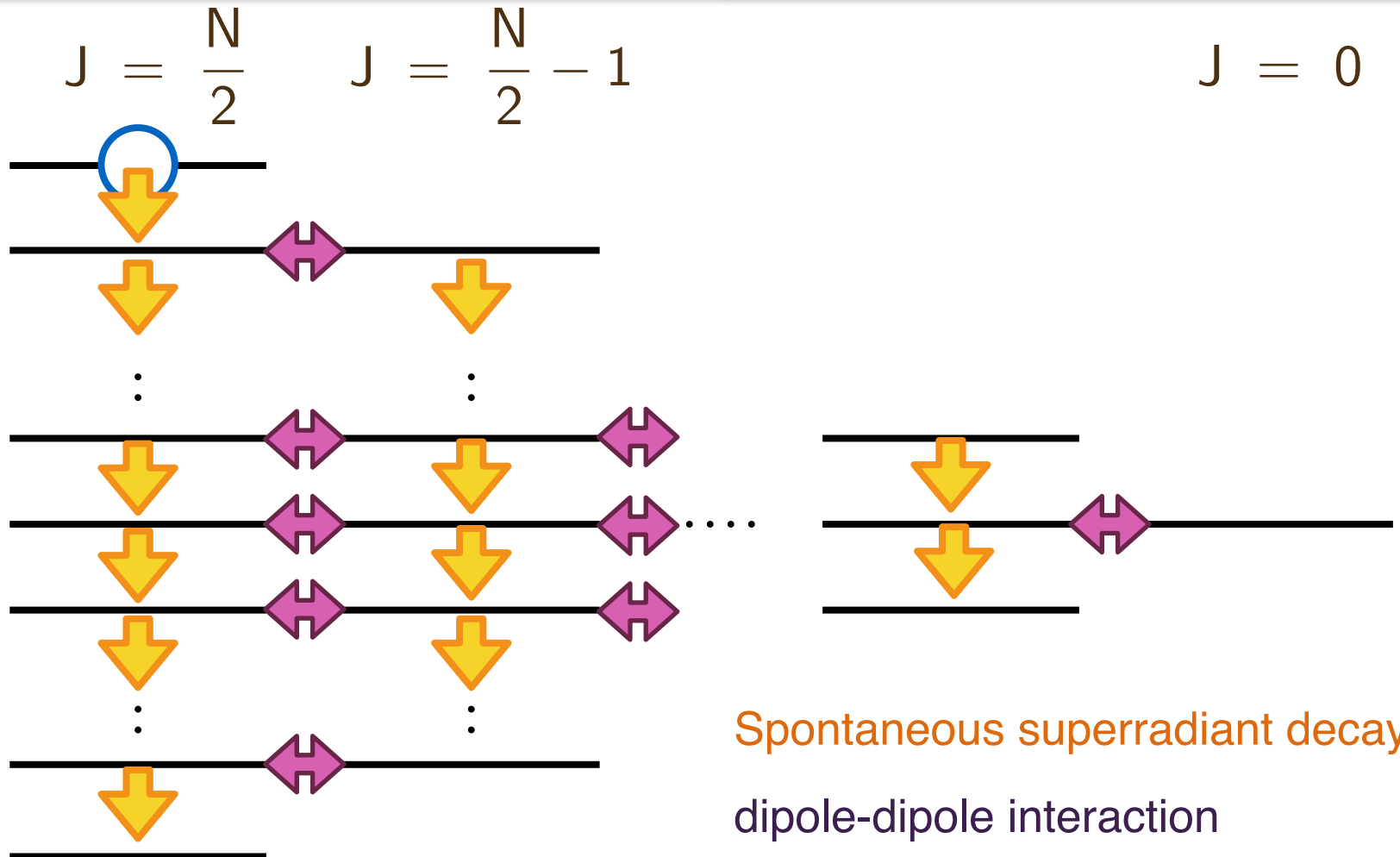
- What possible path could a system starting in  $|11111\dots\rangle$  take,
- with only decay?
- when there is dipole-dipole interaction?

# State connections





# State connections



# Form of dipole-dipole interaction

---

$$V_{\text{dip-dip}} \propto \sum_{i \neq j} \frac{e^{i\theta_{ij}}}{x_{ij}^3} [(1 - \cos^2 \theta_{ij}) x_{ij}^2 + (1 - 3 \cos^2 \theta_{ij}) (ix_{ij} - 1)]$$

$$H_{\text{dip-dip}} = \sum_{i \neq j} V_{ij} \sigma_i^+ \sigma_j^-$$

# “Exchange” term

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- How would this show up in a master equation?

# Form of dipole-dipole interaction

---

$$V_{\text{dip-dip}} \propto \sum_{i \neq j} \frac{e^{i\theta_{ij}}}{x_{ij}^3} [(1 - \cos^2 \theta_{ij}) x_{ij}^2 + (1 - 3 \cos^2 \theta_{ij}) (ix_{ij} - 1)]$$

dipole-dipole interaction:

- distance dependence
- angle dependence
- real + imaginary part

↑  
virtual  
photon exchange

↑  
real

flip-flop  
“exchange”

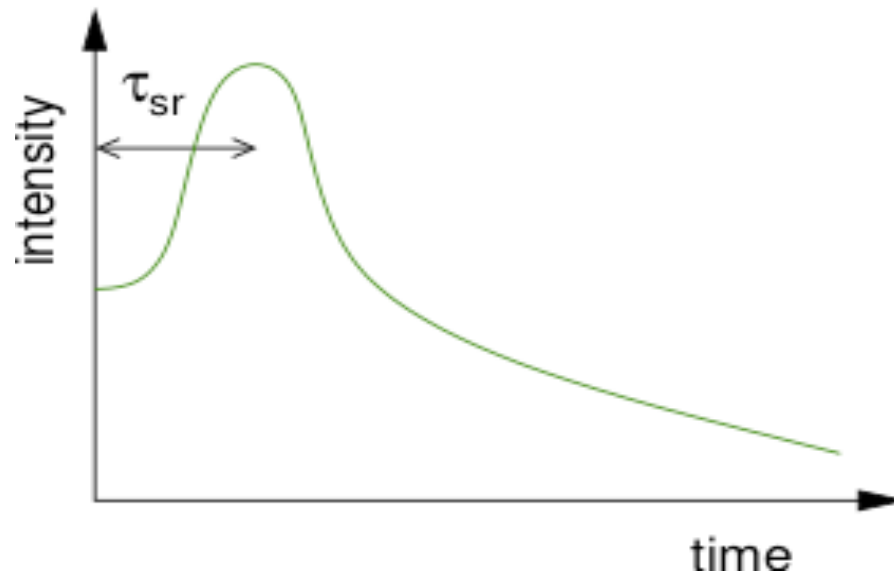
shift/  
dephasing

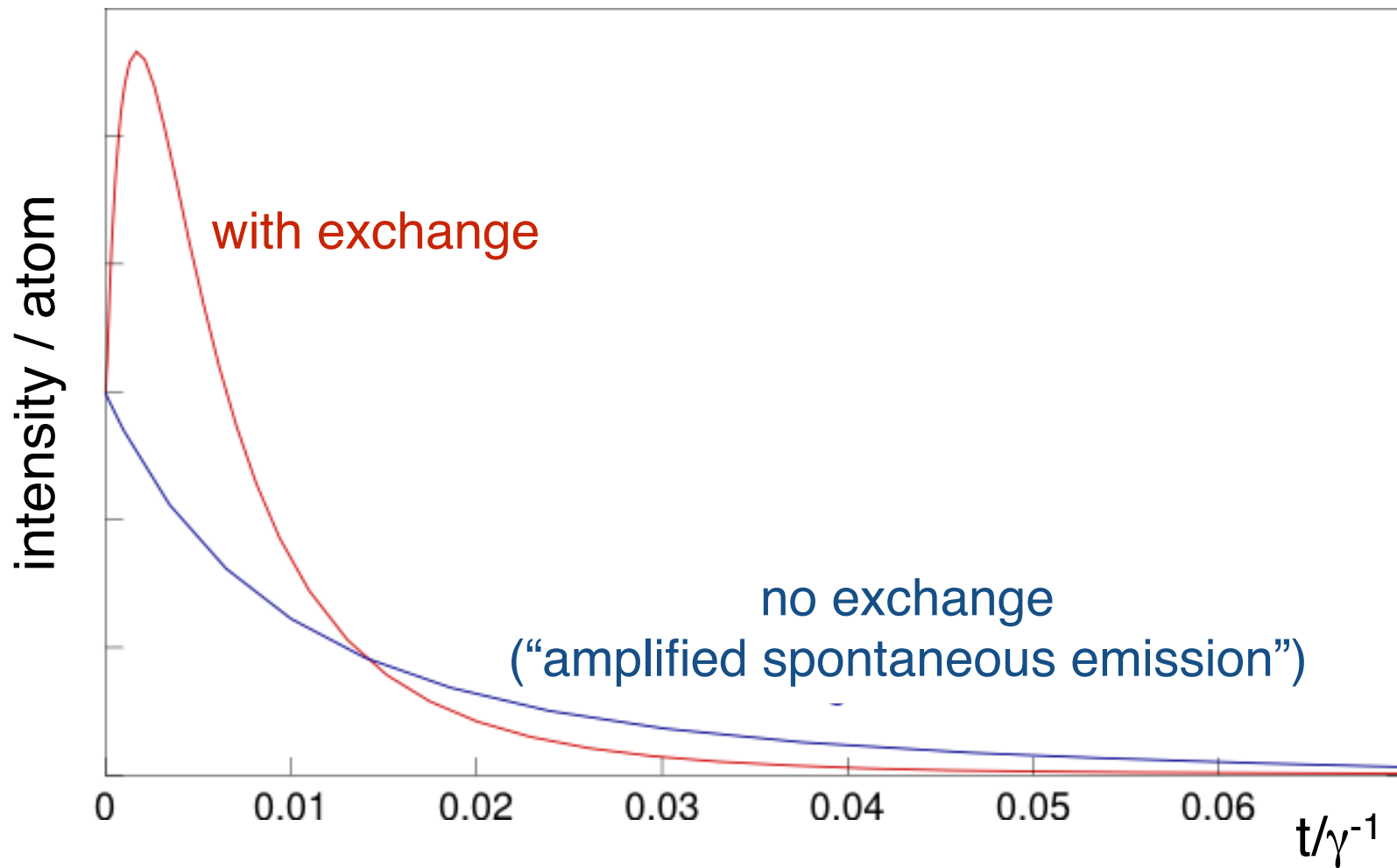
# Atom-atom correlations in superradiance: Classic example



$\lambda$

- Superradiance





# What is “superradiance”?

---

1. Everything that involves Dicke states
  - (e.g., collective  $\sqrt{N}$  effects,
  - bad-cavity limit,
  - ...)
2. Only systems involving cooperative (and nonlinear) effects
  - i.e., effect of exchange interaction
  - more than single excitation

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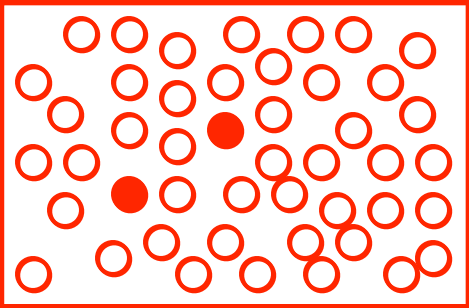
only for purists



# Dynamics of atoms in dense media - Schwinger-Keldysh & Dyson Eq.

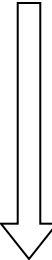
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Full dynamics (all degrees of freedom of atoms, fields)



two probe atoms  
+  
surrounding atoms

Two atoms + field

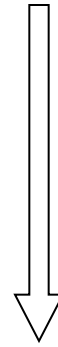
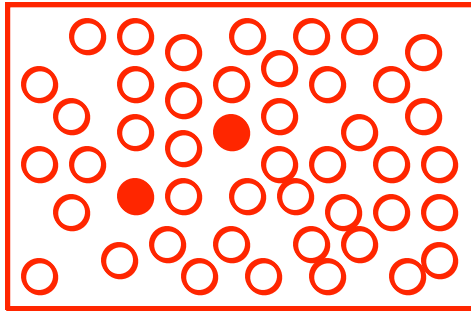


effective two-atom description

# Dynamics of atoms in dense media - Schwinger-Keldysh & Dyson Eq.

Full dynamics (all degrees of freedom of atoms, fields)

$$H = H_{\text{atoms}} + H_{\text{field}} - \sum \mathbf{p}_i \mathbf{E}_i$$



two probe atoms  
+  
surrounding atoms

Two atoms + field

$$V_{\text{probe}} = \sum_{i=1,2} \mathbf{p}_i \mathbf{E}_i \Rightarrow S = T e^{-\frac{i}{\hbar} \int d\tau V_{\text{probe}}(\tau)}$$

- $\langle e^{sX} \rangle = e^{\sum_{m=1,2} \frac{s^m}{m!} \langle\langle X^m \rangle\rangle}$

- Gauss:  $m \leq 2$



field degrees  
of freedom

effective two-atom description

Two atom Master equation

$$\int \langle\langle \hat{E}_i(t_1) \hat{E}_j(t_2) \rangle\rangle$$

trace out field  
degrees of  
freedom

$$V_{\text{dip-dip}} \propto \sum_{i \neq j} \frac{e^{i\theta_{ij}}}{x_{ij}^3} [(1 - \cos^2 \theta_{ij}) x_{ij}^2 + (1 - 3 \cos^2 \theta_{ij}) (ix_{ij} - 1)]$$

# Can one expect superradiance?

Dynamics of atoms in dense media - Schwinger-Keldysh & Dyson Eq

Full dynamics (all degrees of freedom of atoms, fields)

$$H = H_{\text{atoms}} + H_{\text{field}} - \sum_i p_i E_i$$

Two probe atoms + surrounding atoms

Two atoms + field

$$V_{\text{probe}} = \sum_{i,j} p_i E_{ij} \Rightarrow S = T a^{-1} / \det V_{\text{atom}}(\omega)$$

•  $\langle n^2 \rangle = \sum_{i,j} S_{ij} S_{ji}$

• Coher:  $\rho = 0.5$

effective two-atom description  
Two atom Master equation

Symmetric small sample approximation

$$\begin{aligned} \dot{a} &= -(2\Gamma + \gamma) a + \Gamma \\ \dot{n} &= -2(2\Gamma + \gamma) n - 2\gamma(2a - 1) + 8\Gamma x \\ \dot{x} &= -(2\Gamma + \gamma) x + \Gamma n \end{aligned}$$

where  
 $a$ : average excited state population  
 $n$ : two-atom "inversion"  
 $x$ : correlation

Form of  $\Gamma, \Delta$

Solve Dyson equations via series algebra via  $\Gamma$   
"Branens-Kuang" integral via  $\Delta$

$$\Gamma = \frac{\gamma}{2a-1} (e^{2\Delta})$$

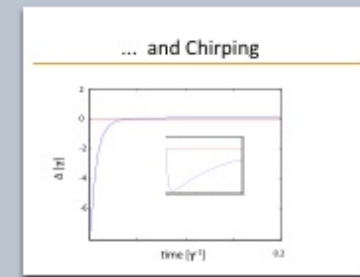
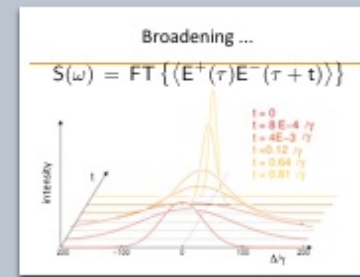
$$\tilde{\Gamma} = 3\gamma \frac{e^{2\Delta}}{1+e^{2\Delta}}$$

$$\Delta = \gamma \int_0^{\infty} dt e^{-\gamma t} \dots$$

optical density

$$O.D. = n \lambda^2 r$$

with  $C = \frac{\gamma}{2(2a-1)}$  and  $f_1 = \frac{\gamma}{2} + \tilde{\Gamma} C$



The important parameter is

$$n \lambda^2 r$$

optical depth

$n$ : density,  $\lambda$ : wavelength,  $r$ : system size

# Can one expect superradiance?

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where  
 $\tilde{n}$ : average excited state population  
 $\tilde{x}$ : two-atom "inversion"  
 $x$ : correlation

Form of  $\Gamma, \Delta$

Solve Dyson equations via series algebra via  $\Gamma$   
"Wannier-Kang" integral via  $\Delta$

$$\Gamma = \frac{\gamma}{2a-1} \langle n^2 \rangle$$

$$\tilde{\Gamma} = 3\gamma \frac{\langle n^2 \rangle}{\langle n \rangle}$$

$$\Delta = \frac{\gamma}{2} \langle n \rangle$$

optical density

$$O.D. = n \lambda^2 L$$

with  $C = \frac{\gamma}{2} \langle 2a - 1 \rangle$  and  $f_1 = \frac{\gamma}{2} \langle F \rangle$

Broadening ...

$$S(\omega) = \text{FT} \{ \langle \{ E^+(\tau) E^-(\tau + t) \} \rangle \}$$

Intensity vs  $\Delta\gamma$

$t = 0$   
 $t = 8 \text{ } \Gamma = 4 \text{ } \gamma$   
 $t = 4 \text{ } \Gamma = 3 \text{ } \gamma$   
 $t = 0.12 \text{ } \gamma$   
 $t = 0.04 \text{ } \gamma$   
 $t = 0.01 \text{ } \gamma$

... and Chirping

$\Delta |n\rangle$  vs time [ $\gamma^{-1}$ ]

The important parameter is

$$n \lambda^2 r$$

or  $n \lambda^3$  ?

optical depth

$n$ : density,  $\lambda$ : wavelength,  $r$ : system size

# Master Equation

---

$$\begin{aligned}
 \dot{\rho}_{i,j} = & \frac{i}{\hbar} \sum_{\mu} \wp_{\mu} \sum_{k=i,j} [\sigma_{k\mu} \mathcal{E}_{L,\mu}^{-}(\vec{r}_k) + \sigma_{k\mu}^{\dagger} \mathcal{E}_{L,\mu}^{+}(\vec{r}_k), \rho] \\
 & + \frac{i}{\hbar} \sum_{\mu,\nu} \sum_{k=i,j} H_{k\mu,k\nu} [[\sigma_{k\mu}, \sigma_{k\nu}^{\dagger}], \rho] \\
 & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu}}{2} \left( [\rho \sigma_{k\mu}, \sigma_{l\nu}^{\dagger}] + [\sigma_{k\mu}, \sigma_{l\nu}^{\dagger} \rho] \right) \\
 & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu} + \gamma_{k\mu,l\nu}}{2} \left( [\rho \sigma_{l\nu}^{\dagger}, \sigma_{k\mu}] + [\sigma_{l\nu}^{\dagger}, \sigma_{k\mu} \rho] \right)
 \end{aligned}$$

# Master Equation

$$\begin{aligned}
 \dot{\rho}_{i,j} = & \frac{i}{\hbar} \sum_{\mu} \vartheta_{\mu} \sum_{k=i,j} [\sigma_{k\mu} \mathcal{E}_{L,\mu}^{-}(\vec{r}_k) + \sigma_{k\mu}^{\dagger} \mathcal{E}_{L,\mu}^{+}(\vec{r}_k), \rho] \\
 & + \frac{i}{\hbar} \sum_{\mu,\nu} \sum_{k=i,j} H_{k\mu,k\nu} [[\sigma_{k\mu}, \sigma_{k\nu}^{\dagger}, \rho] \\
 & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu}}{2} ([\rho \sigma_{k\mu}, \sigma_{l\nu}^{\dagger}] \\
 & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu} + \gamma_{k\mu,l\nu}}{2} ([\rho \sigma_{l\nu}^{\dagger}, \sigma_{k\mu}] + [\sigma_{l\nu}^{\dagger}, \sigma_{k\mu} \rho])
 \end{aligned}$$

$\langle\langle \hat{E}(t) \rangle\rangle$

Mast

$$\langle\langle \hat{E}_1(t_1) \hat{E}_2(t_2) \rangle\rangle$$

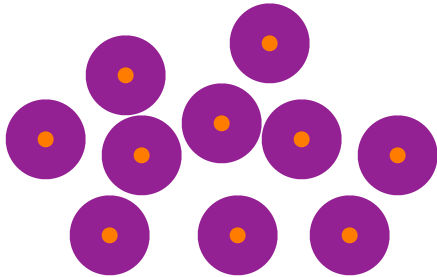
$$\begin{aligned} \dot{\rho}_{i,j} = & \frac{i}{\hbar} \sum_{\mu} \mathcal{D}_{\mu} \sum_{k=i,j} [\sigma_{k\mu} \mathcal{E}_{L,\mu}^{-}(\vec{r}_k) + \sigma_{k\mu}^{\dagger} \mathcal{E}_{L,\mu}^{+}(\vec{r}_k), \rho] \\ & + \frac{i}{\hbar} \sum_{\mu,\nu} \sum_{k=i,j} H_{k\mu,k\nu} [[\sigma_{k\mu}, \sigma_{k\nu}^{\dagger}] \rho] \\ & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu}}{2} ([\rho \sigma_{k\mu}, \sigma_{l\nu}^{\dagger}] + [\sigma_{l\nu}^{\dagger} \rho, \sigma_{k\mu}]) \\ & - \sum_{\mu,\nu} \sum_{k,l=i,j} \frac{\Gamma_{k\mu,l\nu} + \gamma_{k\mu,l\nu}}{2} ([\rho \sigma_{l\nu}^{\dagger}, \sigma_{k\mu}] + [\sigma_{l\nu}^{\dagger}, \sigma_{k\mu} \rho]) \end{aligned}$$

$$\langle\langle \hat{E}(t) \rangle\rangle$$

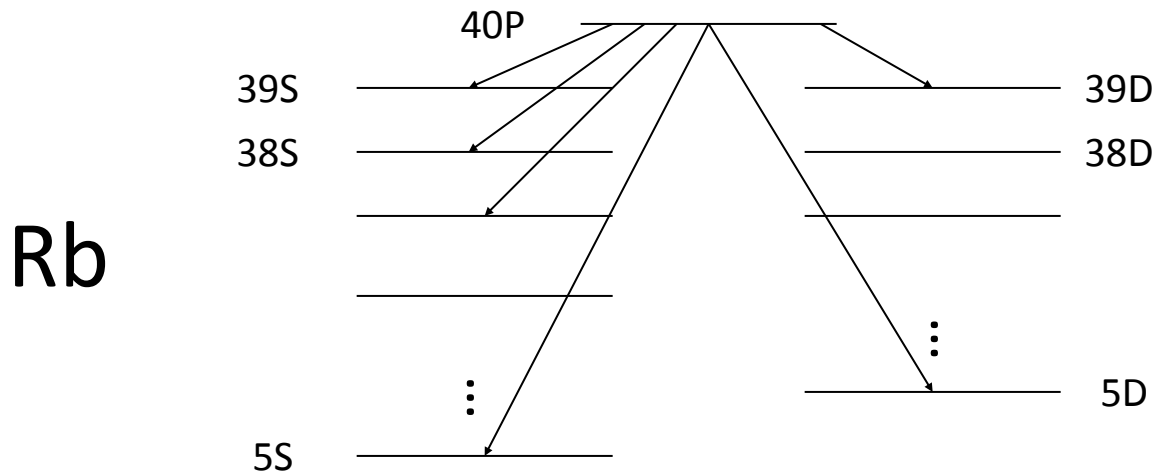


# New experimental systems: example

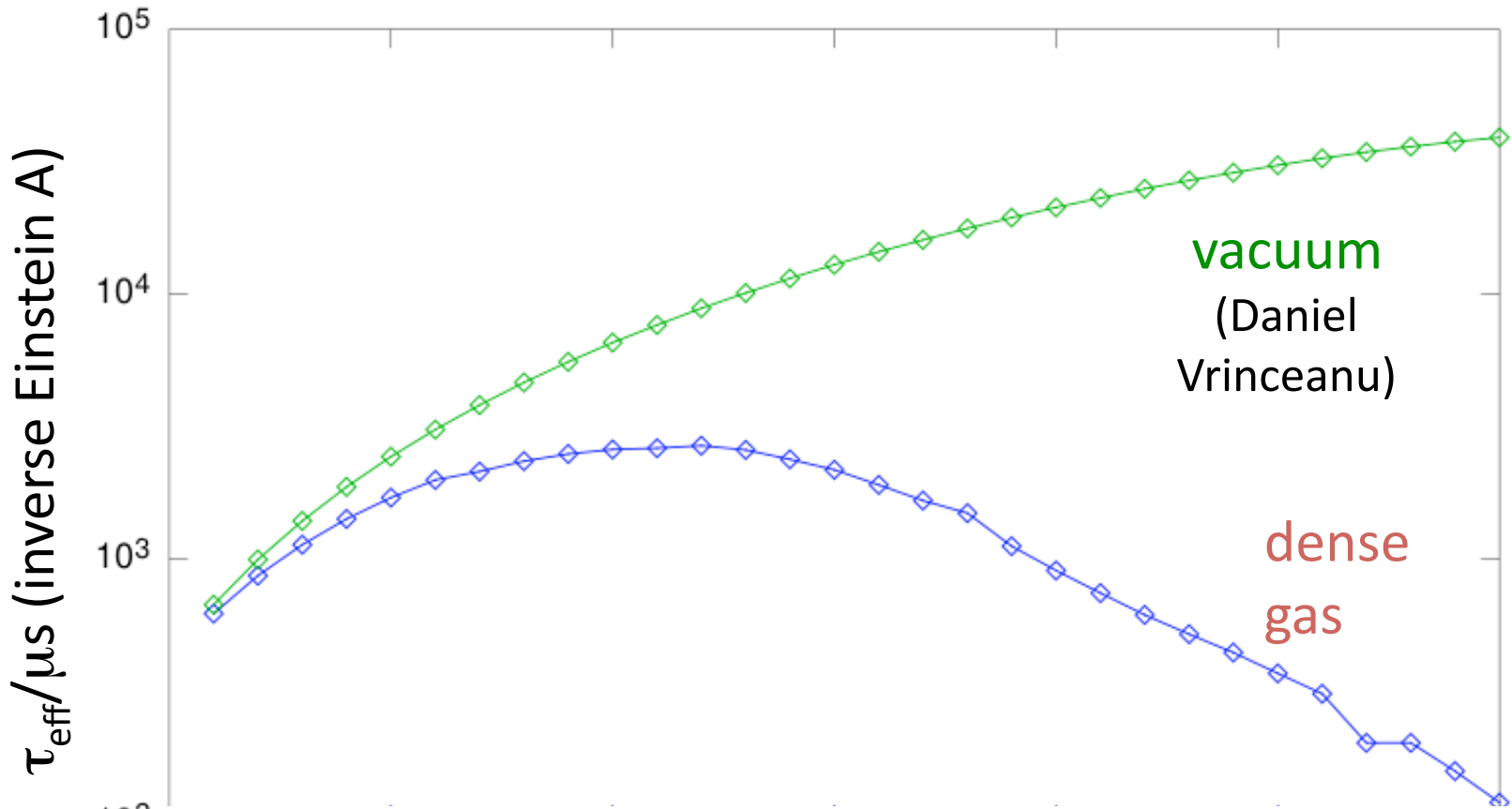
- Ultracold Rydberg atoms



(Phil Gould, Ed Eyler, Uconn)



# Effective decay times from 40P into nS

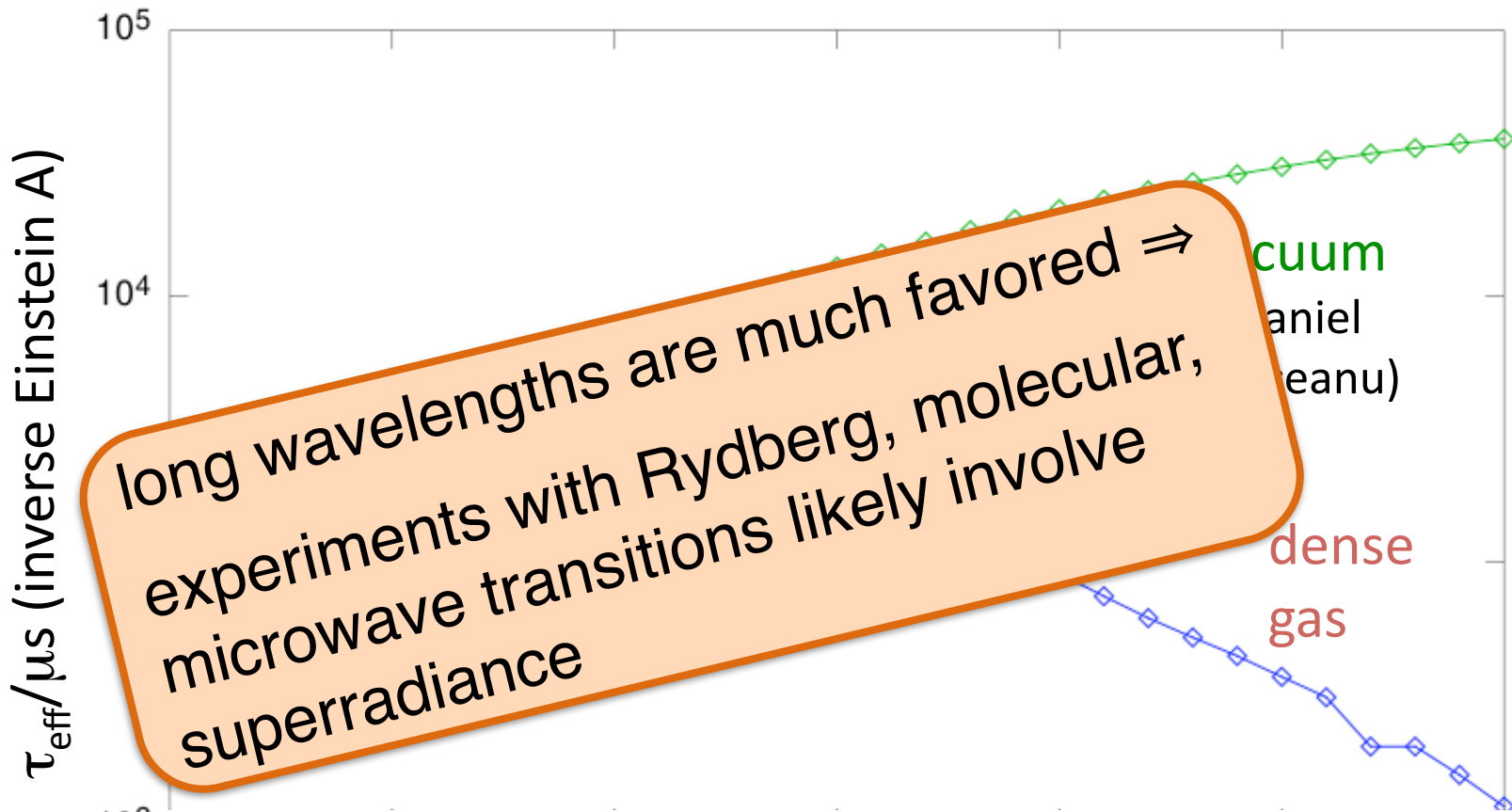


In vacuum: decay into low n is favored

In dense gas: decay into high n is favored  $\Rightarrow \lambda$  large,  $n \lambda^2 r$  large!

**superradiant decay!**

# Effective decay times from 40P into nS

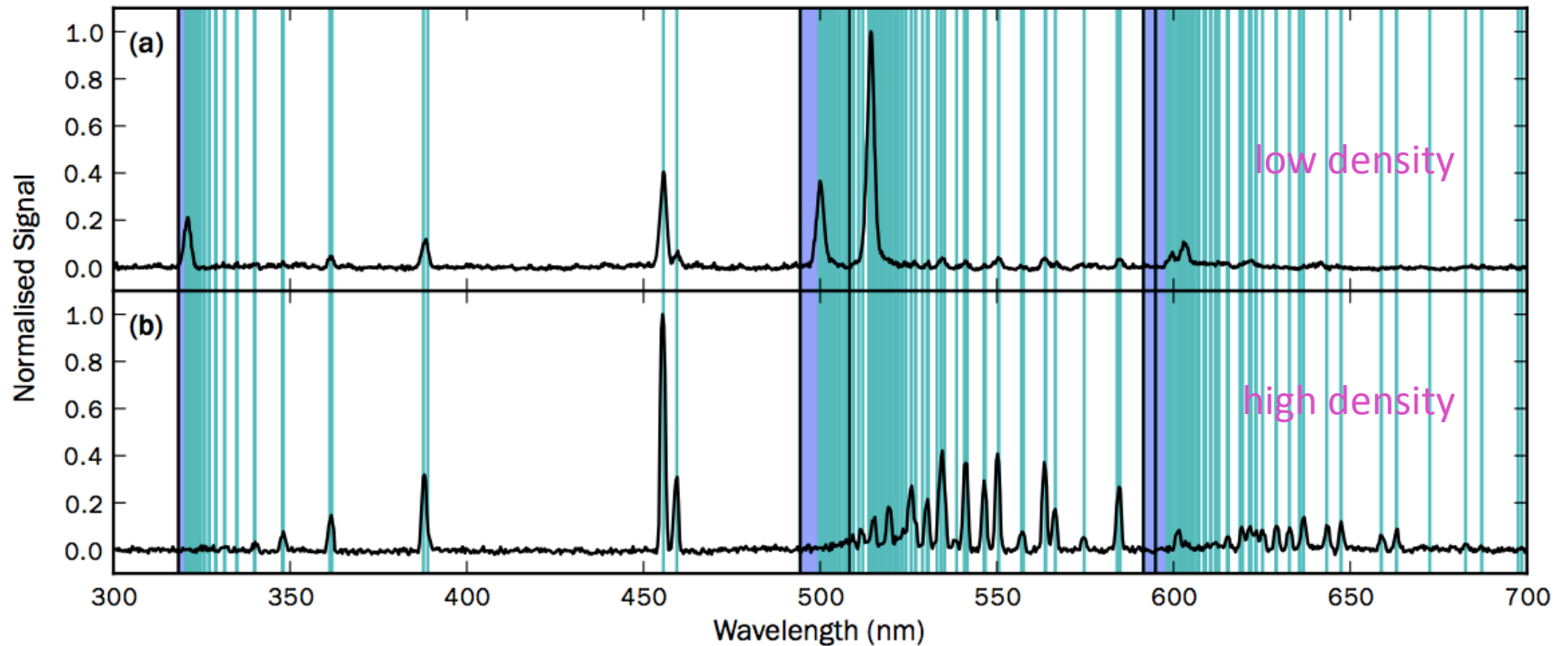


In vacuum: decay into low n is favored

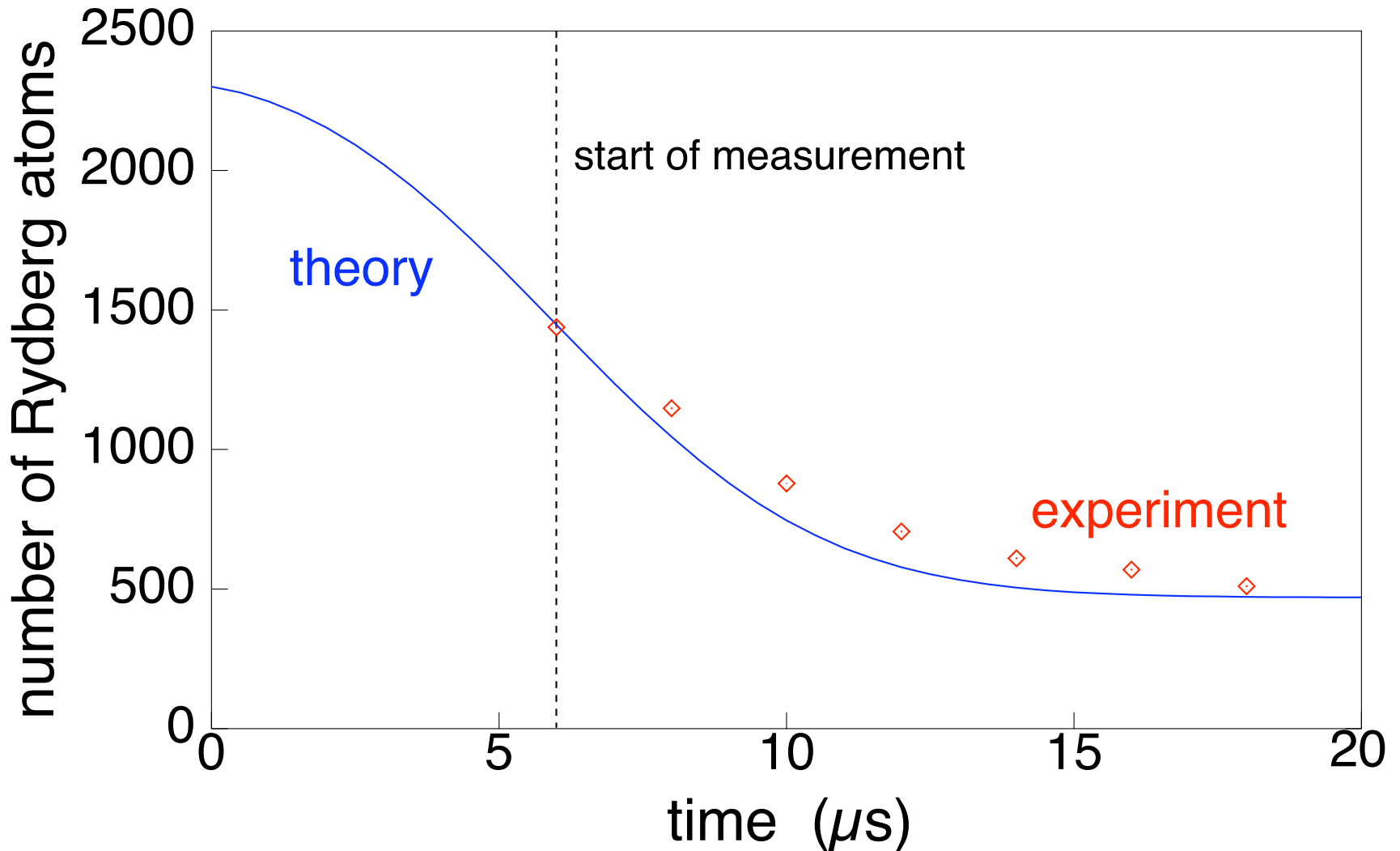
In dense gas: decay into high n is favored  $\Rightarrow \lambda$  large,  $n \lambda^2 r$  large!

**superradiant decay!**

# Experimental Proof!



# Superradiance in Rydberg systems

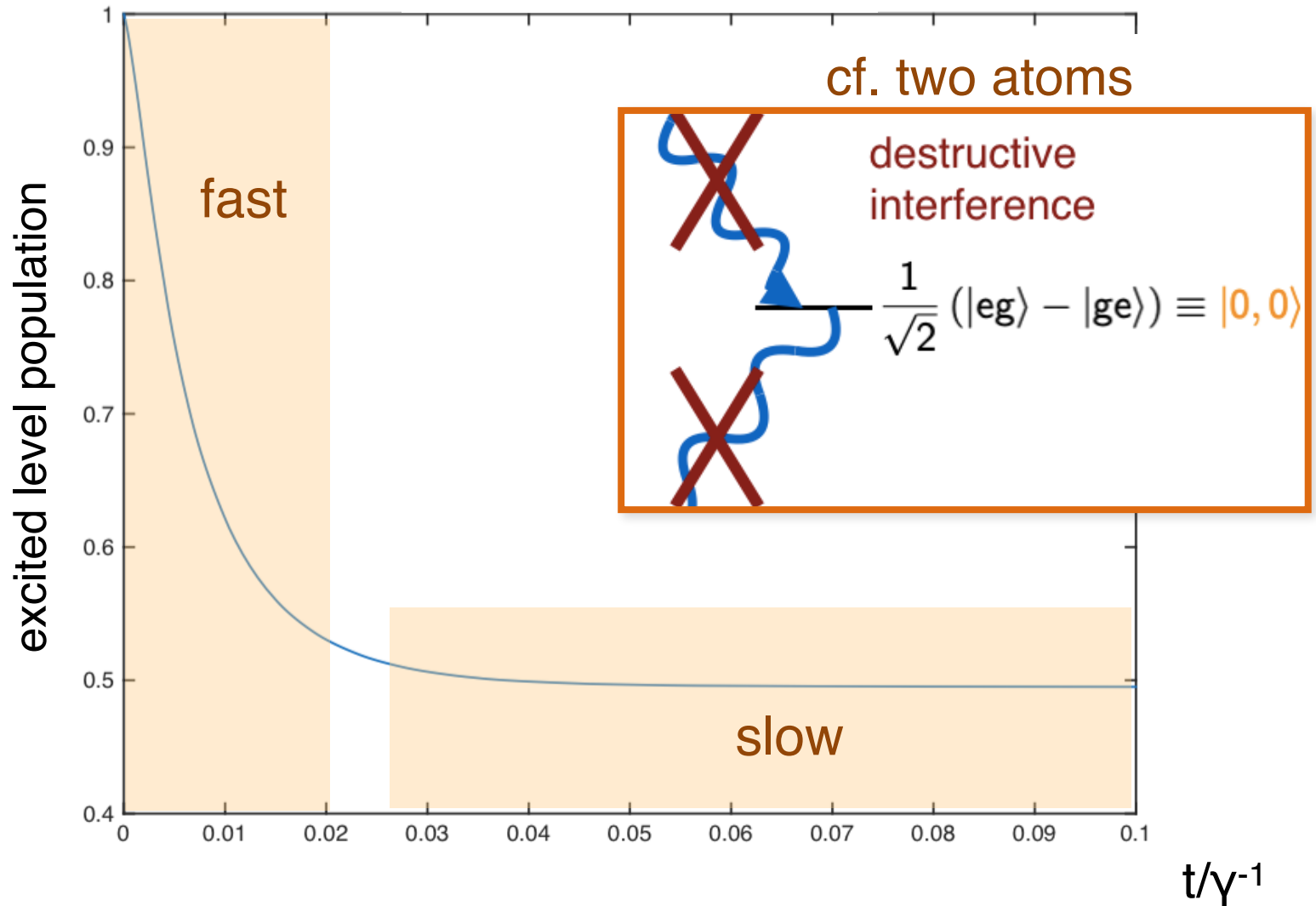


# These lectures

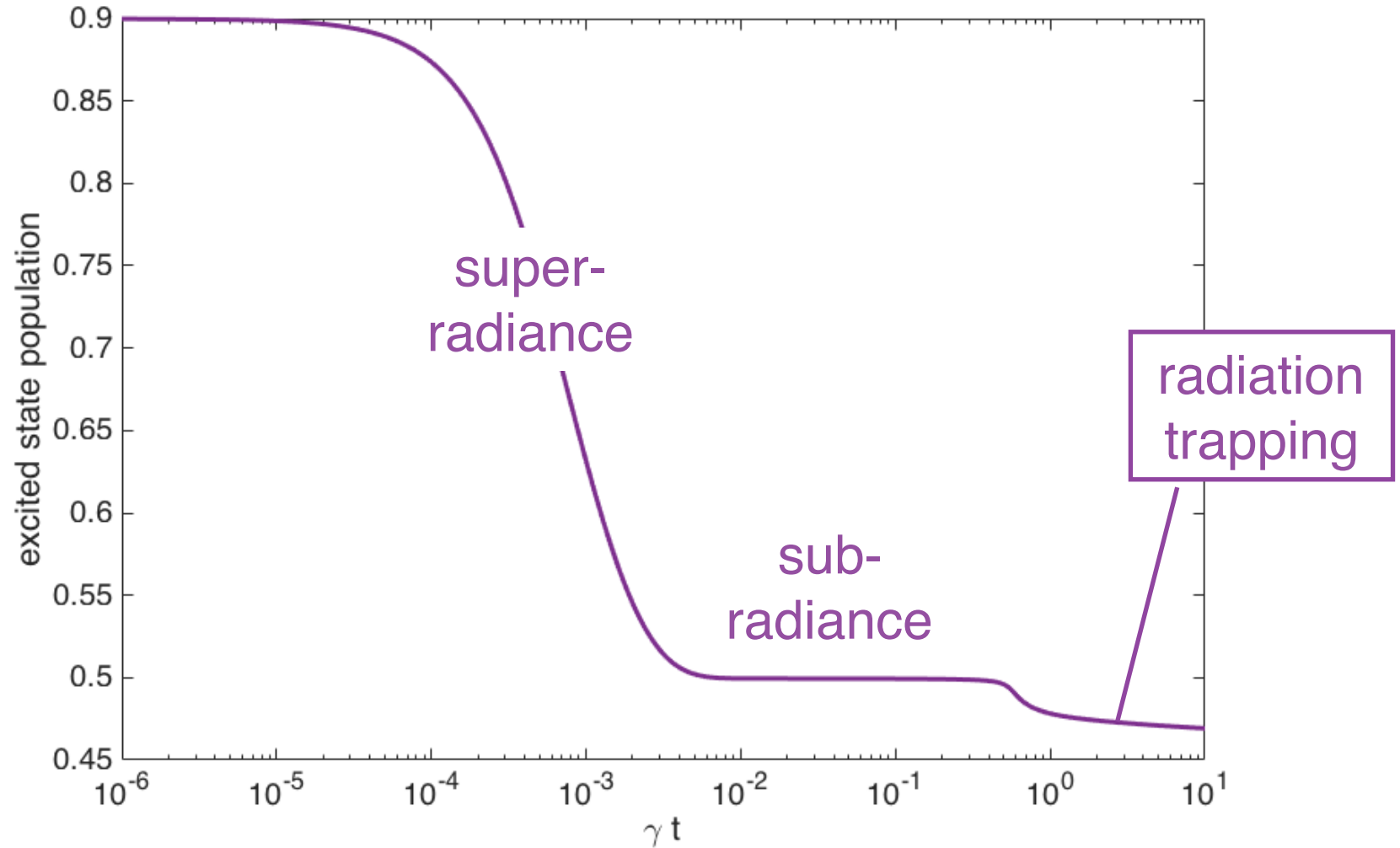
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- Cooperative effects in complex systems
  - ▶ Collective (Lamb) level shifts
  - ▶ Subradiance
  - ▶ Entanglement
- New application: atomically thin mirrors

# Decay dynamics

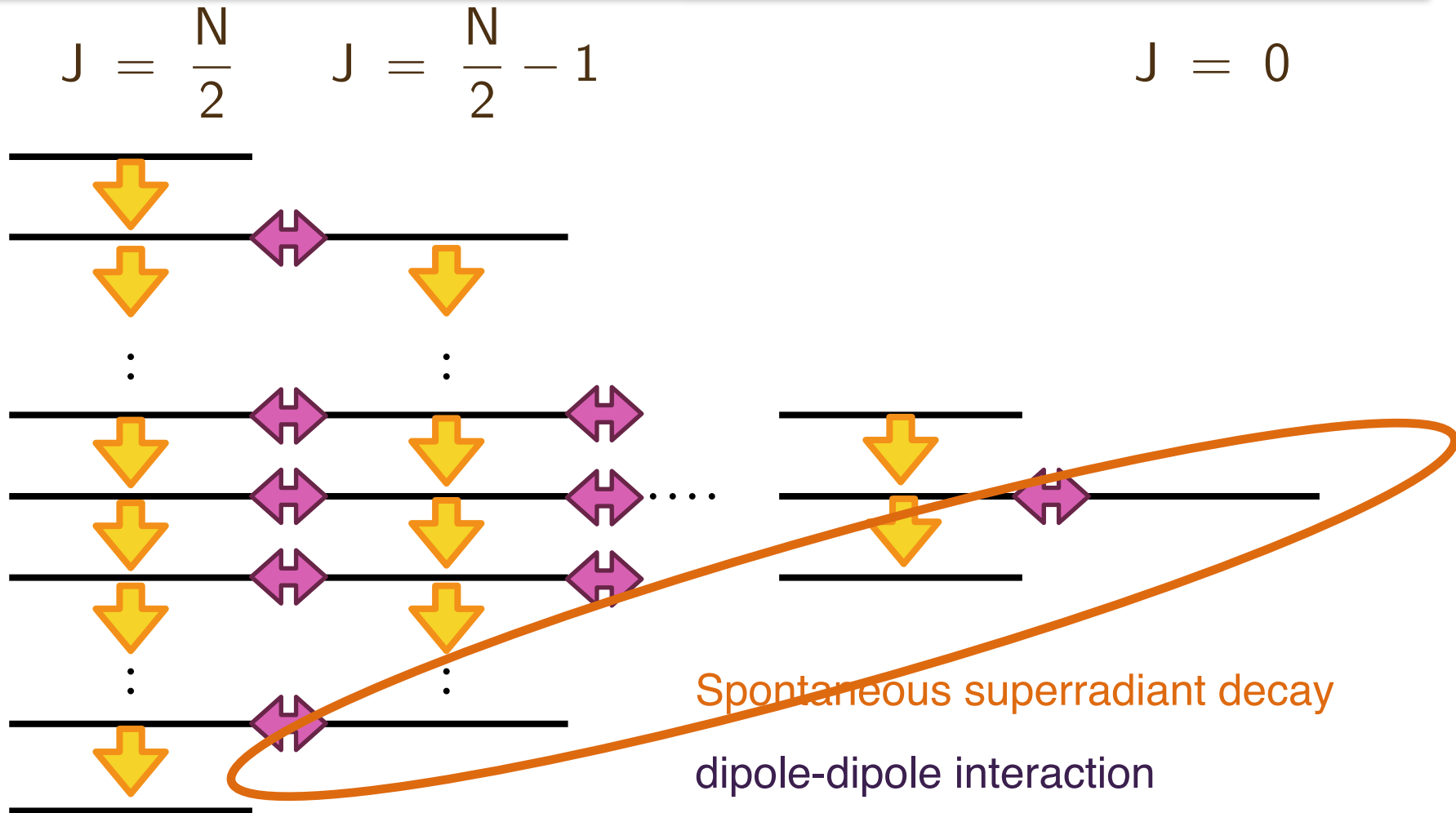


# Subradiance?

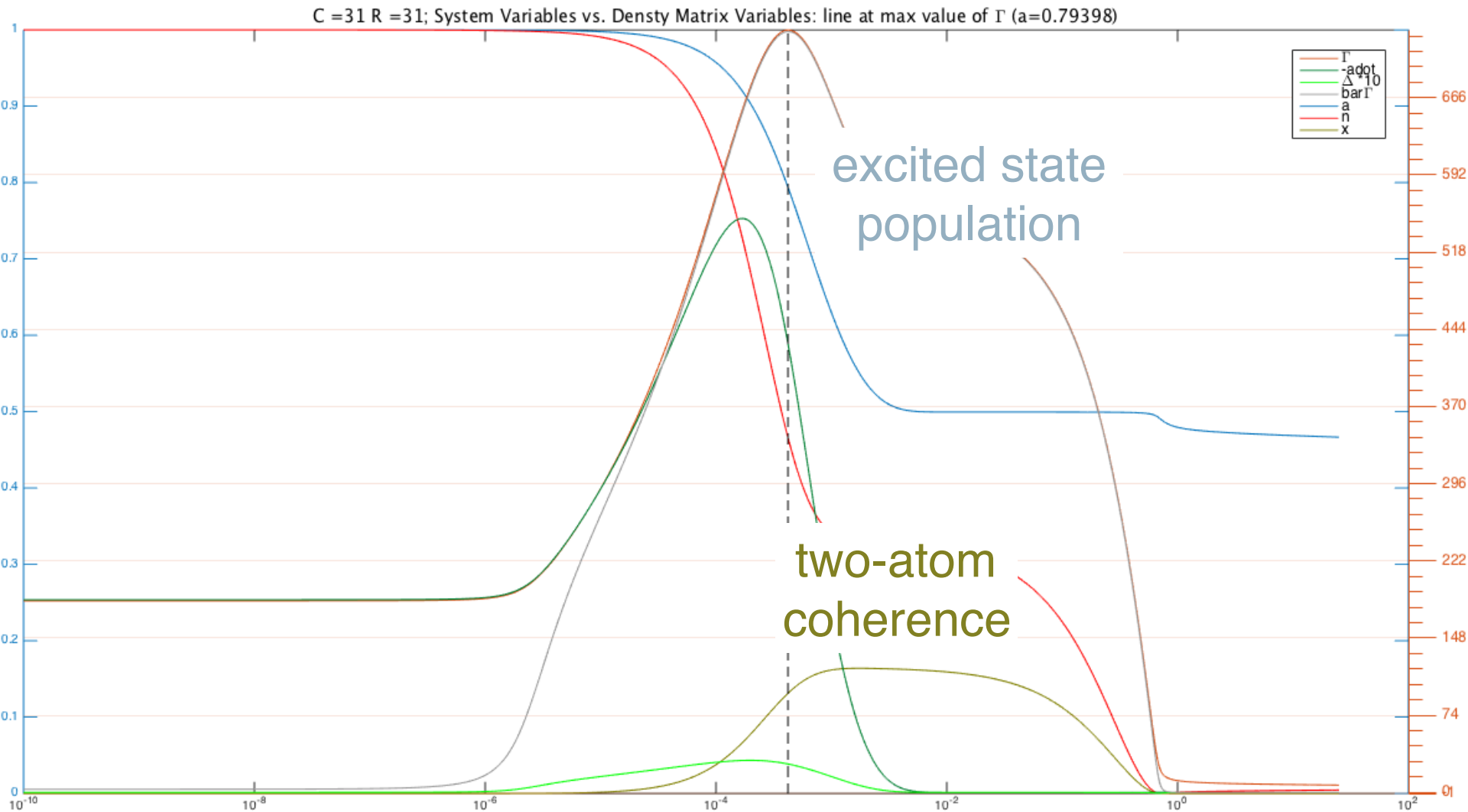




# Subradiant states



# Transitions



# Subradiance: Outlook

---

- Dynamics of subradiance = transition to many-body localized state?
  - ➔ Create, manipulate localization
- Engineered subradiance to create stable states without spontaneous decay
  - ➔ Create, stabilize many-body entangled state (Dissipative non-equilibrium physics)

# Collective Lamb shift

---

- “Lamb shift” is the result of interaction with the vacuum fluctuations
- In the case of altered density of states of the “vacuum” (i.e., the surrounding space), the value of the shift changes
- With a high (superradiant) density of radiators, the density of states inside the medium can be considerably altered



“Collective Lamb shift”

# Collective Shift

has **spontaneous** part....

$$\gamma_{ij}(\omega) = \frac{\rho^2}{\hbar^2} \int d\tau \langle [E_i^-(t), E_j^+(t + \tau)] \rangle e^{i\omega\tau}$$

$$\Delta_{\text{spont}}^{(ij)} = \frac{1}{2\pi} \mathcal{P} \int d\omega' \frac{\gamma_{ij}(\omega')}{\omega - \omega'}$$

$\Delta[\gamma]$

$$\langle [E^-, E^+] \rangle \propto \langle [a, a^\dagger] \rangle = 1$$

independent on number of photons

time  $[\gamma^{-1}]$

0.2

# Collective Shift

has **spontaneous** part....

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independent on number of photons

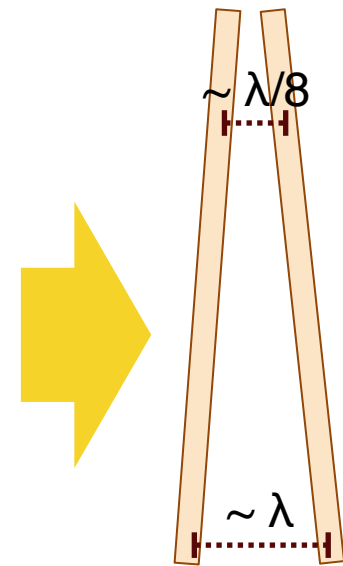
collective Lamb Shift

time [ $\gamma^{-1}$ ]

0.2

# Collective Lamb Shift

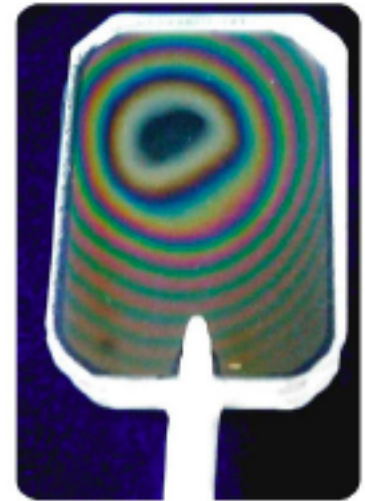
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cavity  
with variable  
thickness

# Collective Lamb Shift

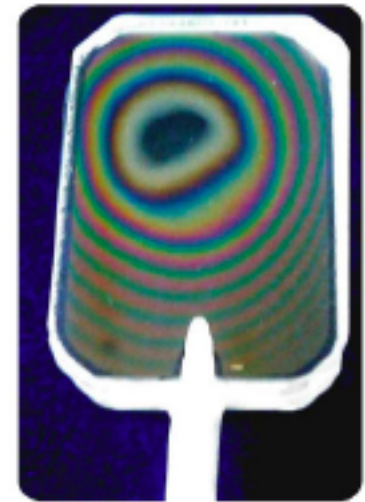
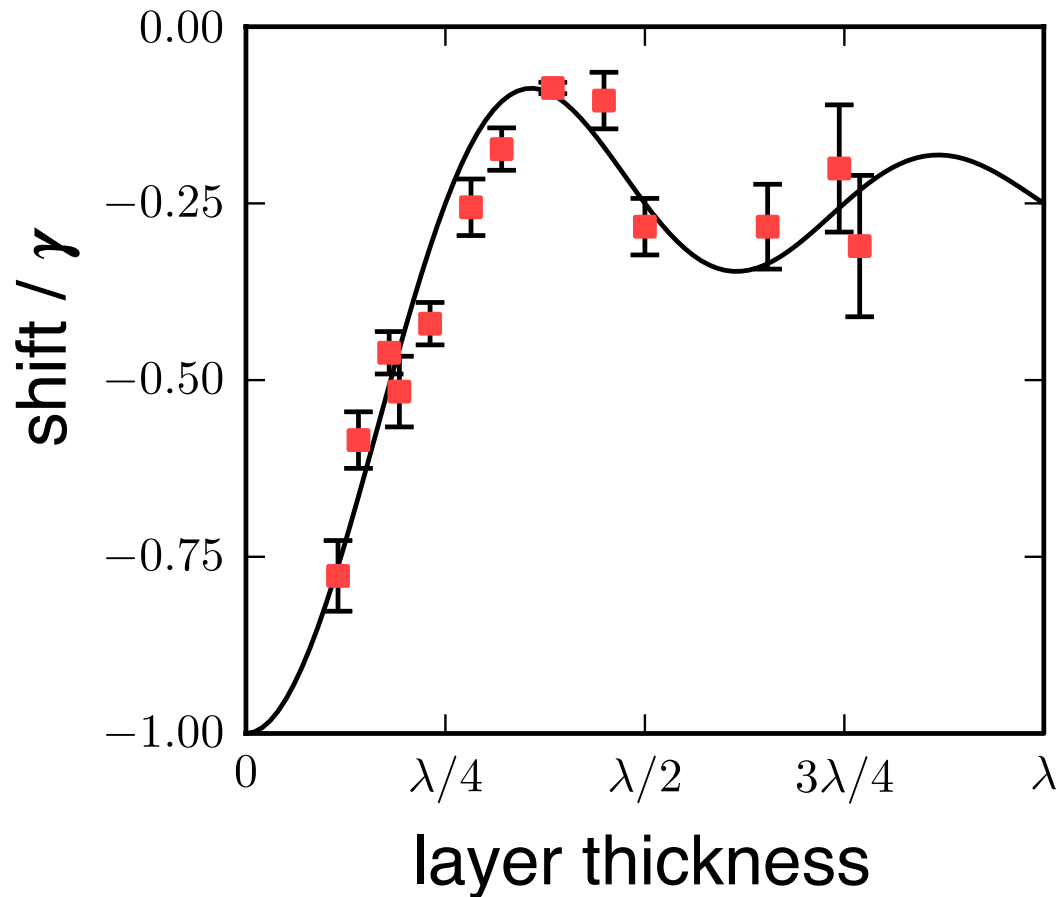
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cavity  
with variable  
thickness

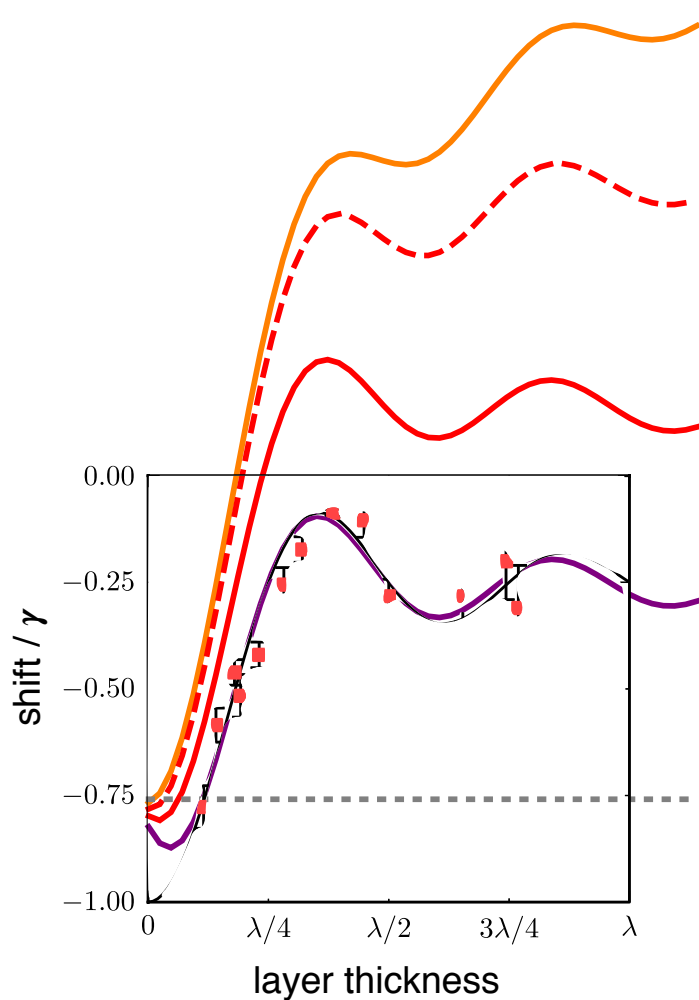


# Collective Lamb Shift



cavity  
with variable  
thickness

# Collective Lamb Shift



Optical depth = highest

Optical depth = lowest

# Collective Shift

has **spontaneous** part....

$$\frac{\wp^2}{\hbar^2} \int d\tau \langle [E_i^-(t), E_j^+(t + \tau)] \rangle e^{i\omega\tau}$$

$$\langle\langle E^- E^+ \rangle\rangle \approx \langle\langle a^\dagger a \rangle\rangle = n$$

dependent on number of photons

... and **“stimulated”** part

$$\Gamma_{ij}(\omega) = \frac{\wp^2}{\hbar^2} \int d\tau \langle\langle E_i^-(t) E_j^+(t + \tau) \rangle\rangle e^{i\omega\tau}$$

$$\Delta_{\text{stim}}^{(ij)} = \frac{1}{2\pi} \mathcal{P} \int d\omega' \frac{\Gamma_{ij}(\omega')}{\omega - \omega'}$$

$\Delta[\gamma]$

-4

-6

$\omega'$

# Collective Shift

has **spontaneous** part....

$$\frac{\wp^2}{\hbar^2} \int d\tau \langle [E_i^-(t), E_j^+(t + \tau)] \rangle e^{i\omega\tau}$$

$$\langle\langle E^- E^+ \rangle\rangle \approx \langle\langle a^\dagger a \rangle\rangle = n$$

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$$\Gamma_{ij}(\omega) = \frac{\wp^2}{\hbar^2} \int d\tau \langle\langle E_i^-(t) E_j^+(t + \tau) \rangle\rangle e^{i\omega\tau}$$

$$\Delta_{\text{stim}}^{(ij)} = \frac{1}{2\pi} \mathcal{P} \int d\omega' \frac{\Gamma_{ij}(\omega')}{\omega - \omega'}$$

Induced shift

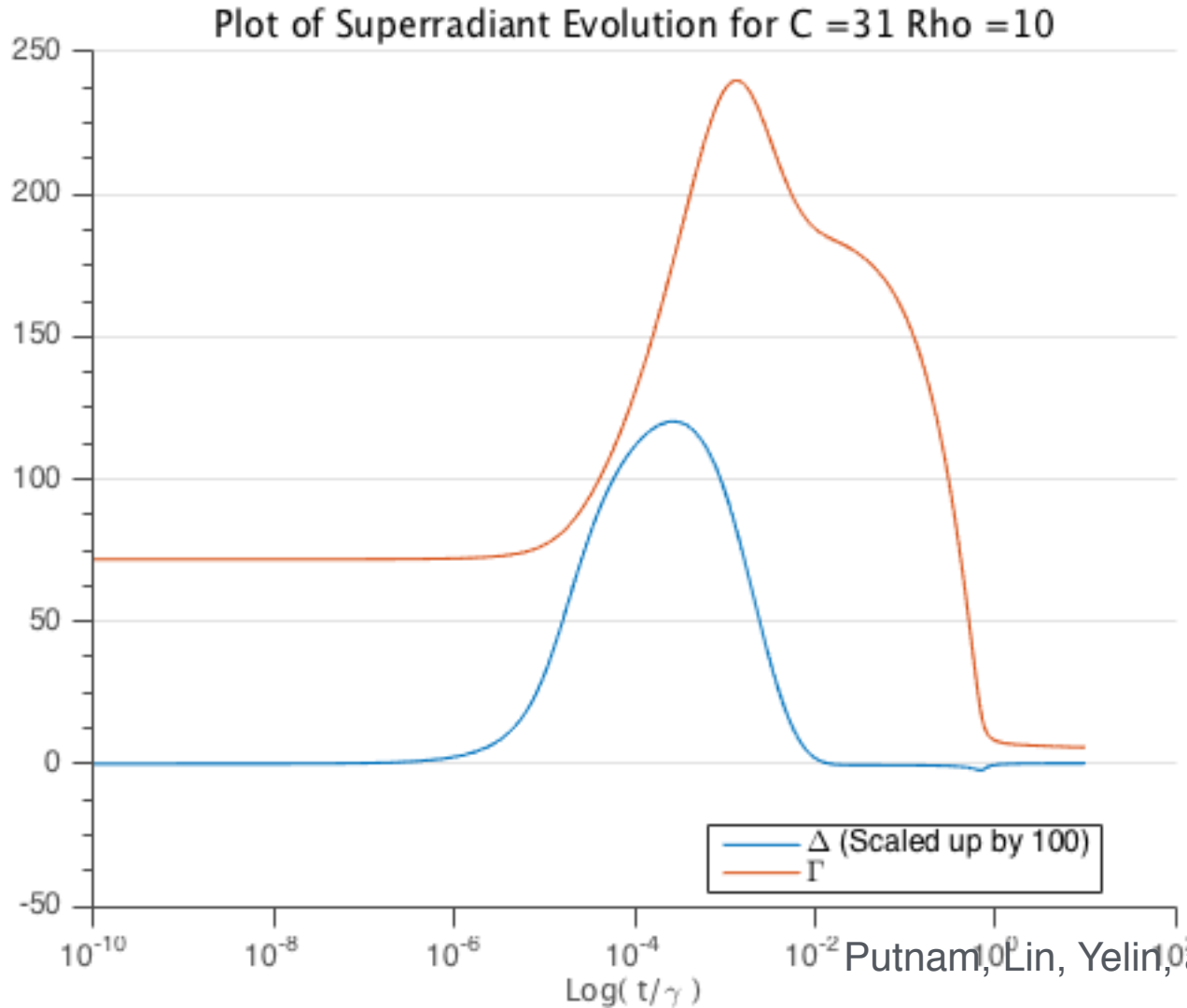
$\Delta[\gamma]$

-4

-6

$\omega'$

# Collective Shift: decay of inverted TLS



# These lectures

---

- Cooperative effects in complex systems
  - ▶ Collective (Lamb) level shifts
  - ▶ Subradiance
  - ▶ Entanglement
- New application: atomically thin mirrors

# Superradiance and Entanglement

---

Does (Dicke) superradiance need/create entanglement?

NO

# Example: 2-atom Dicke

---

- PPT (Peres-Horodecki) criterion:  
Eigenvalues of partial positive transpose  $\geq 0$

$$\rho = \sum_{ijkl} p_{kl}^{ij} |i\rangle\langle j| \otimes |k\rangle\langle l|$$

$$\rho^{\text{PPT}} = \sum_{ijkl} p_{kl}^{ij} |i\rangle\langle j| \otimes |l\rangle\langle k|$$



# Superradiance and Entanglement

---

How to define/calculate many-particle entanglement?

O Gunne, G Toth

Multipartite entanglement in four-qubit class states

YK Bai, ZD Wang

Exact and asymptotic measures of multipartite pure-state entanglement

CH Bennett, S Popescu, JA Smolin

Geometric measures of entanglement and applications to bipartite and multipartite quantum states

TC Wei, PM

Scalable multipartite entanglement of trapped ions

H Häffner, V Vesselinovik, CF Roos, J Benhelm, M Chwalla

**10,000 for "definition of multipartite entanglement"**

# Superradiance and Entanglement

---

Does (Dicke) superradiance need/create entanglement? (Initial state: no entanglement)

Dicke superradiant  
time evolution

=

separable states

**constructive proof**

# Superradiance and Entanglement

---

Does (Dicke) superradiance need/create entanglement? (Initial state: no entanglement)

Dicke superradiant  
states

=

separable states

# Superradiance and Entanglement

---

Does (Dicke) superradiance need/create entanglement? (Initial state: no entanglement)

our system: mixed state of  
N-atom Dicke states with  
N+1 known independent  
coefficients  $p_i$

=

compare to mixture of  
symmetric product states of  
N (two-level) atoms (needs  
N+1 coefficients  $y_i$ )

(N+1) - dim.  
equation  
system

# Form of equations

---

- General Dicke states:

$$\rho_{\text{GDS}} = \sum_{\mathbf{n}} \chi_{\mathbf{n}} |D_{\mathbf{n}}\rangle \langle D_{\mathbf{n}}|$$

- Separable diagonally symmetric:

$$\rho_{\text{SDS}} = N! \sum_{\mathbf{n}} \sum_{j=1}^{j_{\max}} \frac{x_j y_j^{n_0} (1 - y_j)^{n_1}}{n_0! n_1!} |D_{\mathbf{n}}\rangle \langle D_{\mathbf{n}}|$$

# Superradiance and Entanglement

Does (Dicke) superradiance need/create entanglement? (Initial state: no entanglement)

our system: mixture of  
N-atom Dicke states  
N+1 known independent  
coefficients

compare to mixture of  
symmetric product states of  
(two-level) atoms (needs  
N+1 coefficients  $y_i$ )

condition:  
all coefficients  
 $0 \leq p_i \leq 1$

equation  
system

# Superradiance and Entanglement

Does (Dicke) superradiance need/create entanglement? (Initial state: no entanglement)

our system: mixed state of  
N-atom Dicke states with  
N+1 known independent  
coefficients  $p_i$

=

complete mixture of  
symmetric states of  
N (two terms (needs  
N terms  $y_i$ )



condition:  
all coefficients  
 $0 \leq p_i \leq 1$

(N+1) - dim.  
equation  
system

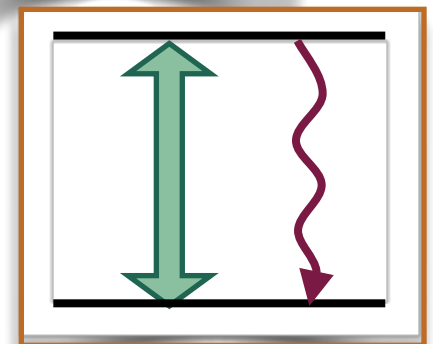
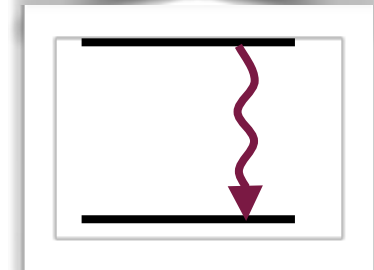
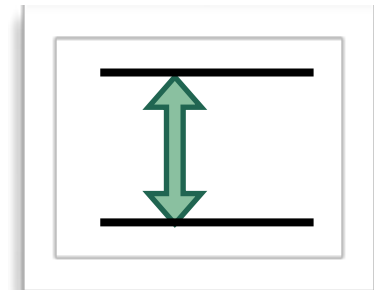


# Superradiance and Entanglement

---

Driven superradiant system:

- Driving alone does not entangle atoms
- Superradiance alone does not entangle atoms
- Driving and superradiance together entangle atoms!





# Fuzzy Bunny?

---



# Spin Squeezing

---

- Correlated (“squeezed”) spins could improve resolution in one direction (“quadrature”).

# (Spin) Squeezing

---

- How to measure squeezing/measurement improvement?

$$\xi^2 \equiv \frac{\text{optimal variance}}{\text{unsqueezed optimal variance}}$$

# Spin squeezing

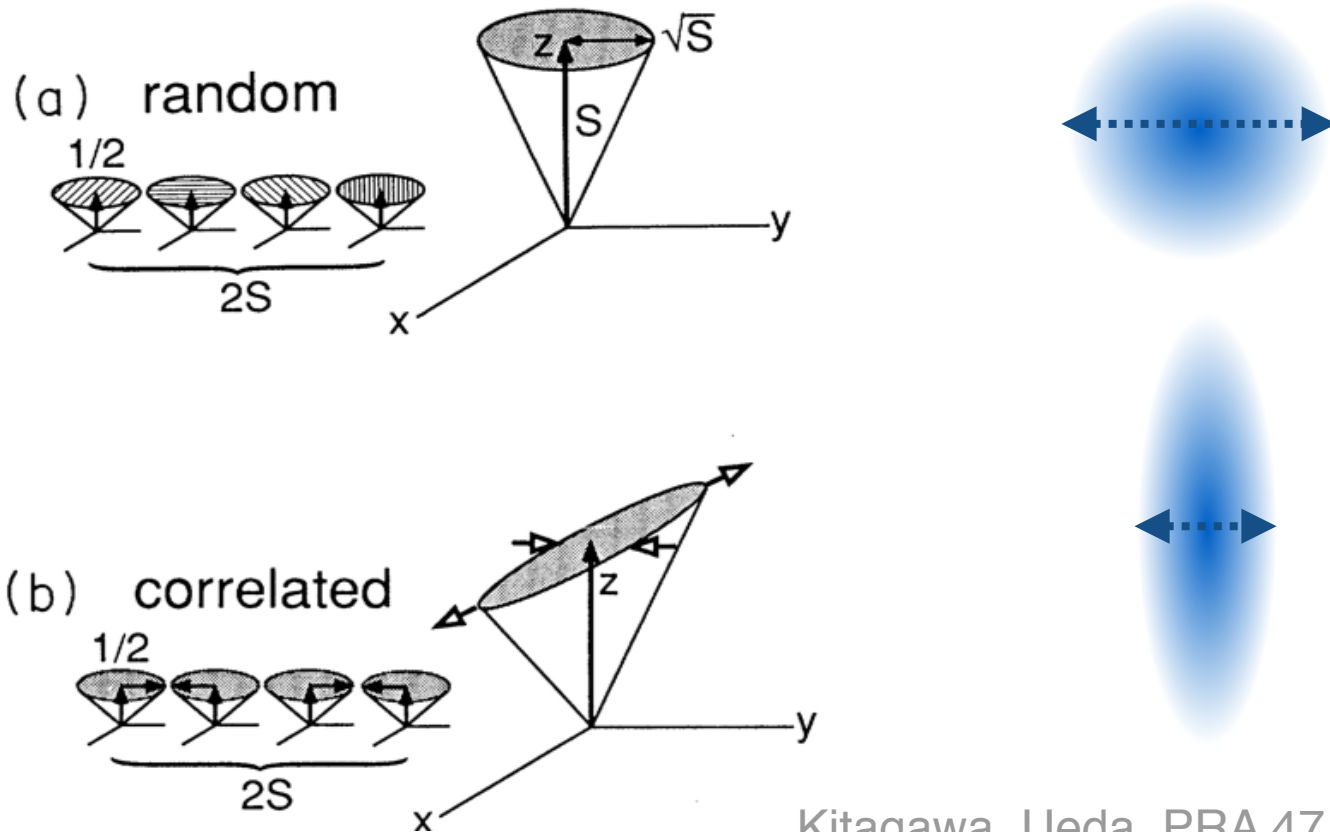
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Old problem: How to improve metrology by spin squeezing ensembles

- ➔ Groups of Bigelow, Kuzmich, Lewenstein, Mølmer, Polzik, Sanders, Sørensen, Vuletic, Wineland,...

# Spin Squeezing

- Correlated (“squeezed”) spins could improve resolution in one direction (“quadrature”).



# (Spin) Squeezing

---

- How to measure squeezing/measurement improvement?

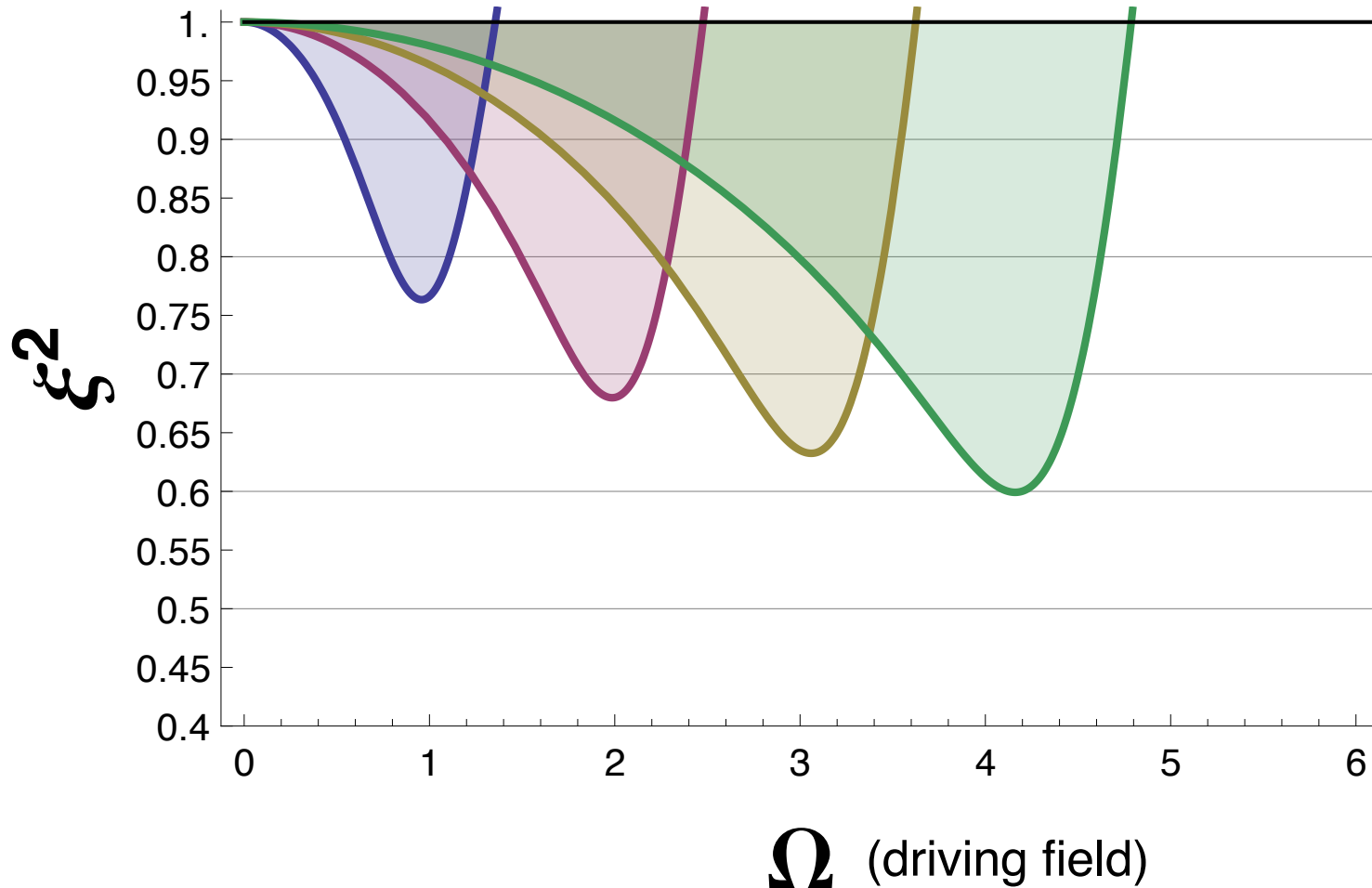
$$\xi^2 \equiv \frac{\text{optimal variance}}{\text{unsqueezed optimal variance}}$$

$$\xi^2 = \frac{N}{2} \left[ \langle J_1^2 + J_2^2 \rangle - \sqrt{\langle J_1^2 - J_2^2 \rangle^2 + \langle J_1 J_2 + J_2 J_1 \rangle^2} \right]$$

( $J_1, J_2$ , are uncertainties in the two directions orthogonal to the total spin  $\mathbf{J}$ )

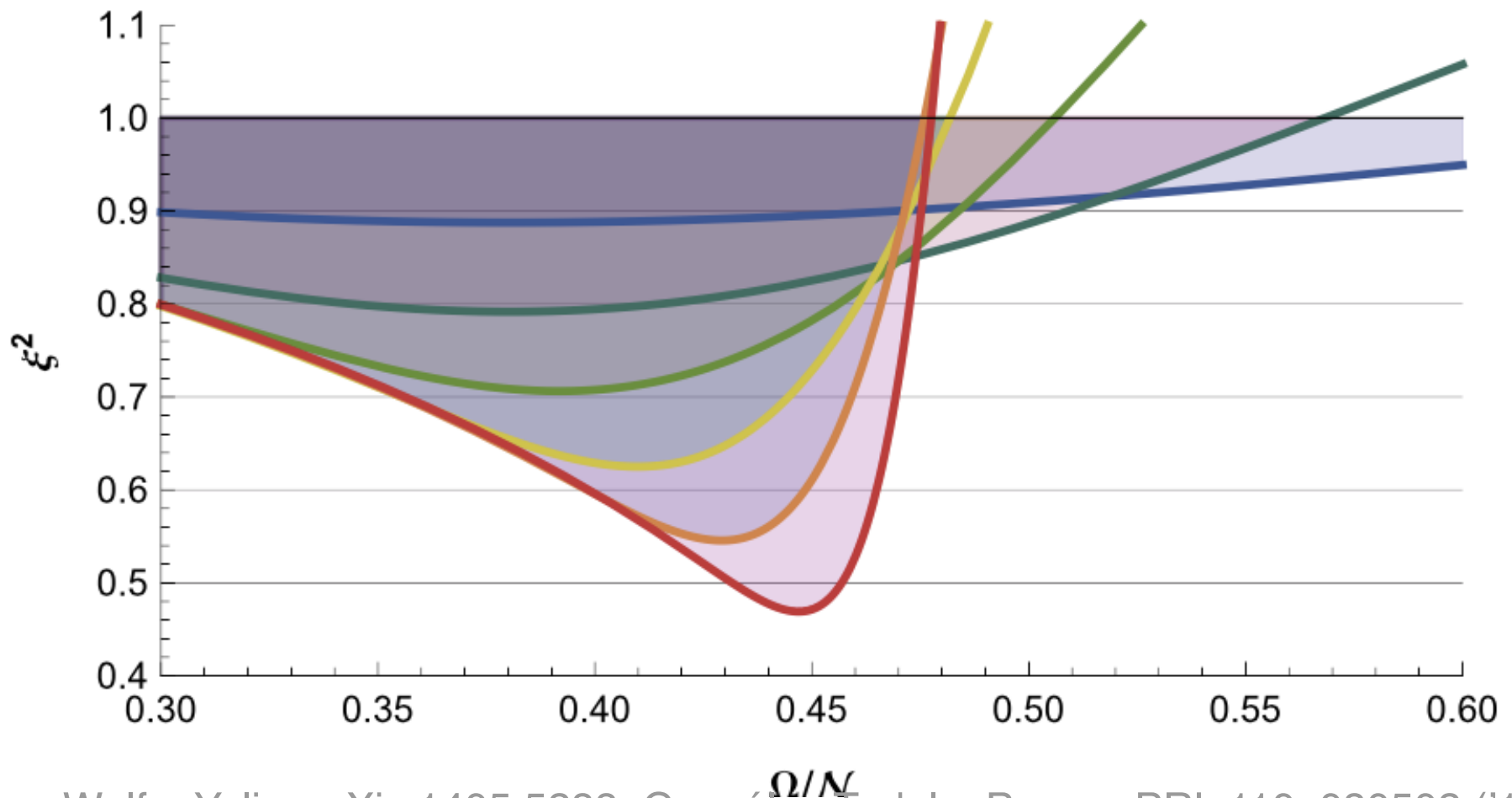
# Superradiant Spin Squeezing

$\mathcal{N} =$  ■ 5 ■ 10 ■ 15 ■ 20



# Superradiant Spin Squeezing

$\xi^2$  vs.  $\Omega/N$ ,  $N =$  ■ 2 ■ 4 ■ 8 ■ 16 ■ 32 ■ 64

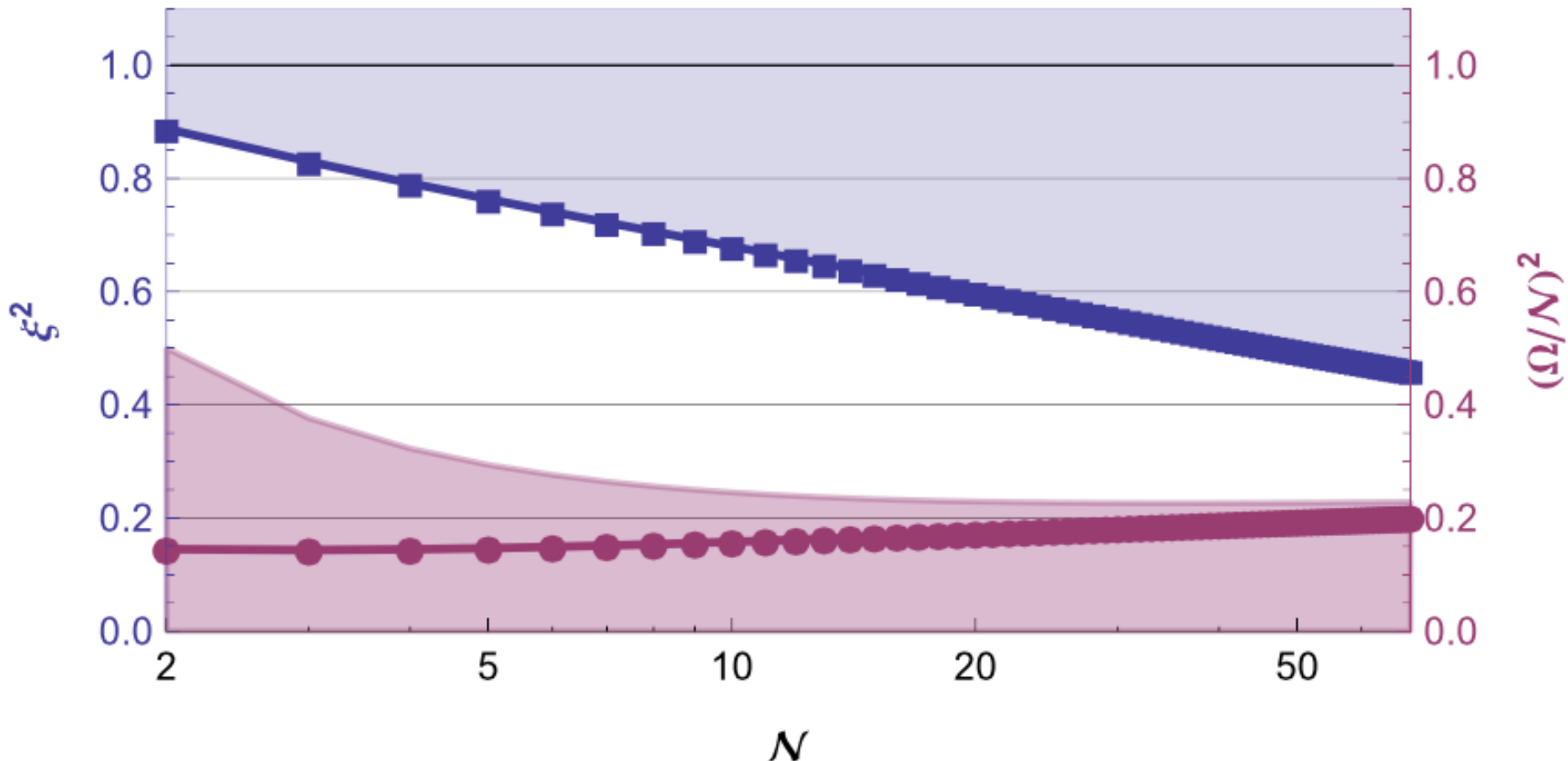




# Best case for Dicke ensemble

■ minimal possible  $\xi^2$

● most optimal  $(\Omega/\mathcal{N})^2$



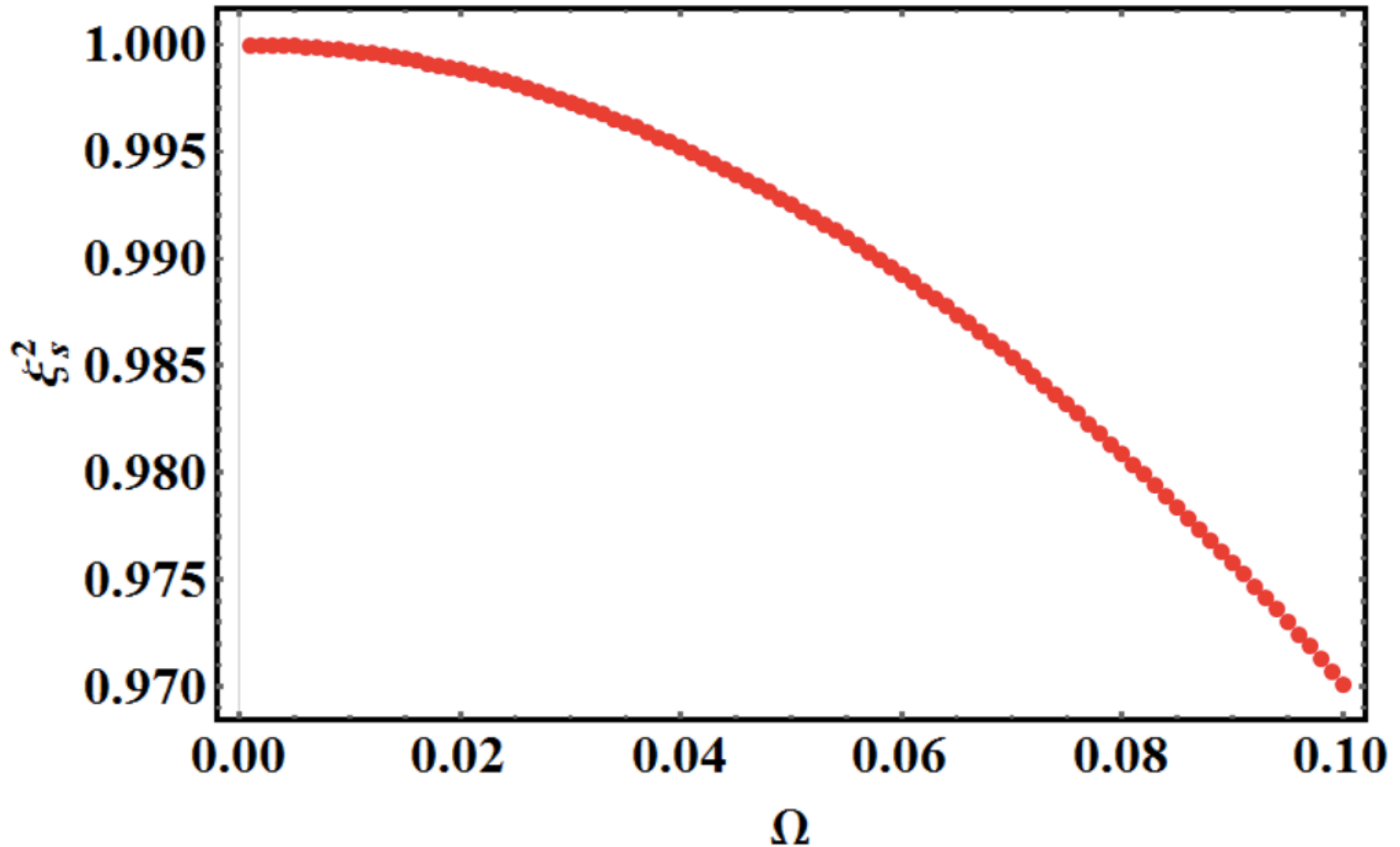
# What about realistic systems?

---

- Dicke: 3 parameters ( $N, \Gamma, \Omega$ )
  - Realistic systems: (OD, rel. density,  $\Gamma, \Omega, \gamma_{ij}, \Delta, \delta_{ij}$ )
  - Is it possible to find parallels?
  - minimize “incoherent” aspects?
- ➔ key: spontaneous decay, shift instead of induced!

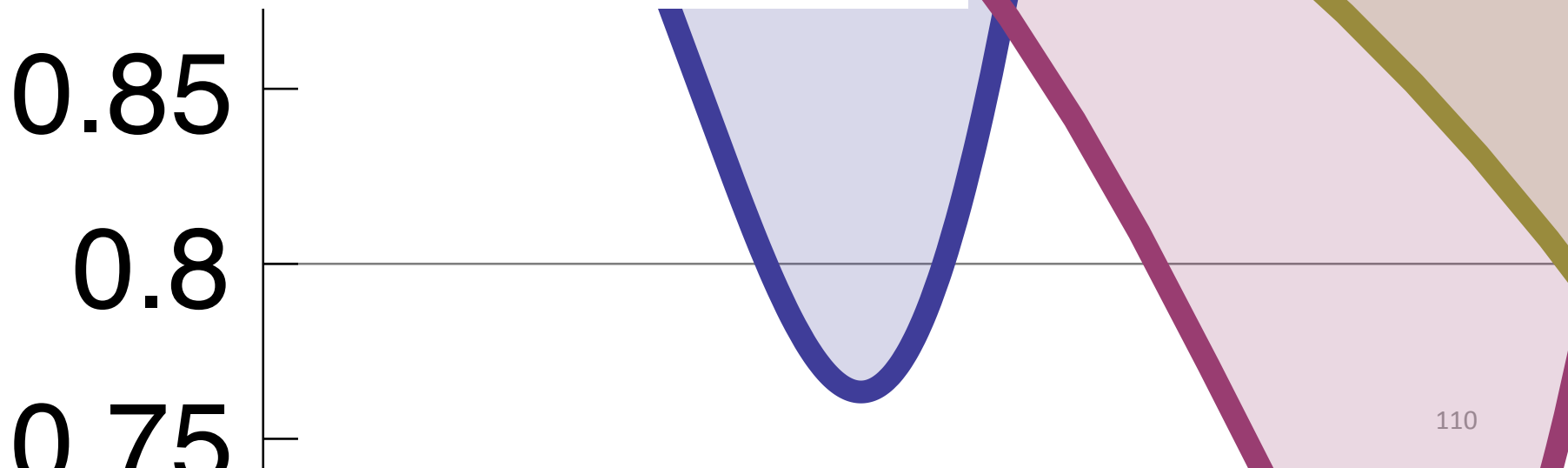
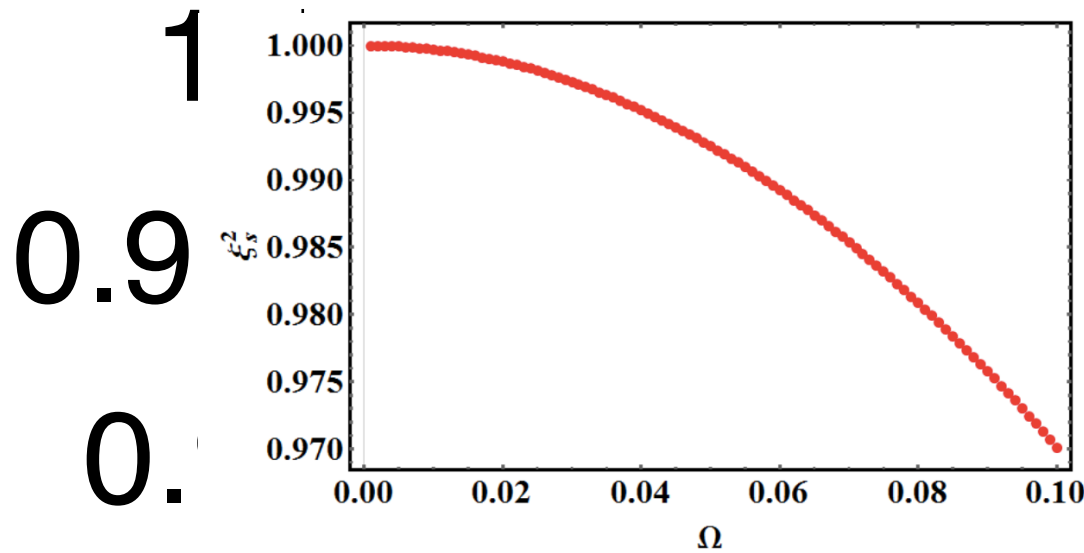
# Spin squeezing in realistic systems?

---



$N =$ 

# Spin squeezing in realistic systems?



Full dynamics (all degrees of freedom of atoms, fields)

---

$$H = H_{\text{atoms}} + H_{\text{field}} - \sum \mathbf{p}_i \mathbf{E}_i$$



two probe atoms  
scattered atoms

**Thank**

**you!**

